

# ***The Status of the keV-scale sterile neutrino search with KATRIN***

**Anton Huber for the KATRIN Collaboration**

Chalonge-de Vega Meudon Workshop 2016 - Paris



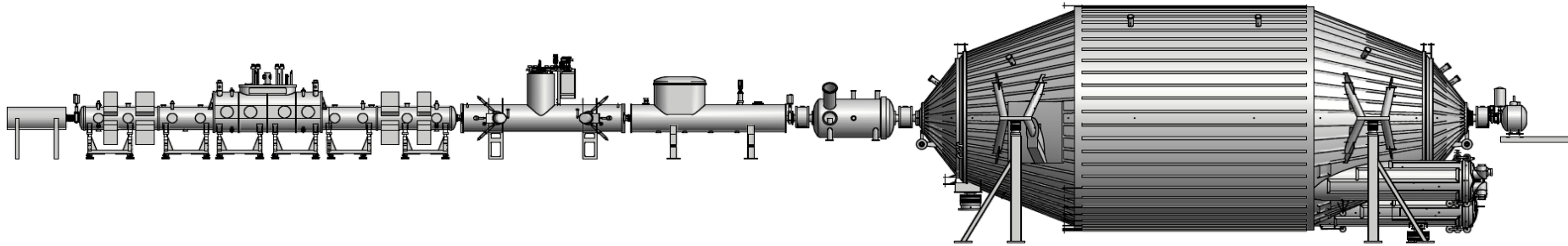
# Overview

- The KATRIN Experiment and the neutrino mass
  - Main goal
  - Measurement principle
  - Status
  
- The KATRIN Experiment and sterile neutrinos
  - Imprint of a sterile neutrino
  - How to use KATRIN
  - Two main planned measurements
  
- Conclusion and time to discuss





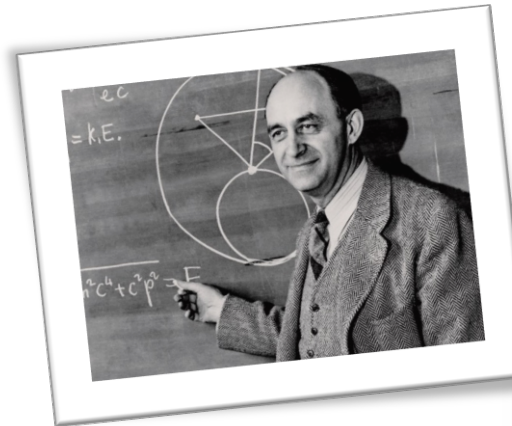
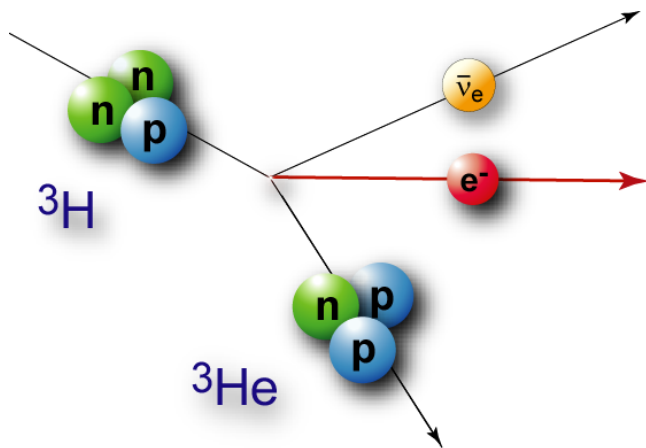
# KATRIN Experiment



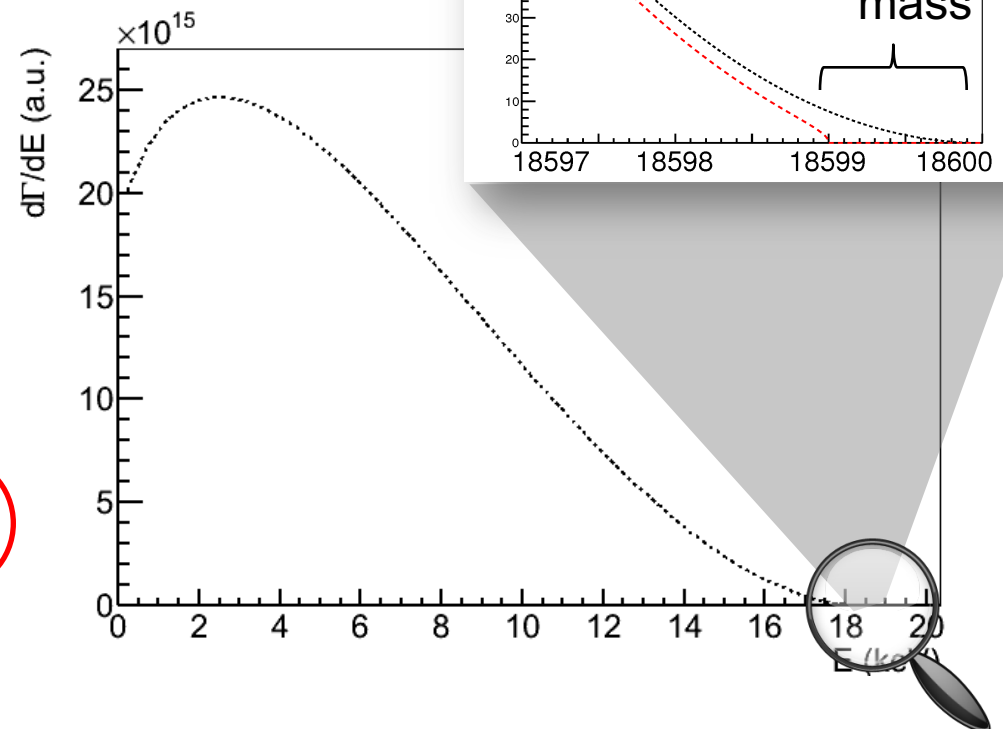
- Probing the neutrino mass with a sensitivity of  $m(\nu_e) = 200 \text{ meV}$  (90% CL), after 3 years
- Improvement of a factor of 10

KATRIN Collaboration, FZKA Scientific Report  
7090 (2004)

# Tritium beta decay



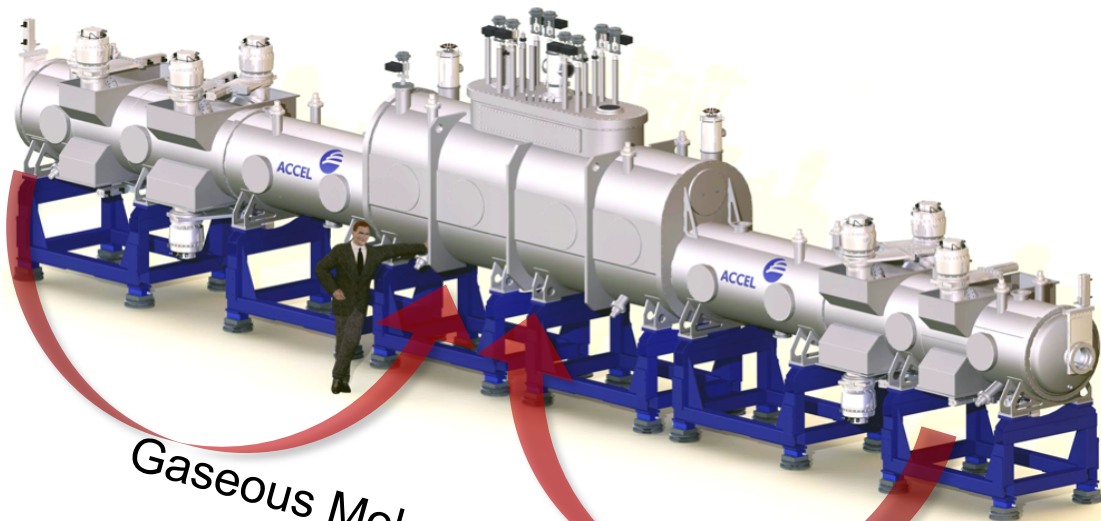
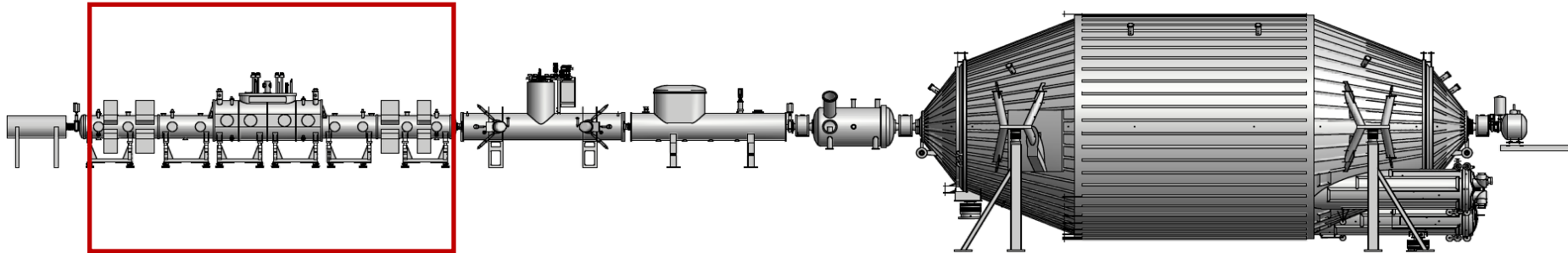
$10^{-13}$  of all decays in last eV



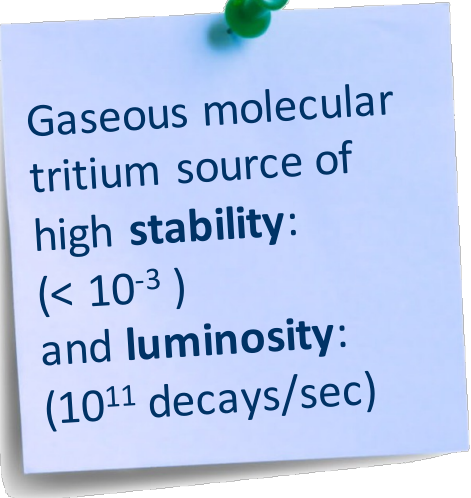
$$\frac{d\Gamma}{dE} = C \cdot F(E, Z) \cdot p \cdot (E + m_e) \cdot (E - E_0) \cdot \sqrt{(E - E_0)^2 - m_\nu^2}$$

# KATRIN Overview

(windowless) Source Section

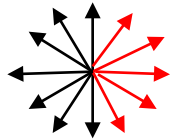


Gaseous Molecular Tritium

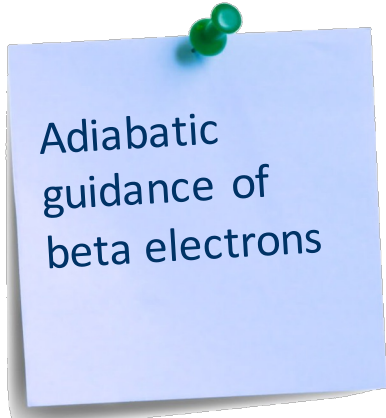
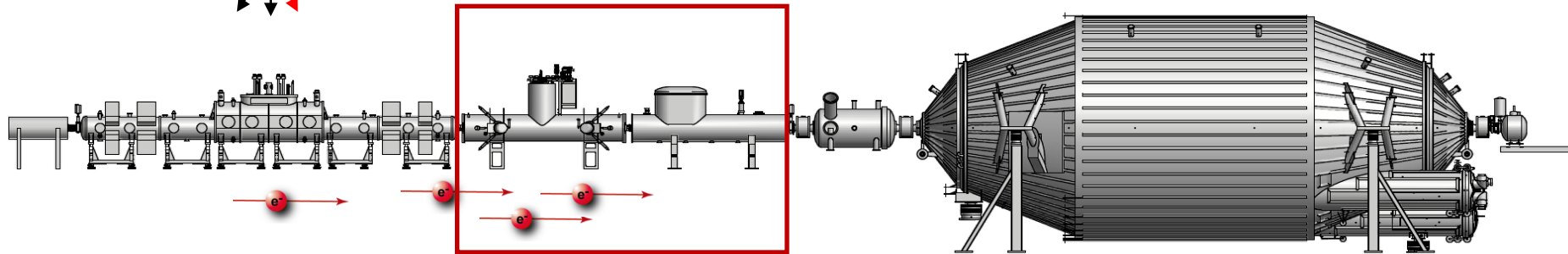


Gaseous molecular tritium source of high **stability**: ( $< 10^{-3}$ ) and **luminosity**: ( $10^{11}$  decays/sec)

# The KATRIN experiment



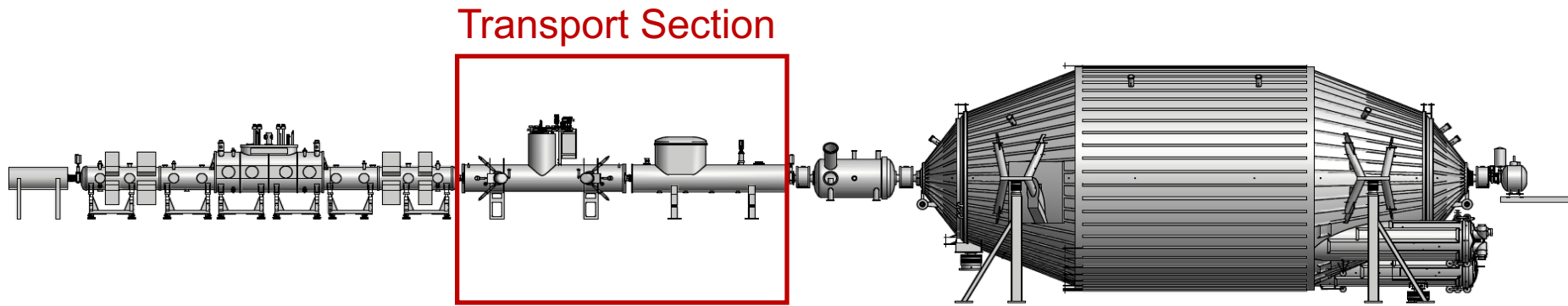
Transport Section



Adiabatic  
guidance of  
beta electrons



# The KATRIN experiment



$R > 10^7$

$R > 10^7$



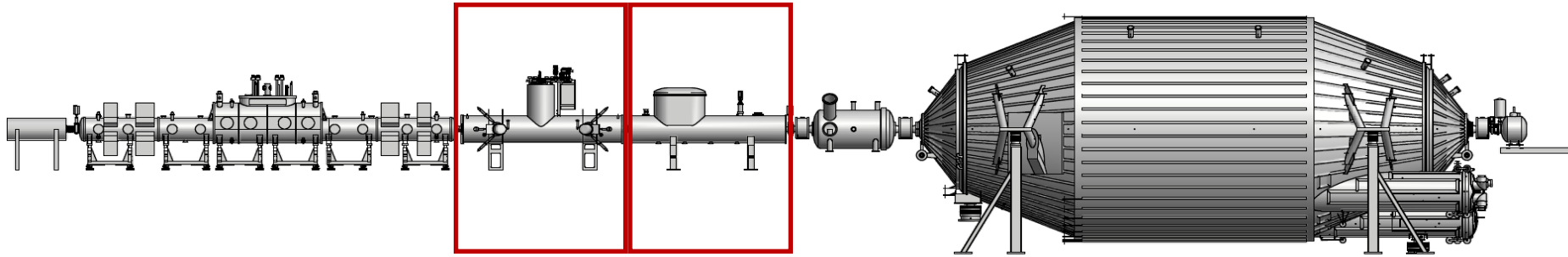
Reduction of tritium flow by 14 orders of magnitude



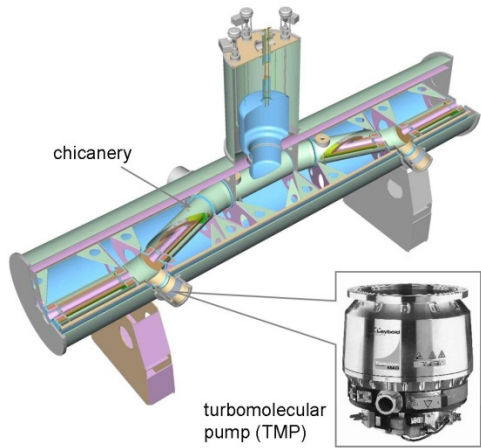
$p(T_2) < 10^{-20}$  mbar  
 $p = 10^{-11}$  mbar

# KATRIN Overview

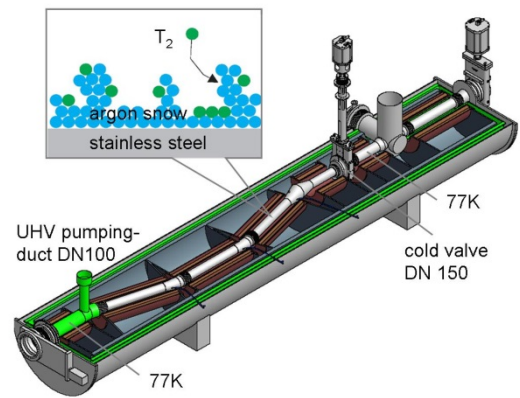
## Transport Section



## Differential pumping section

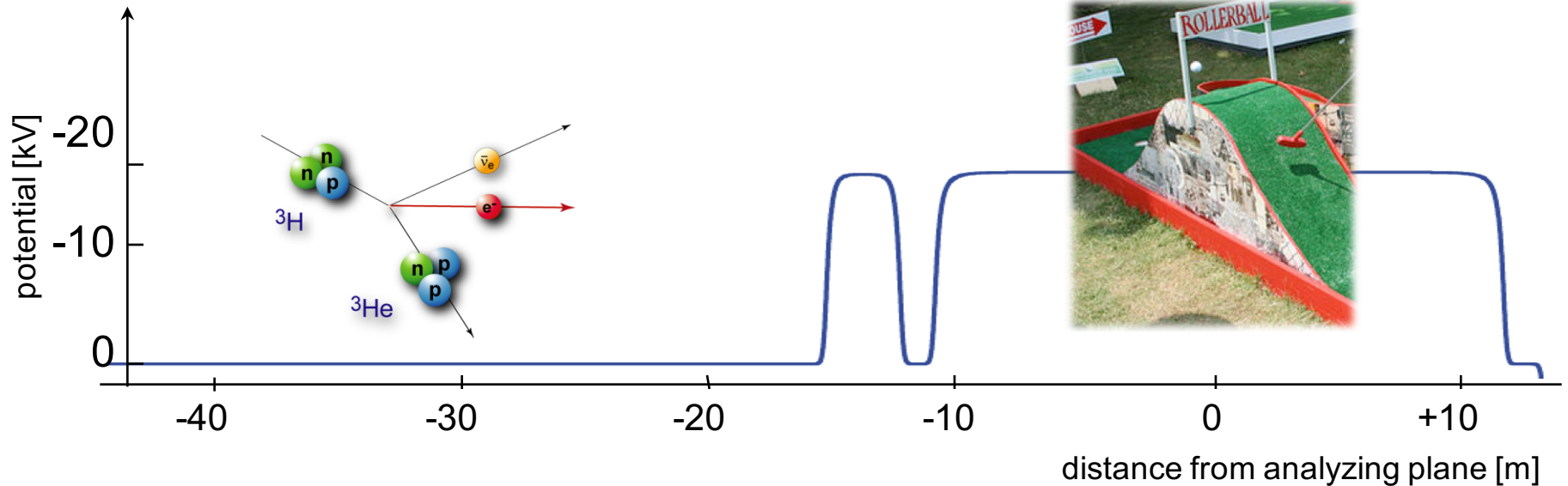
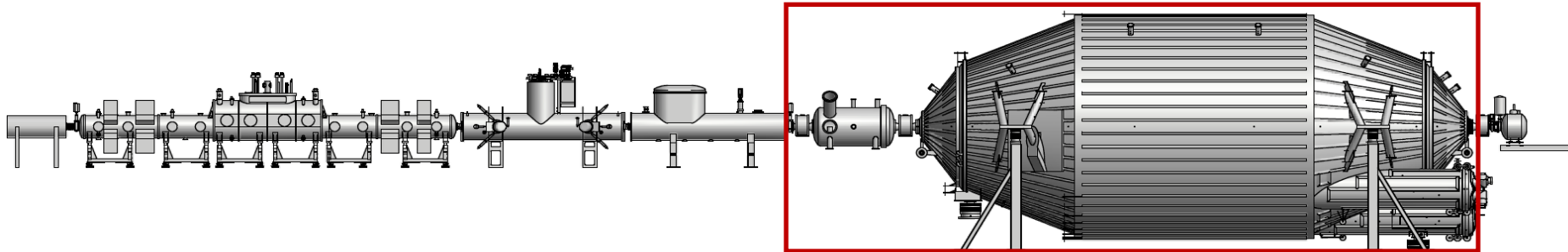


## Cryogenic pumping section

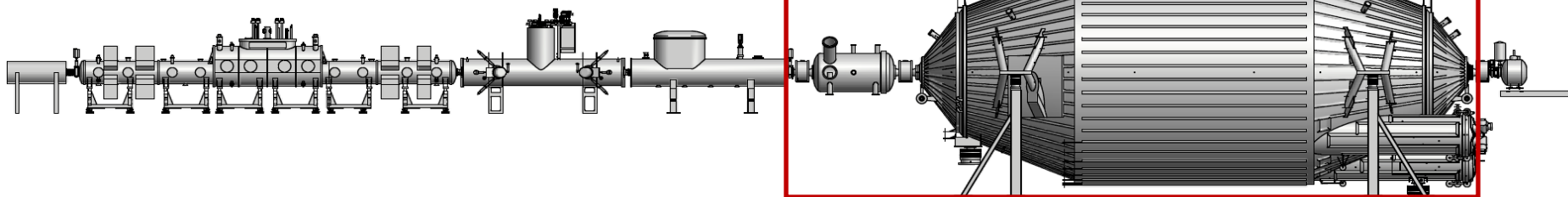


# KATRIN Overview

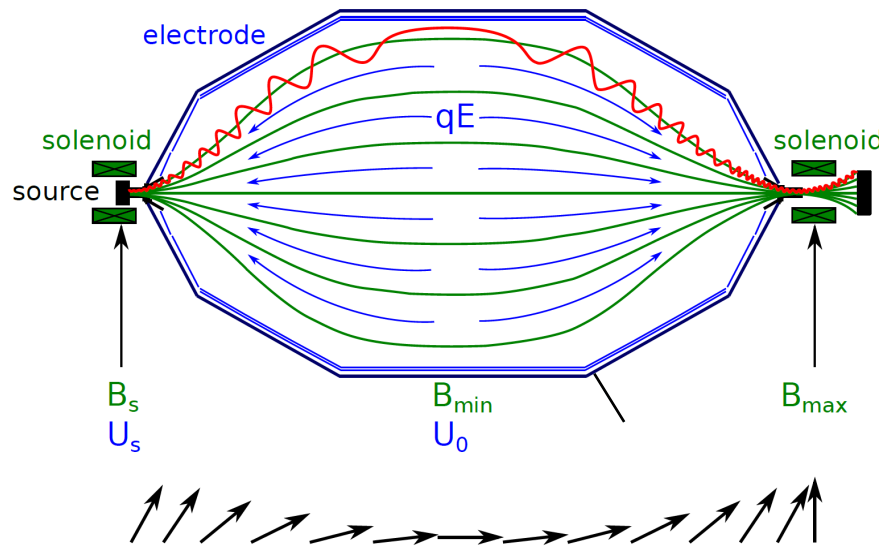
## Spectrometer Section



# KATRIN Overview



## Spectrometer Section



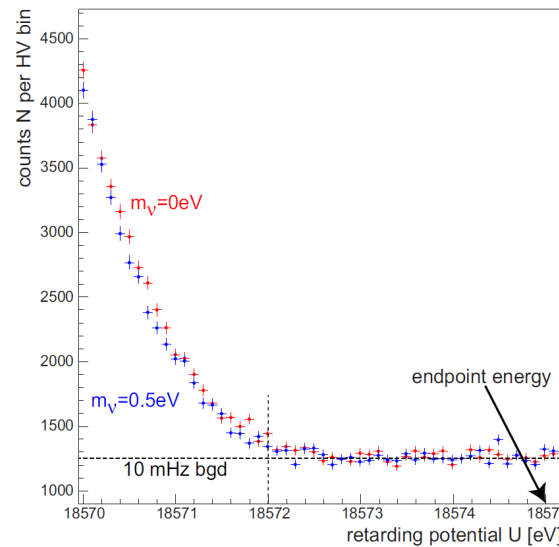
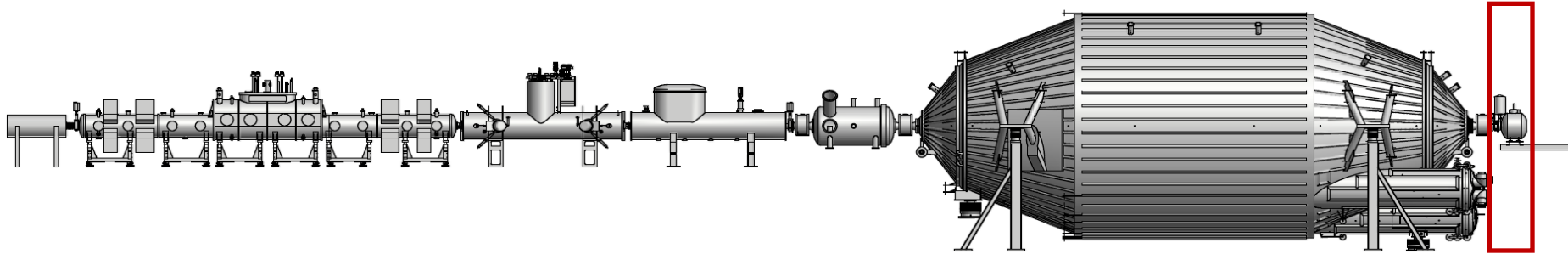
### MAC-E Principle:

Large angle acceptance

High energy resolution  
(0.93 eV)

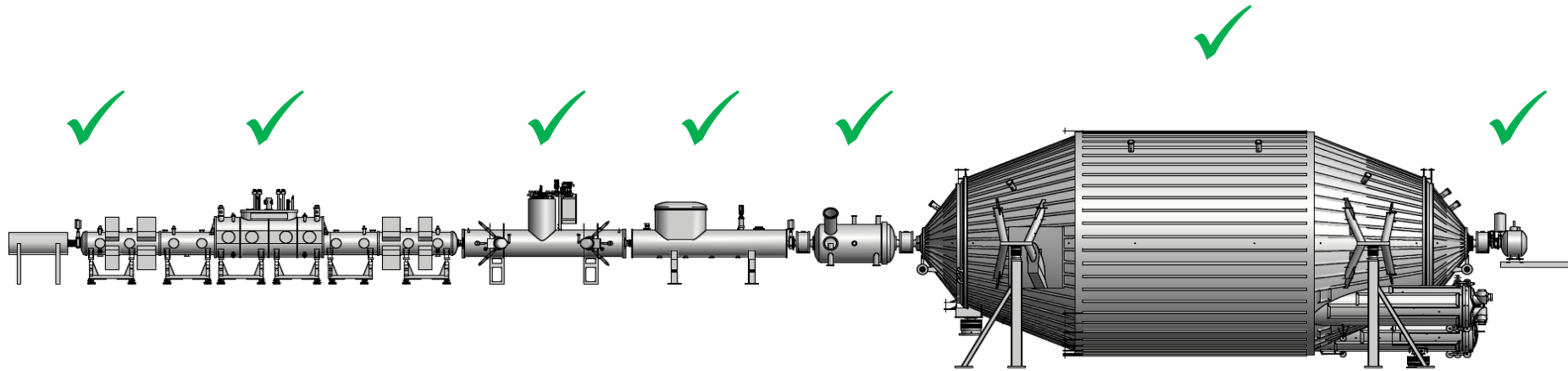
# KATRIN Overview

## Detector Section



**Integral**  
measurement of  
spectrum down to  
**30 eV** below the  
endpoint

# KATRIN Overview

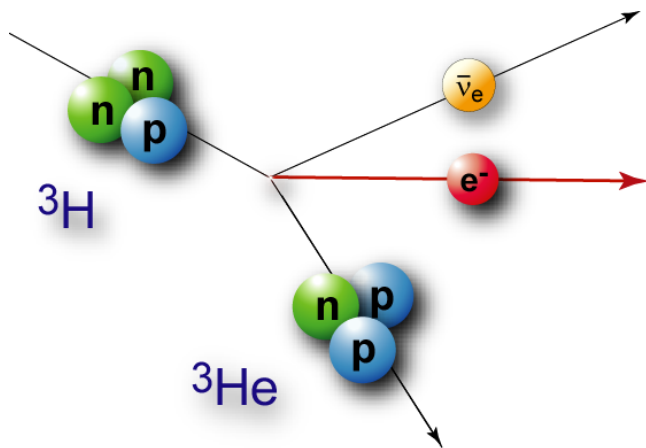


- All Components are on the site
- Alignment and connection are prepared at the moment
- First Light measurement with e-gun in October this year

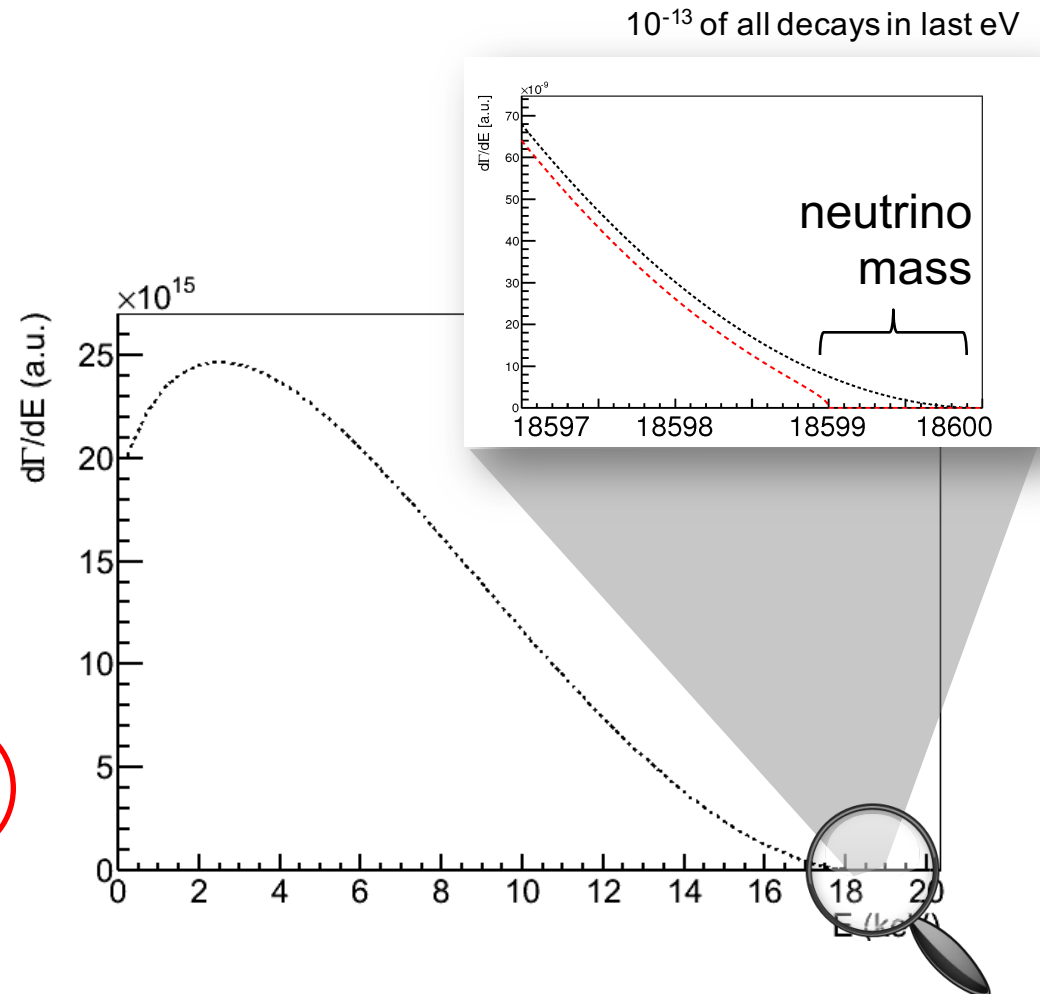
# KATRIN and the keV-scale sterile neutrino



# Tritium beta decay

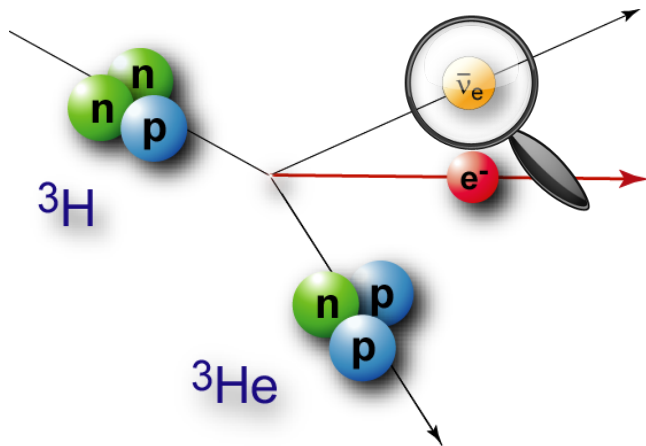


$$\frac{d\Gamma}{dE} = C \cdot F(E, Z) \cdot p \cdot (E + m_e) \cdot (E - E_0) \cdot \sqrt{(E - E_0)^2 - m_\nu^2}$$

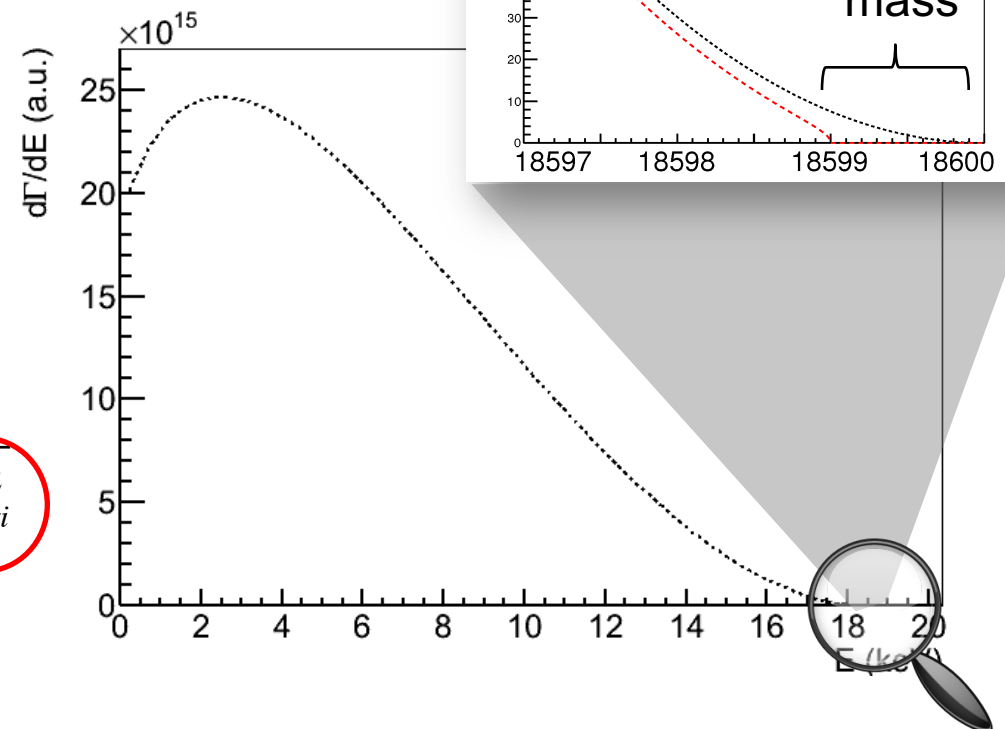




# Tritium beta decay

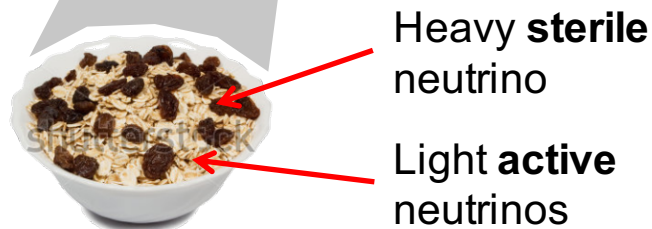
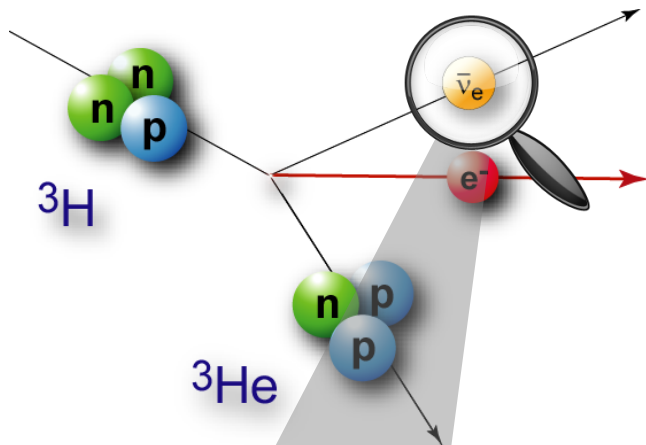


$$m_{\beta}^2 = \sum_i |U_{ei}|^2 m_{\nu_i}^2$$

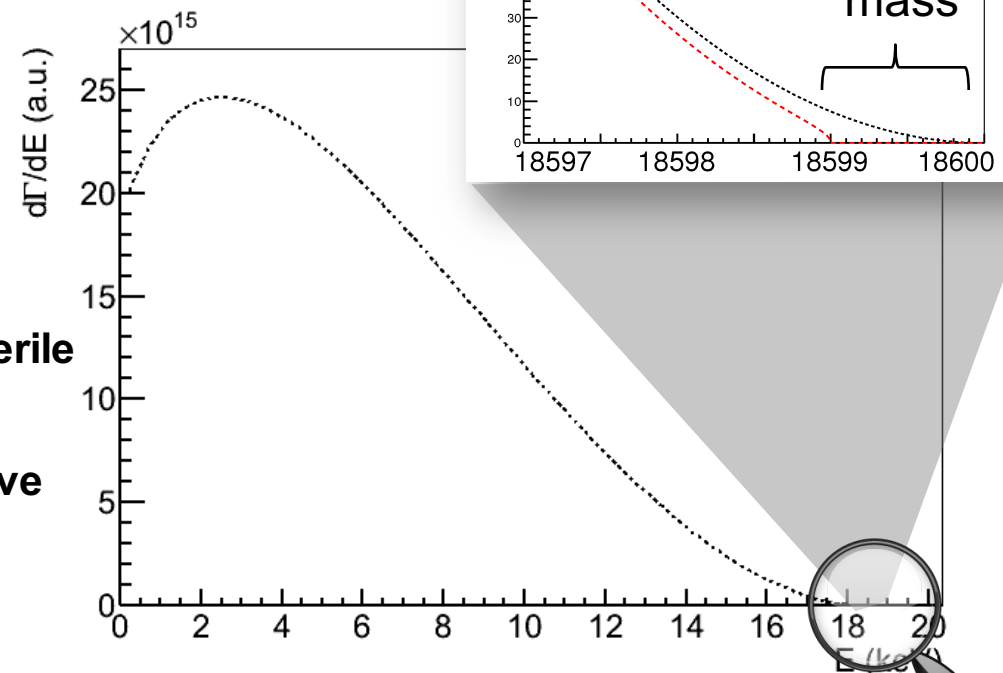


$$\frac{d\Gamma}{dE} = C \cdot F(E, Z) \cdot p \cdot (E + m_e) \cdot (E - E_0) \cdot \sum_i |U_{ei}|^2 \sqrt{(E - E_0)^2 - m_{\nu_i}^2}$$

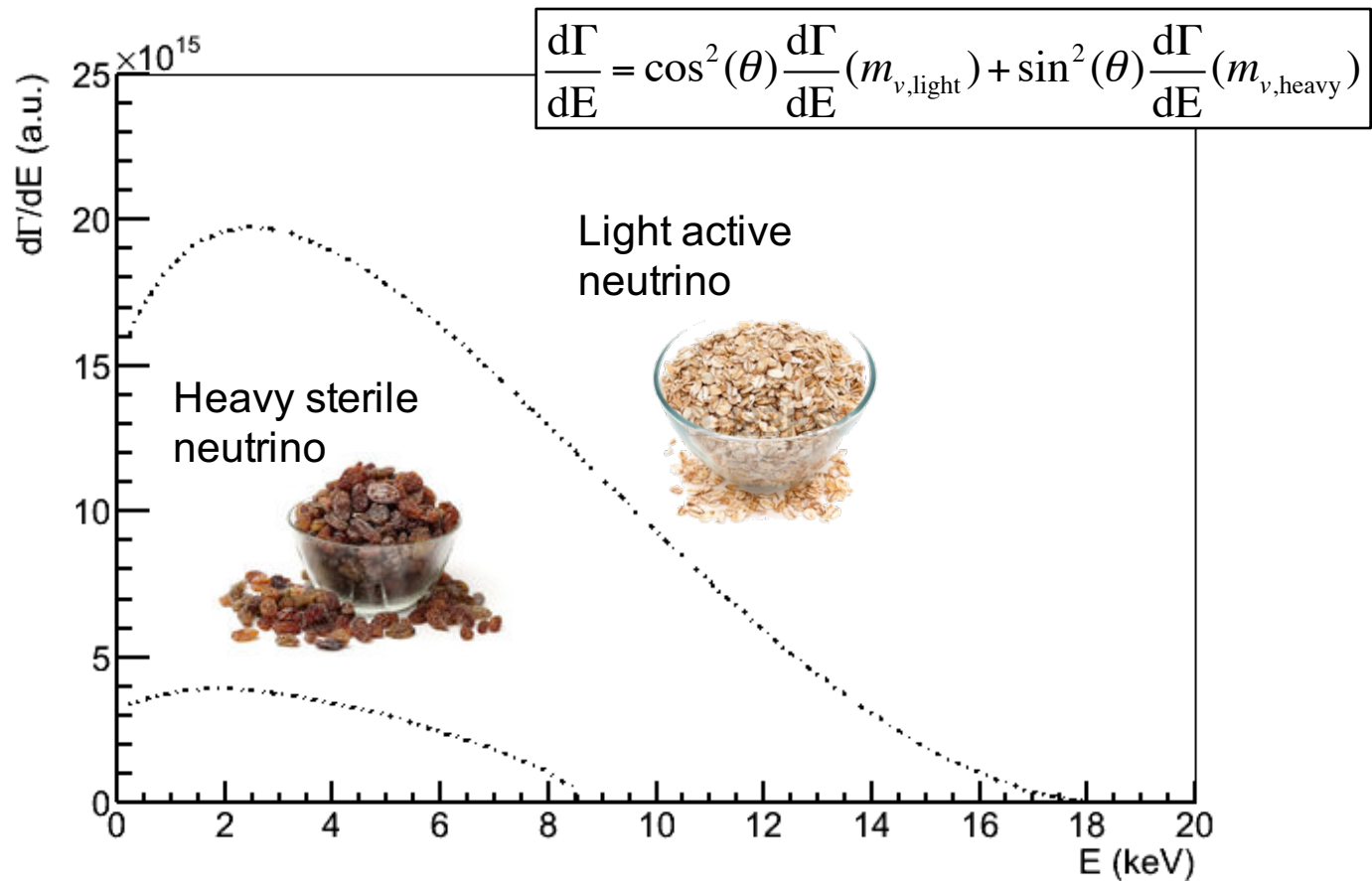
# Tritium beta decay



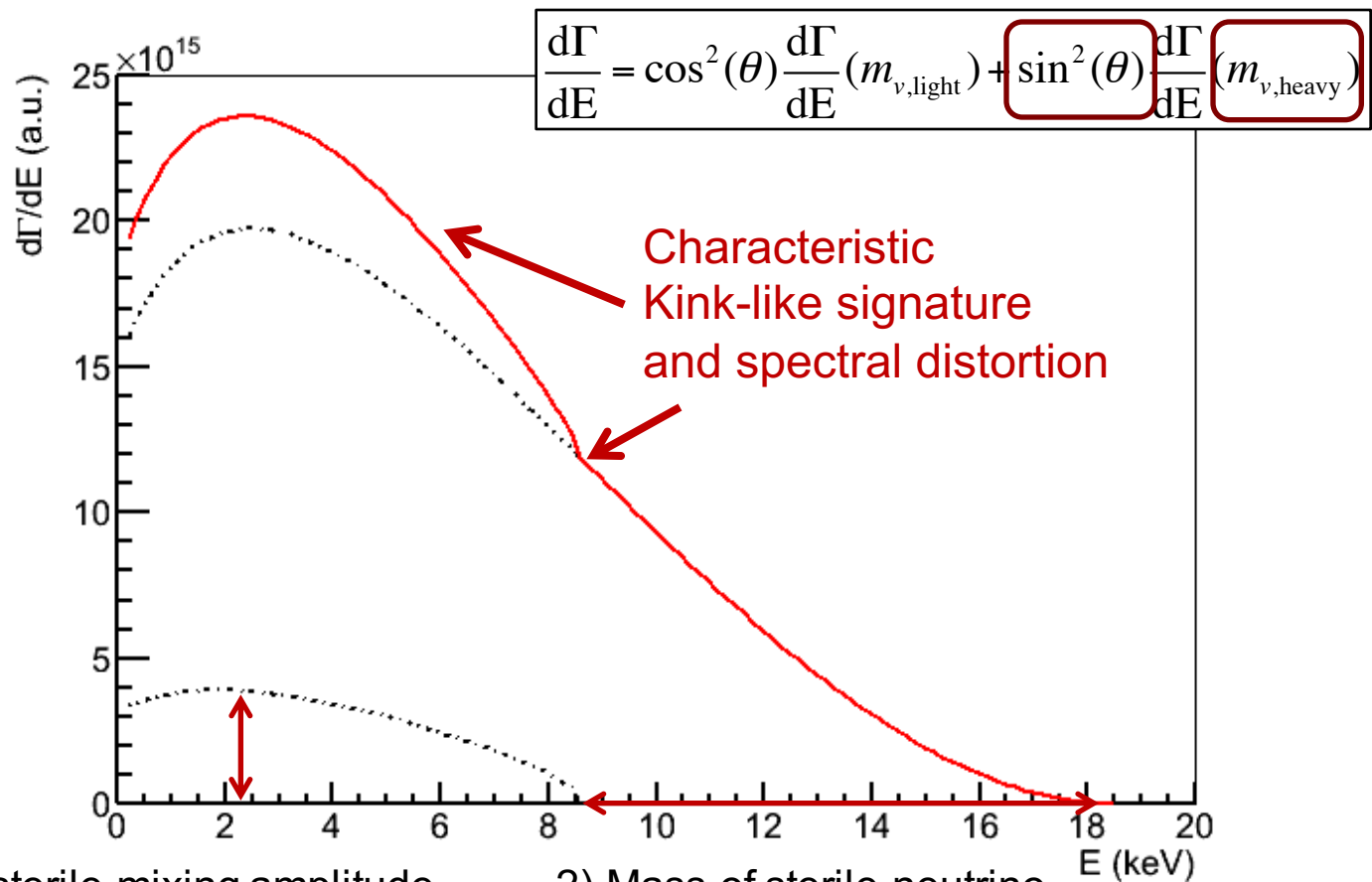
Neutrino from tritium  $\beta$ -decay



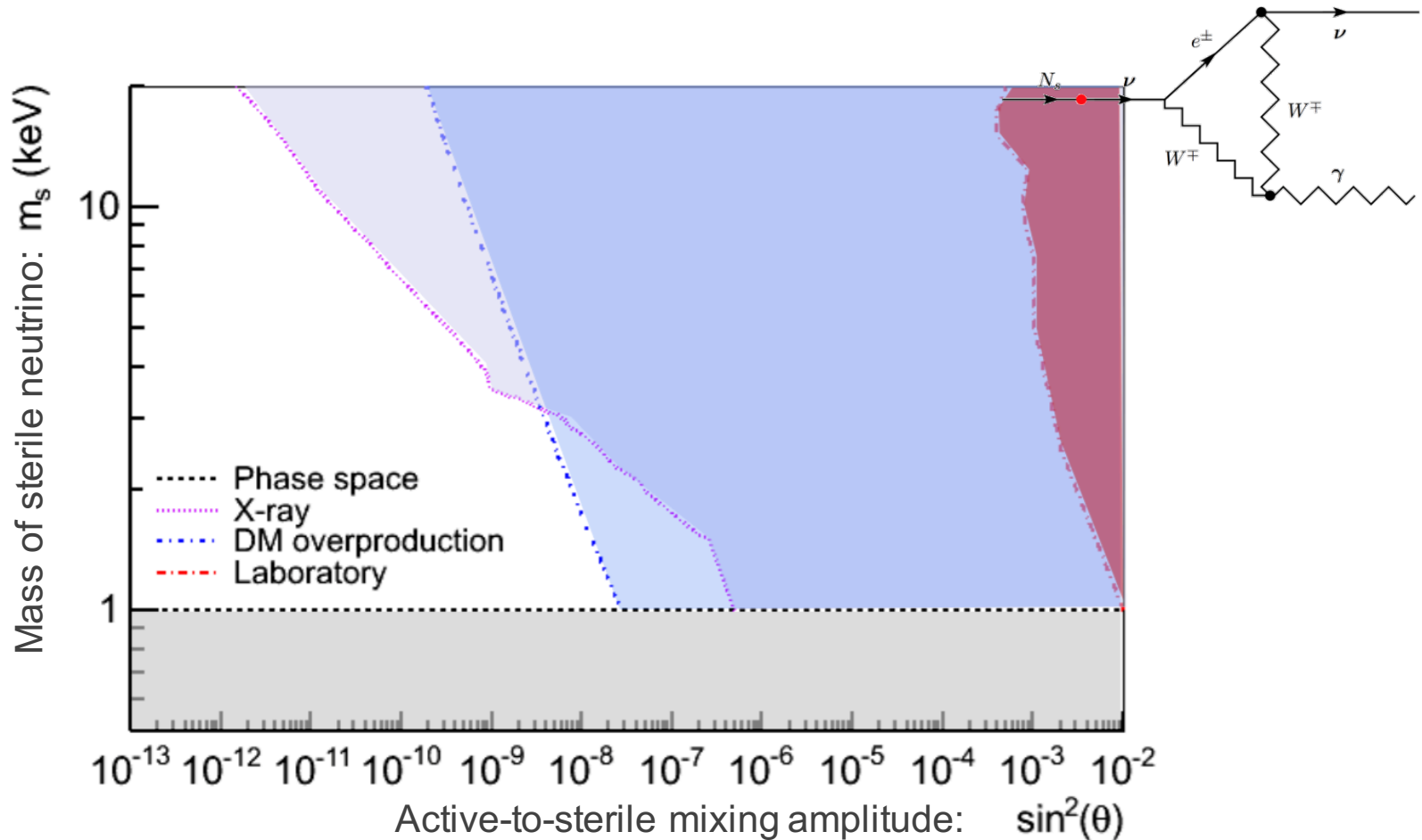
# Imprint of sterile $\nu$ 's on $\beta$ -spectrum



# Imprint of sterile $\nu$ 's on $\beta$ -spectrum

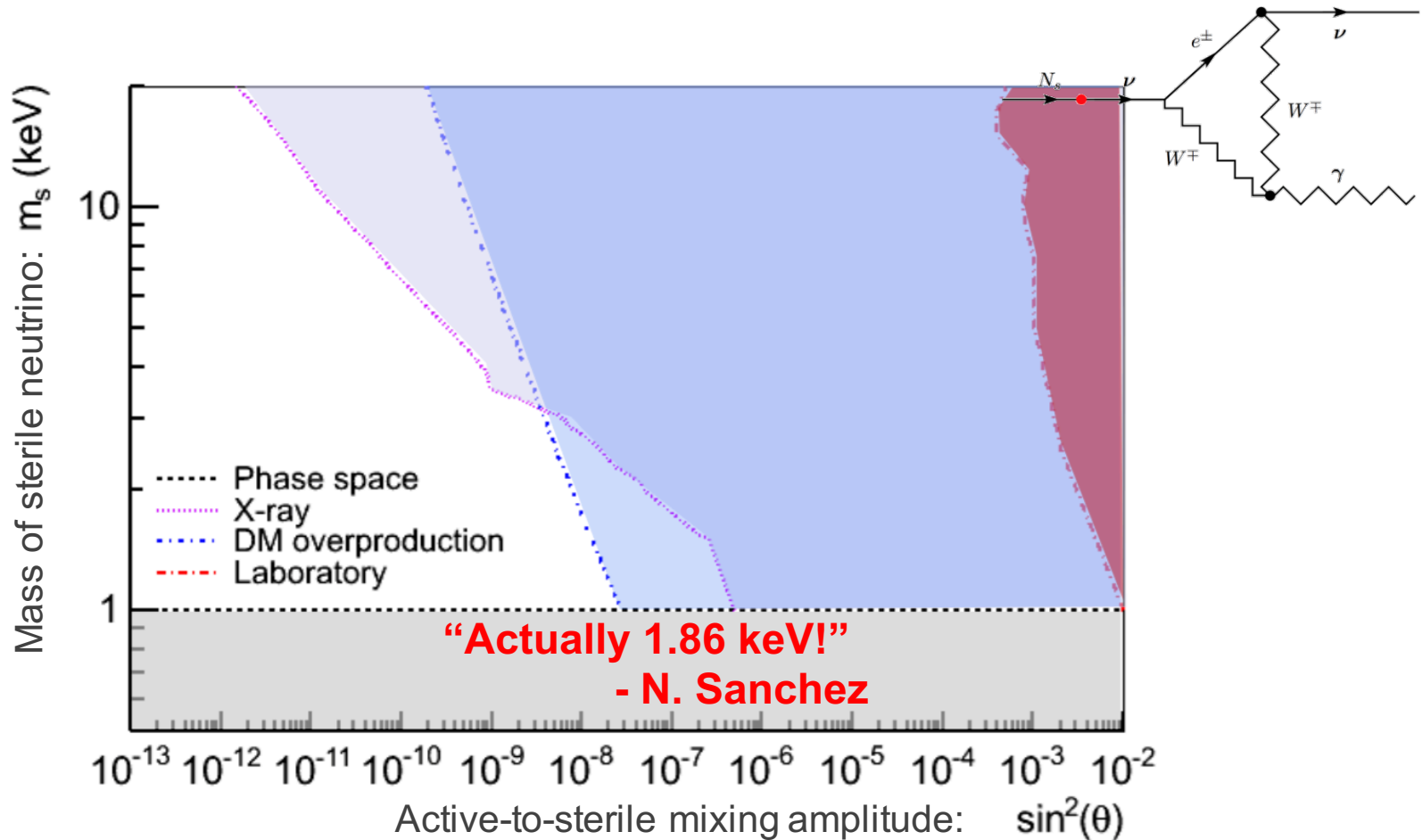


# Cosmological constraints



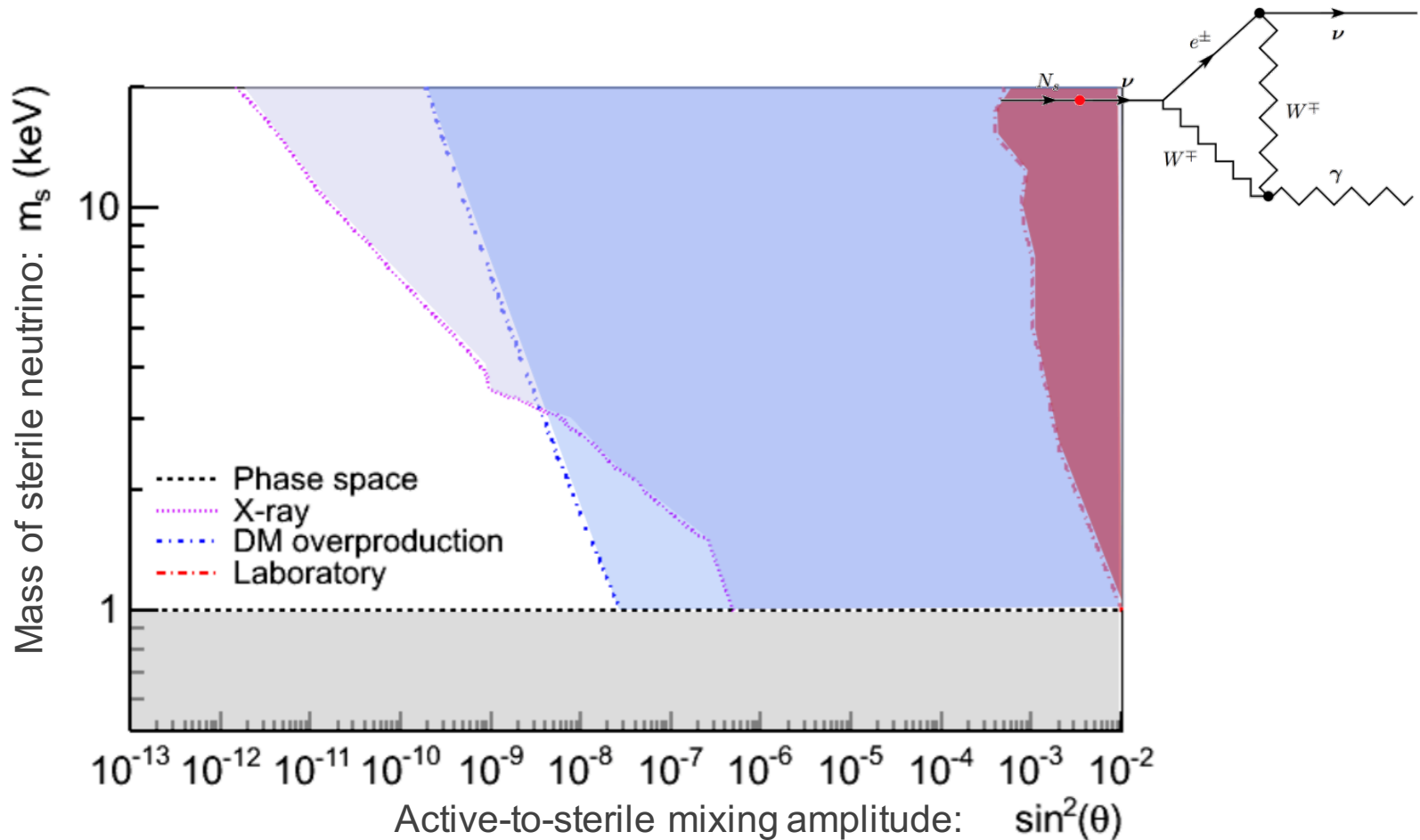
O. Ruchayskiy, A. Ivashko  
JHEP 1206 (2012) 100

# Cosmological constraints



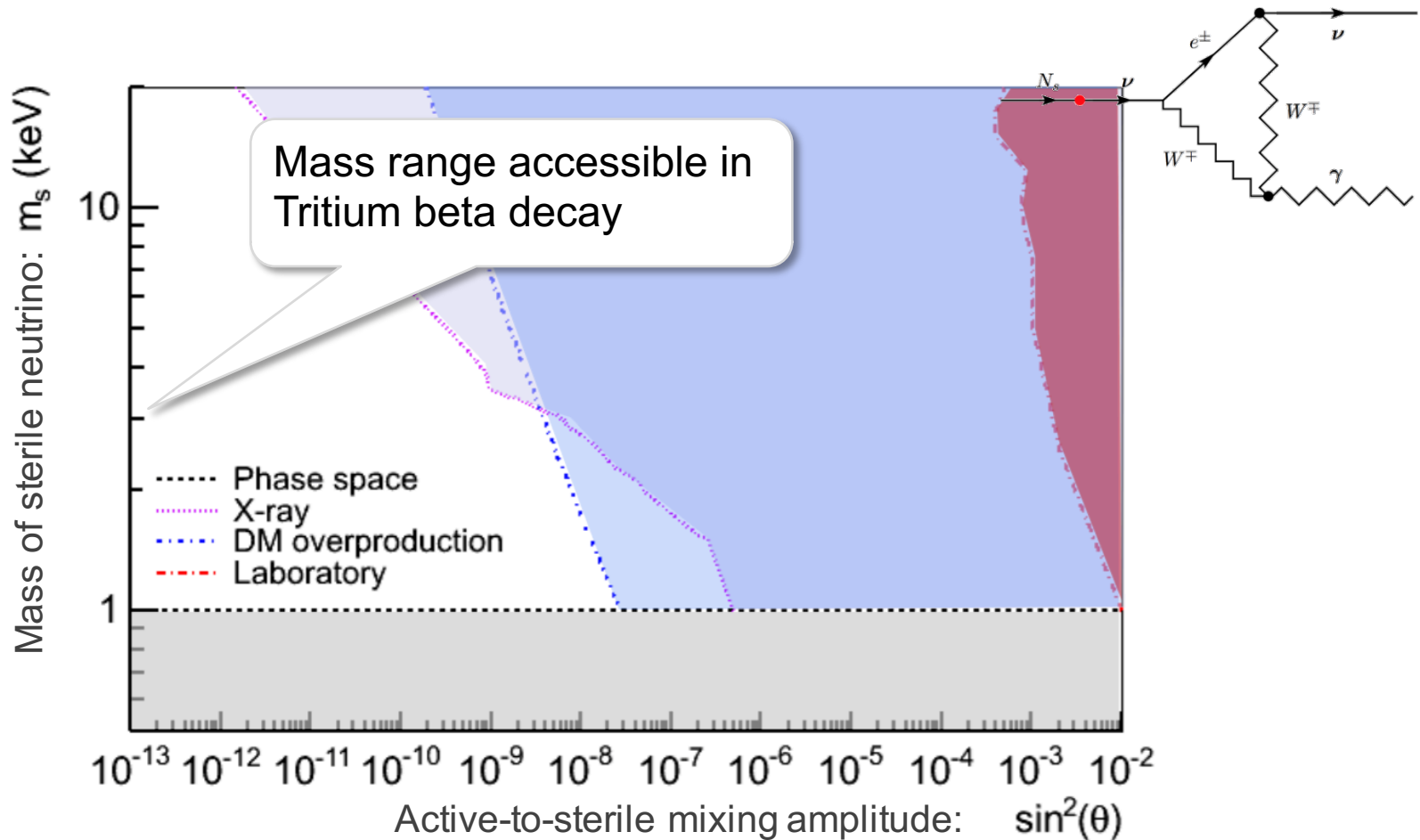
O. Ruchayskiy, A. Ivashko  
JHEP 1206 (2012) 100

# Cosmological constraints



O. Ruchayskiy, A. Ivashko  
JHEP 1206 (2012) 100

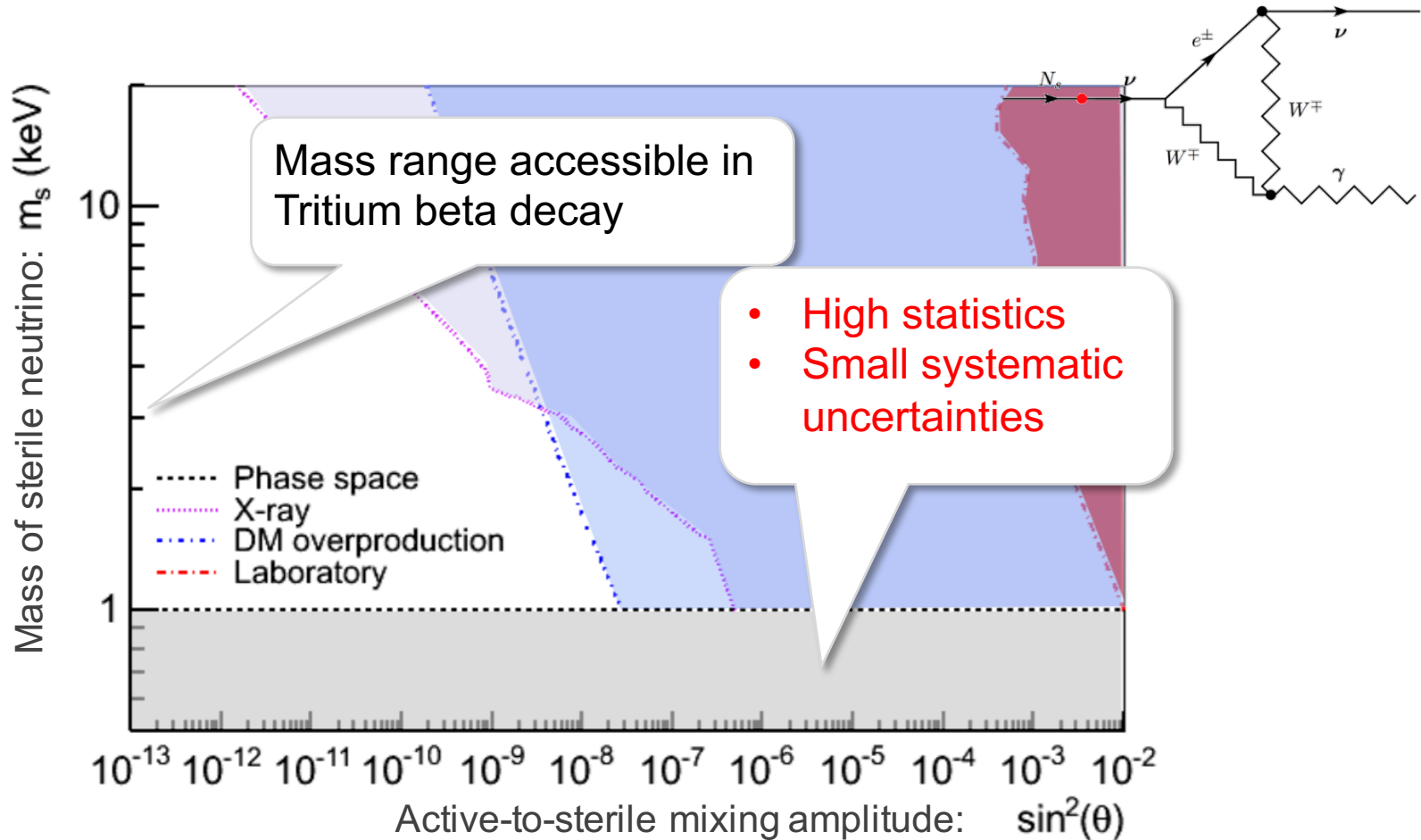
# Cosmological constraints



O. Ruchayskiy, A. Ivashko  
JHEP 1206 (2012) 100



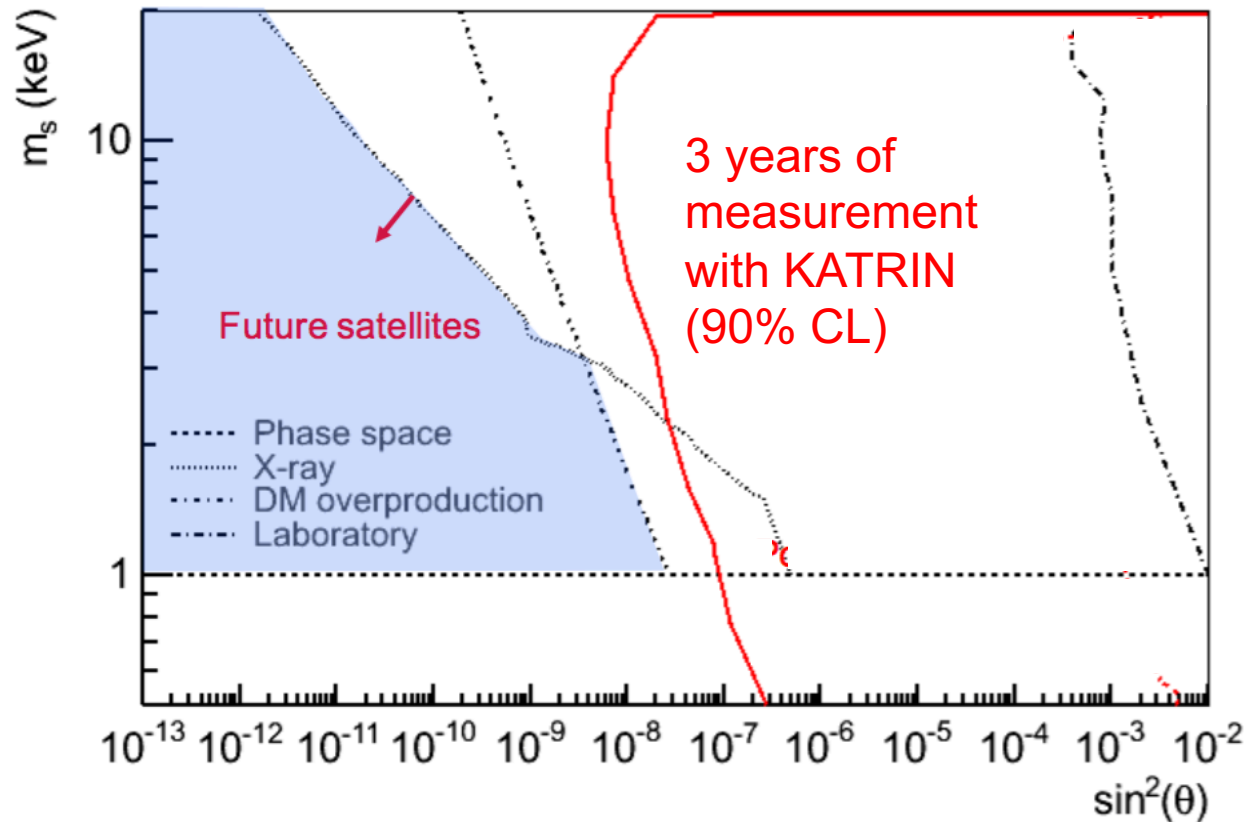
# The challenge of sterile $\nu$ search



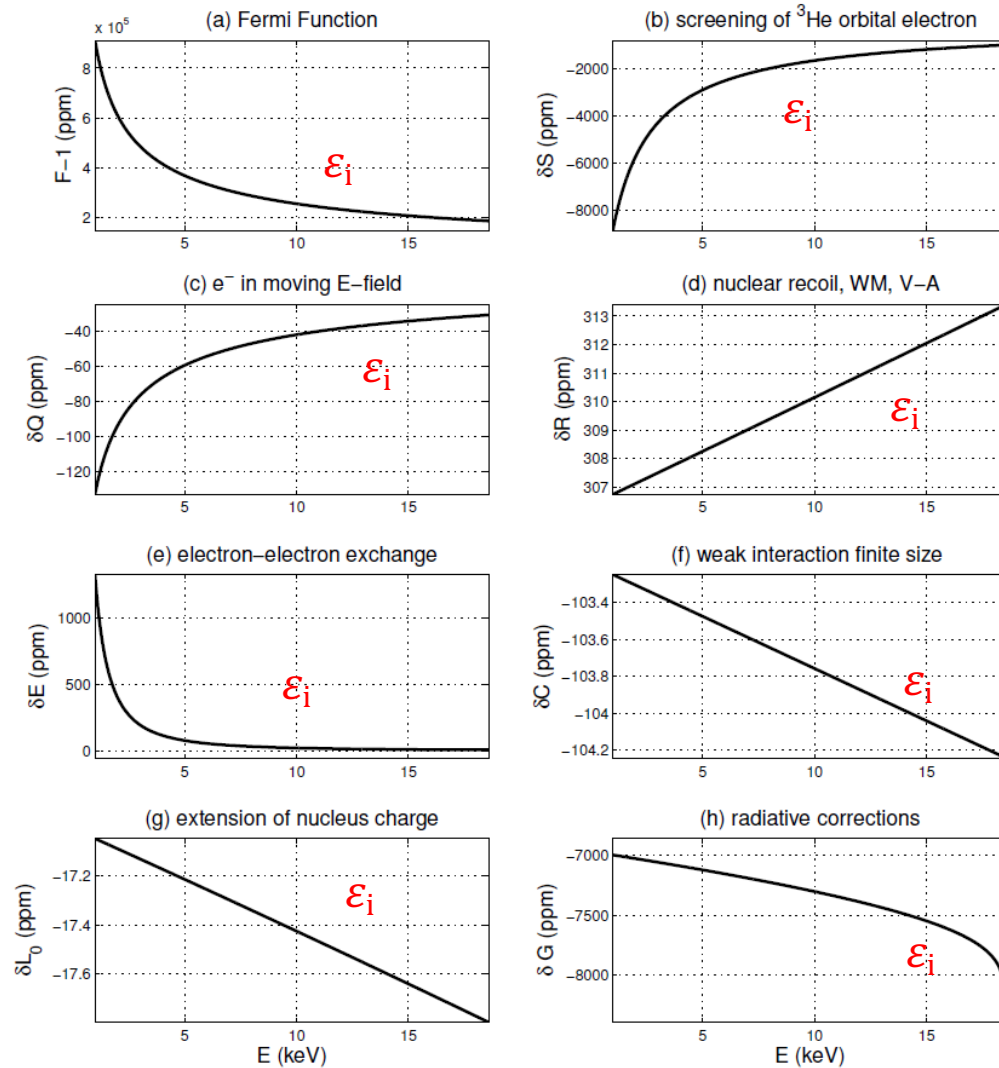
O. Ruchayskiy, A. Ivashko  
JHEP 1206 (2012) 100

# Statistical sensitivity

PRELIMINARY



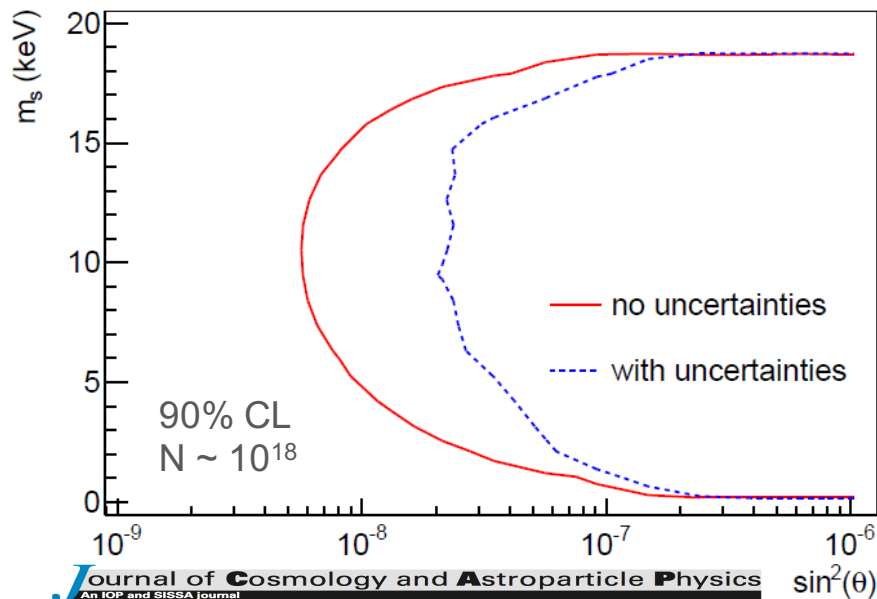
# Theoretical corrections to the $\beta$ -spectrum



# Detailed sensitivity studies

## Spectral fit approach:

„How do theoretical uncertainties impact the sensitivity to find the signature of a sterile neutrino ?“



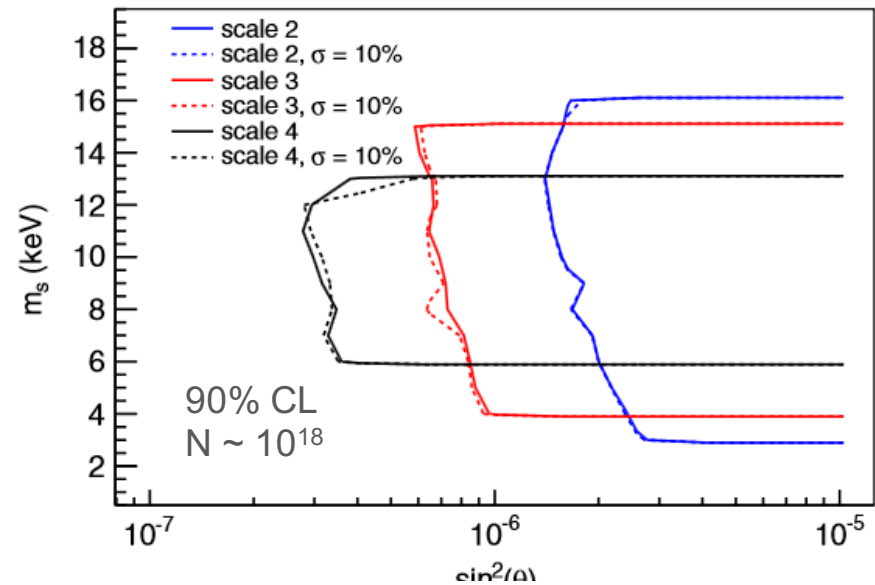
Journal of Cosmology and Astroparticle Physics  
 An IOP and SISSA journal

### Sensitivity of next-generation tritium beta-decay experiments for keV-scale sterile neutrinos

S. Mertens,<sup>a,d</sup> T. Lasserre,<sup>b,c</sup> S. Groh,<sup>d</sup> G. Drexlin,<sup>d</sup> F. Glück,<sup>d,f</sup>  
 A. Huber,<sup>d</sup> A.W.P. Poon,<sup>a</sup> M. Steidl,<sup>d</sup> N. Steinbrink<sup>e</sup>  
 and C. Weinheimer<sup>e</sup>

## Wavelet approach:

„Is a precise knowledge of the spectrum necessary to find the signature of a sterile neutrino ?“



### Wavelet Approach to Search for Sterile Neutrinos in Tritium $\beta$ -Decay Spectra

S. Mertens,<sup>1,2</sup> K. Dolde,<sup>2</sup> M. Korzeczek,<sup>2</sup> F. Glueck,<sup>2,3</sup> S. Groh,<sup>2</sup> R. D. Martin,<sup>1,\*</sup> A. W. P. Poon,<sup>1</sup> and M. Steidl<sup>2</sup>

<sup>1</sup>Institute for Nuclear and Particle Astrophysics,

Nuclear Science Division, Lawrence Berkeley National Laboratory, USA

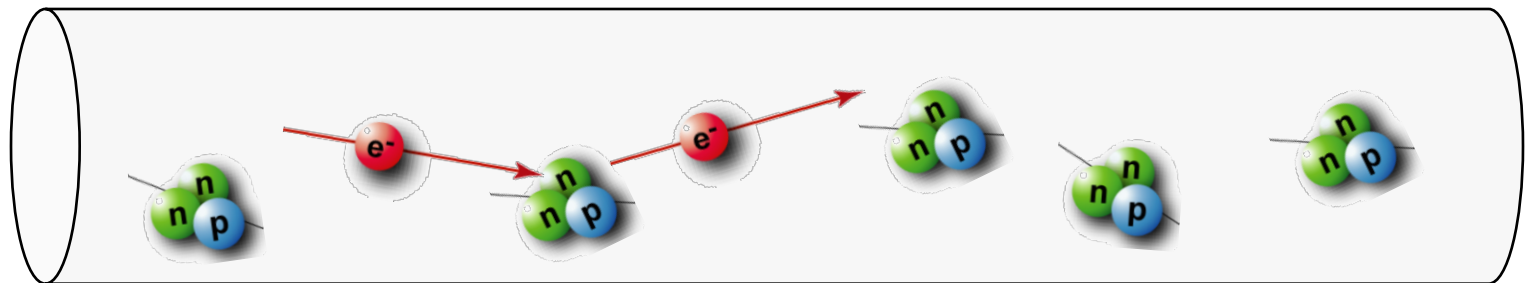
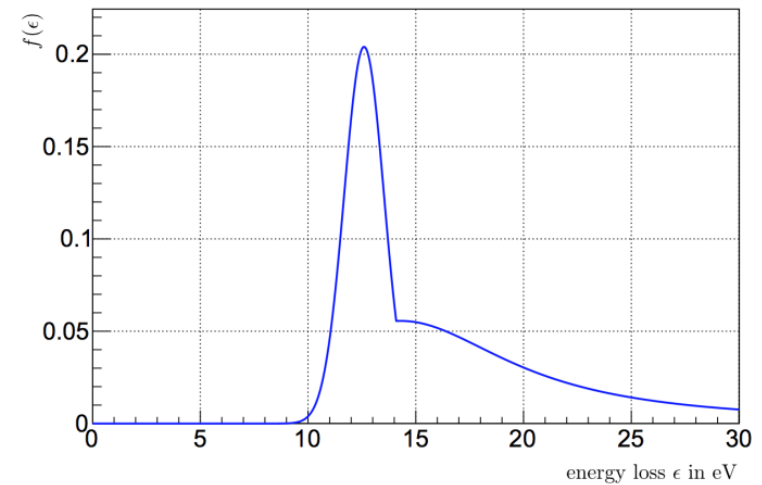
<sup>2</sup>Institute for Nuclear Physics (IKP), Karlsruhe Institute of Technology, Germany

<sup>3</sup>Wigner Research Institute for Physics, P. O. B. 49, H-1525 Budapest, Hungary

# Ongoing sensitivity studies

Systematic effects related to :

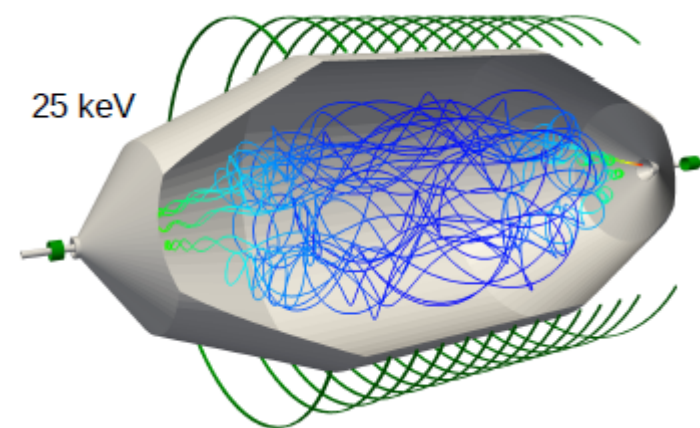
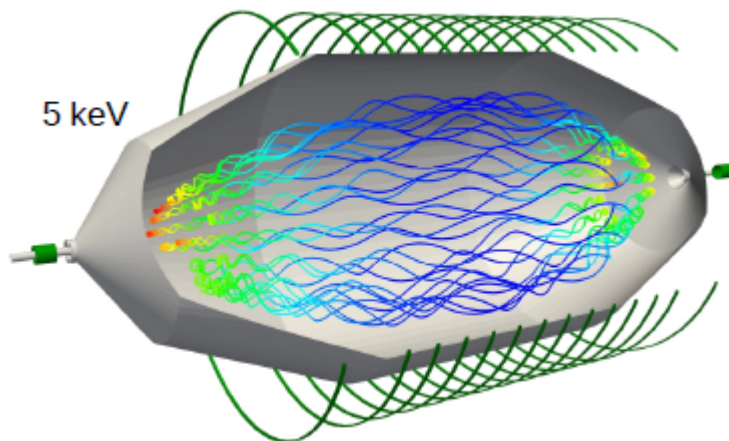
- Source Section
  - scattered electrons arrive at detector



# Ongoing sensitivity studies

Systematic effects related to :

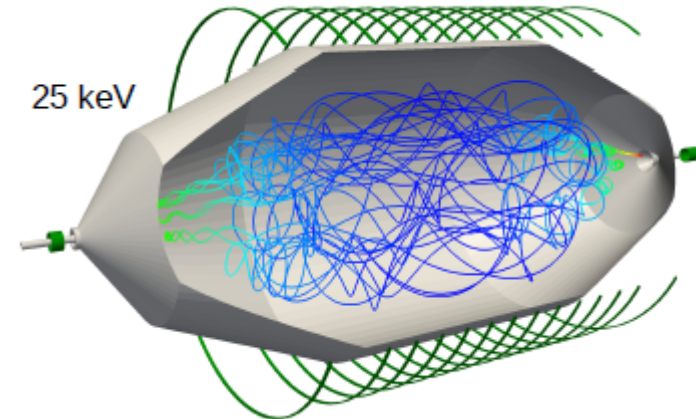
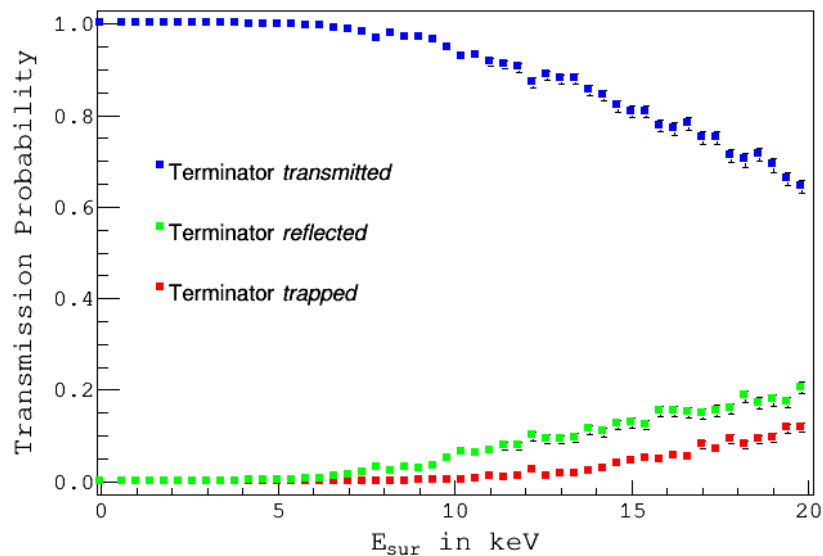
- Source Section
  - scattered electrons arrive at detector
- Spectrometer Section
  - electrons pass through spectrometer with high surplus energy



# Ongoing sensitivity studies

Systematic effects related to :

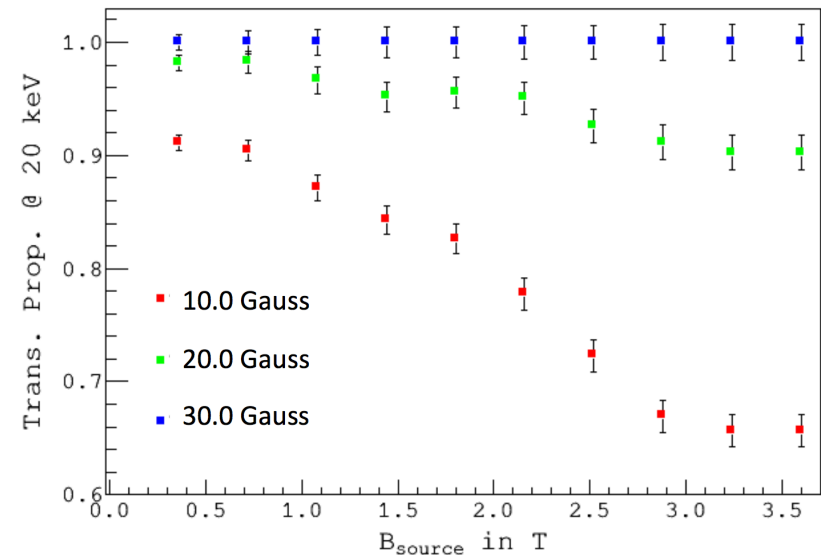
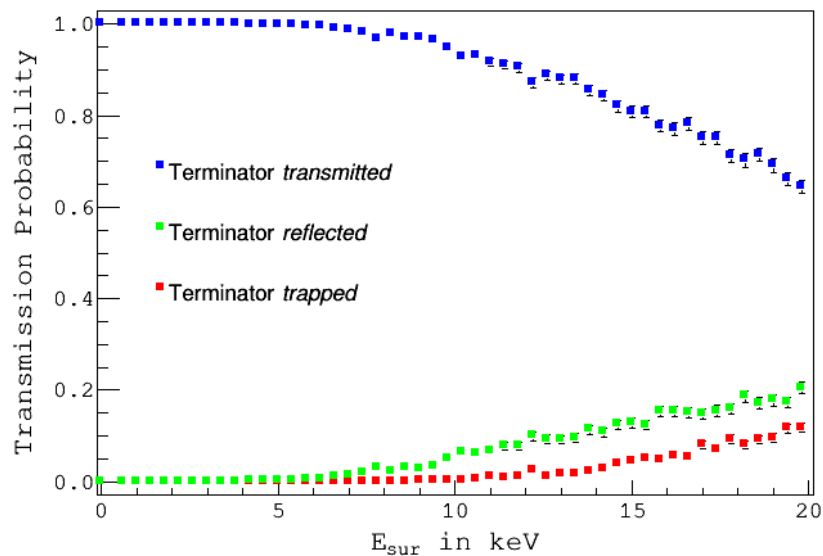
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# Ongoing sensitivity studies

Systematic effects related to :

- Source Section
  - scattered electrons arrive at detector
- Spectrometer Section
  - electrons pass through spectrometer with high surplus energy





# Ongoing sensitivity studies

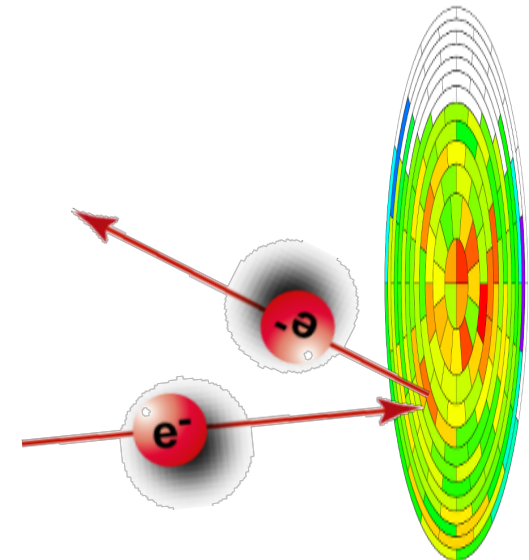
Systematic effects related to :

- Source Section
  - scattered electrons arrive at detector
- Spectrometer Section
  - electrons pass through spectrometer with high surplus energy
- Detector Section
  - Backscattering
  - Charge sharing
  - Pile-up
  - Etc.

**completed Anton Huber  
Masters Thesis**

**completed Kai Dolde  
Masters Thesis**

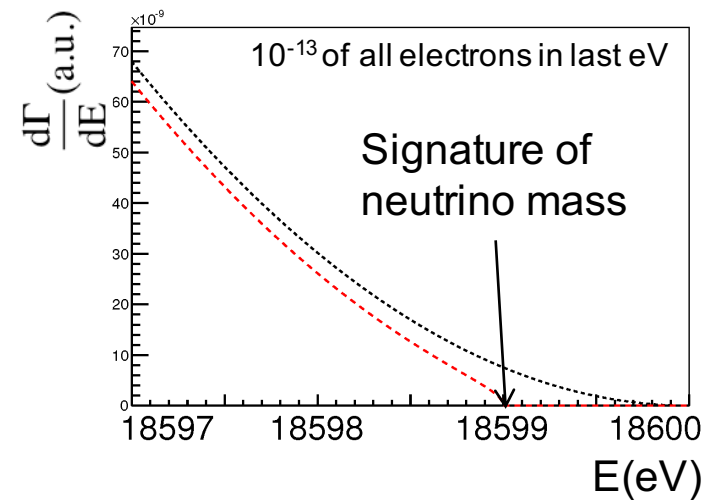
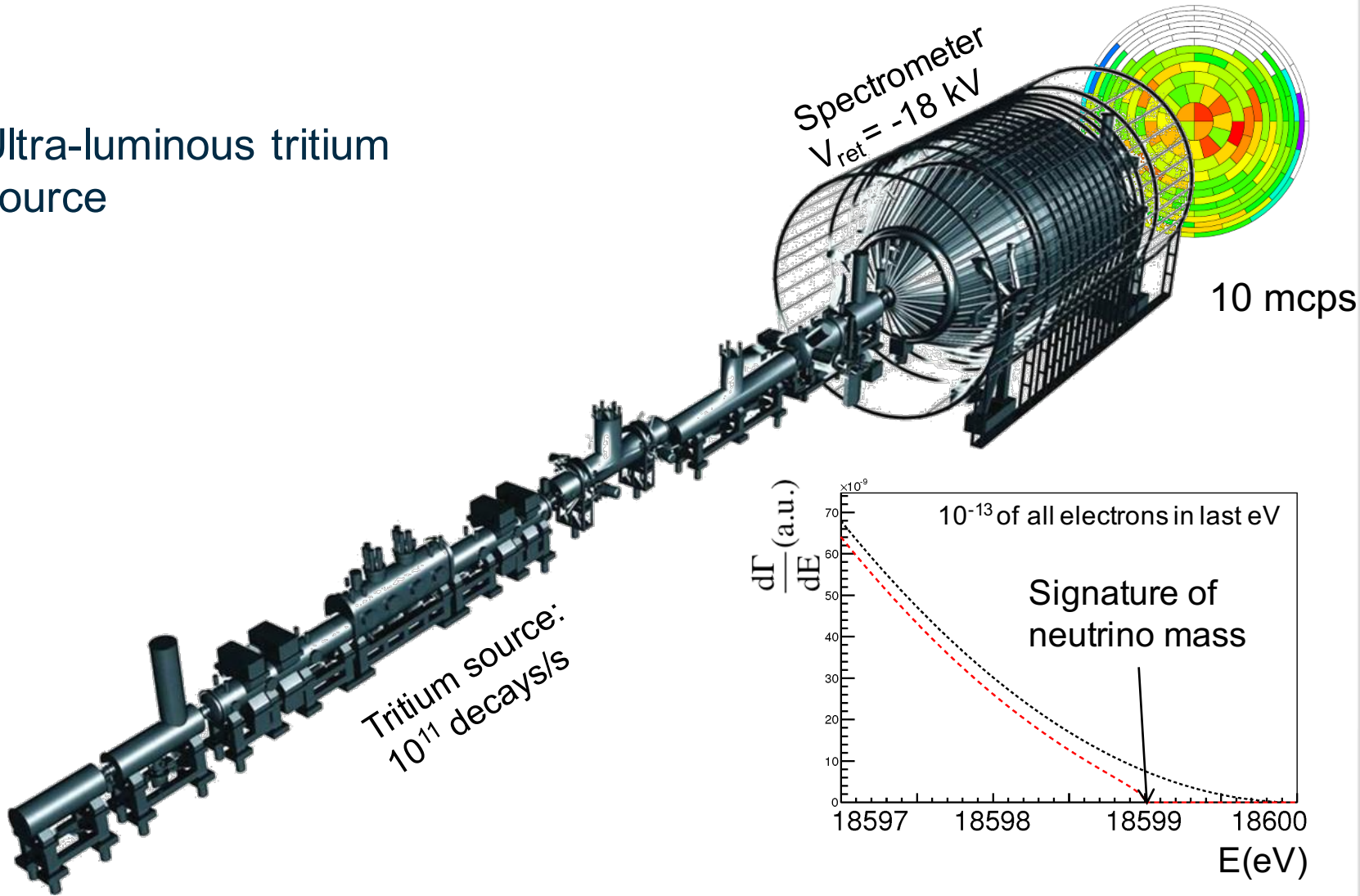
**... to be published soon**



# How to use KATRIN



Ultra-luminous tritium source



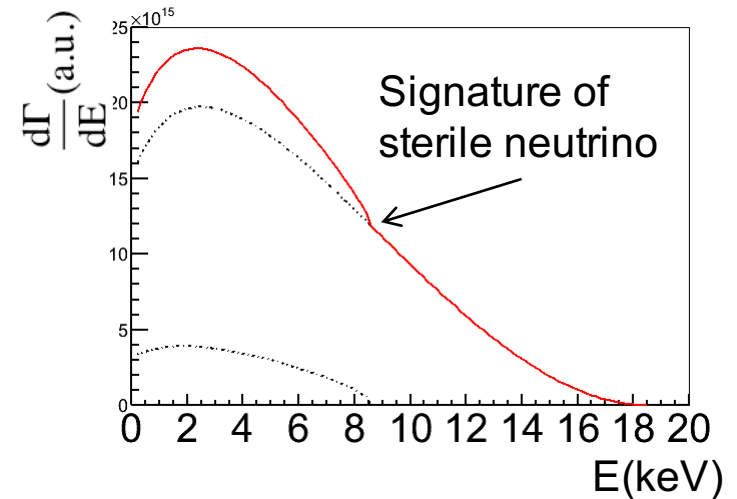
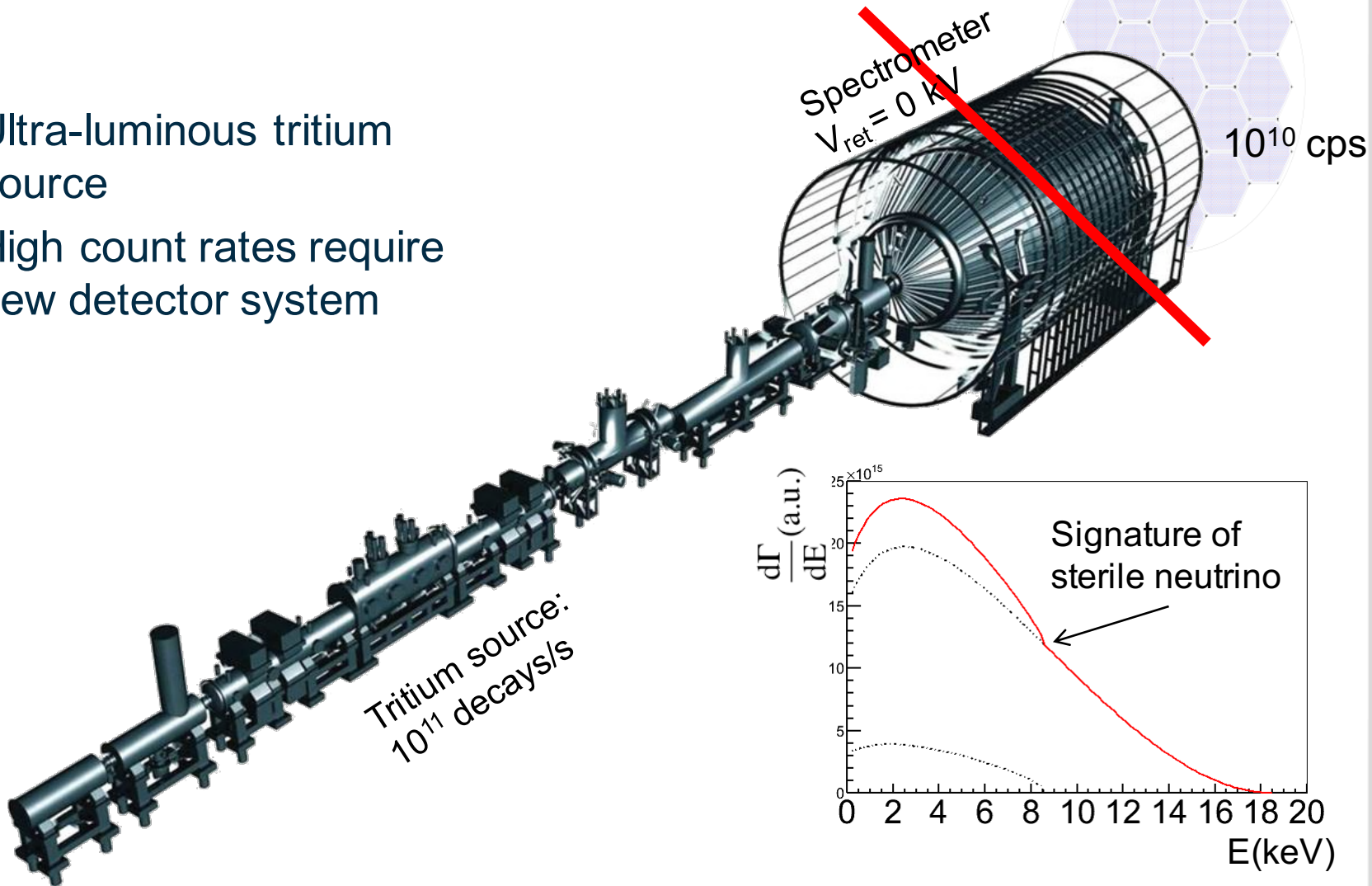
# How to use KATRIN



Ultra-luminous tritium source



High count rates require new detector system



# How to use KATRIN

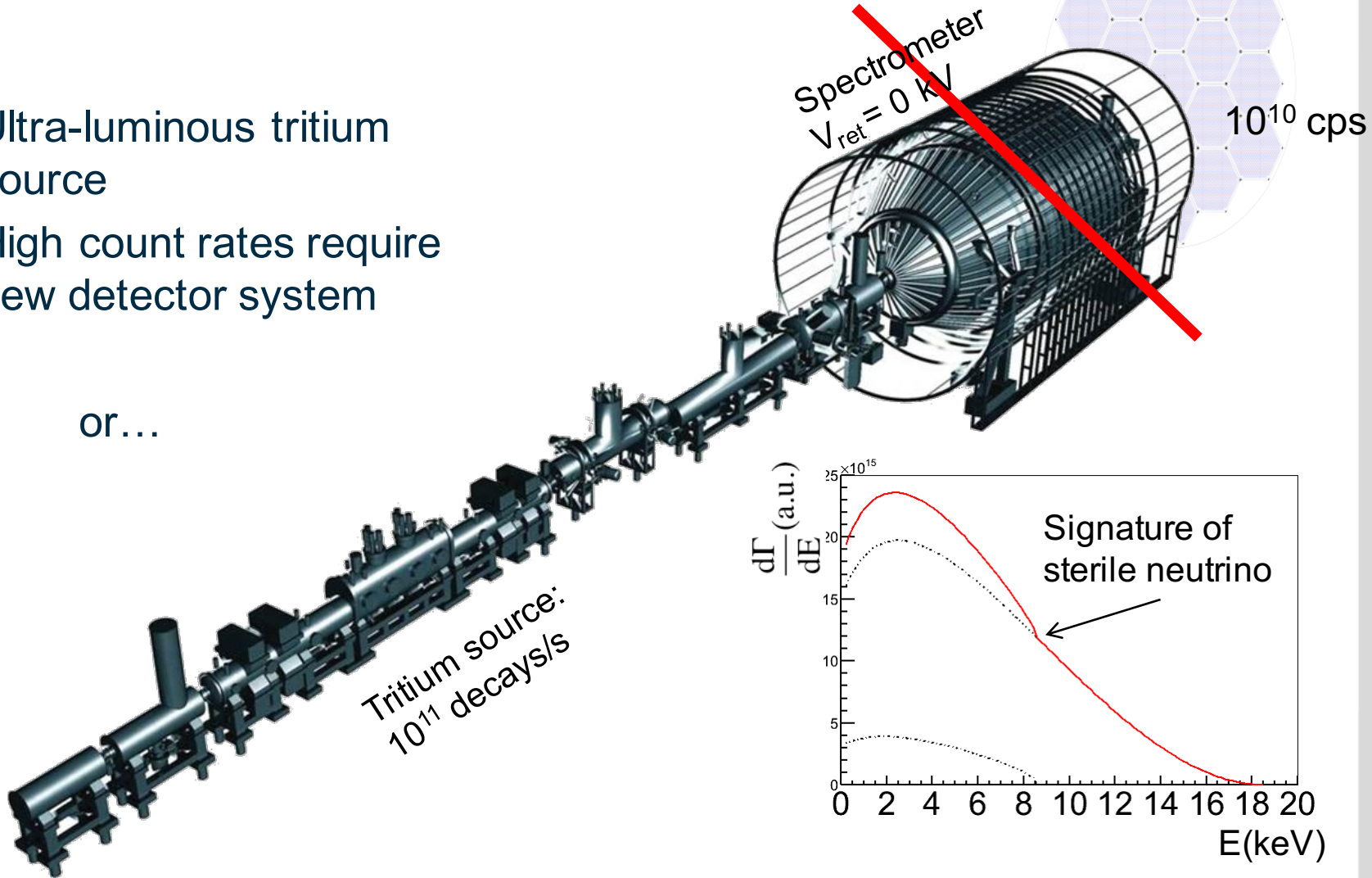


Ultra-luminous tritium source



High count rates require new detector system

or...



# How to use KATRIN

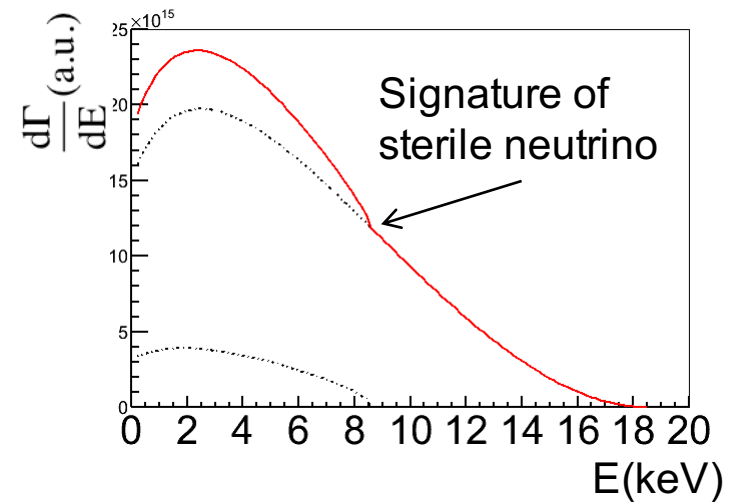
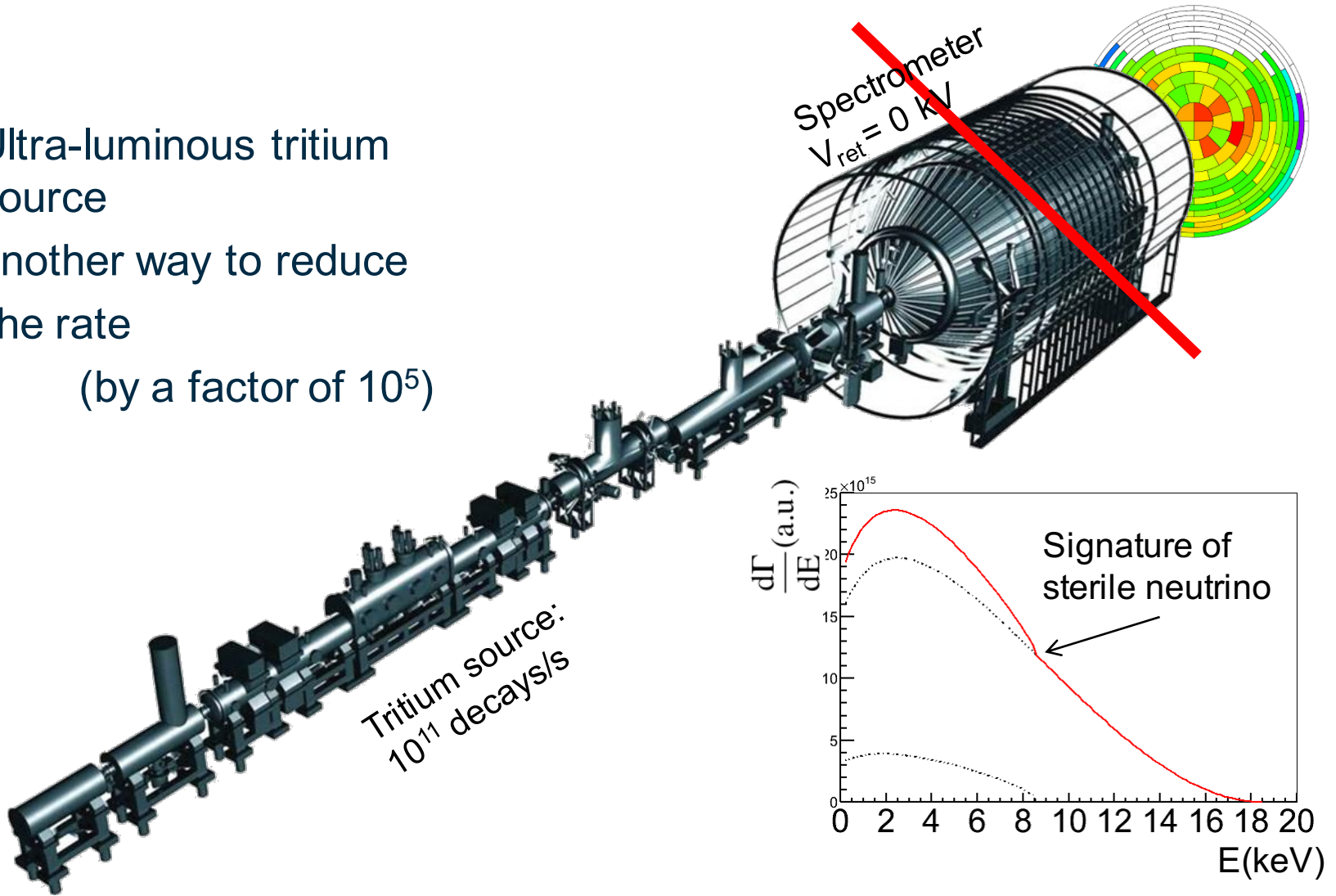


Ultra-luminous tritium source

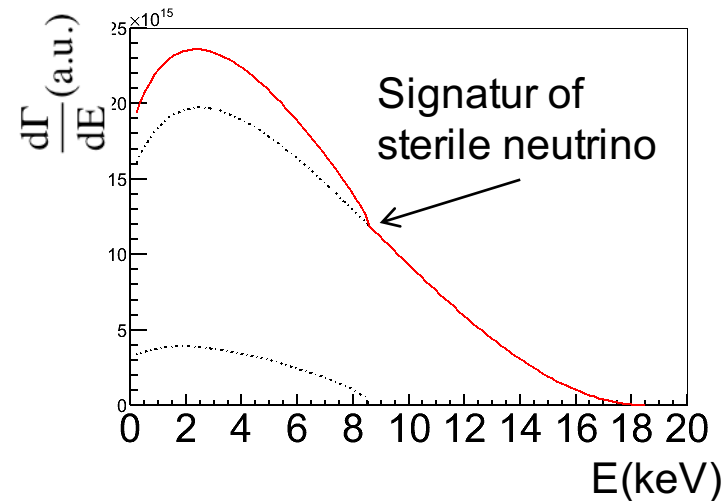


another way to reduce the rate

(by a factor of  $10^5$ )



# How to use KATRIN



A **Pre** KATRIN Measurement

- > will be performed very soon
- > requires a 'improvised' rate reduction

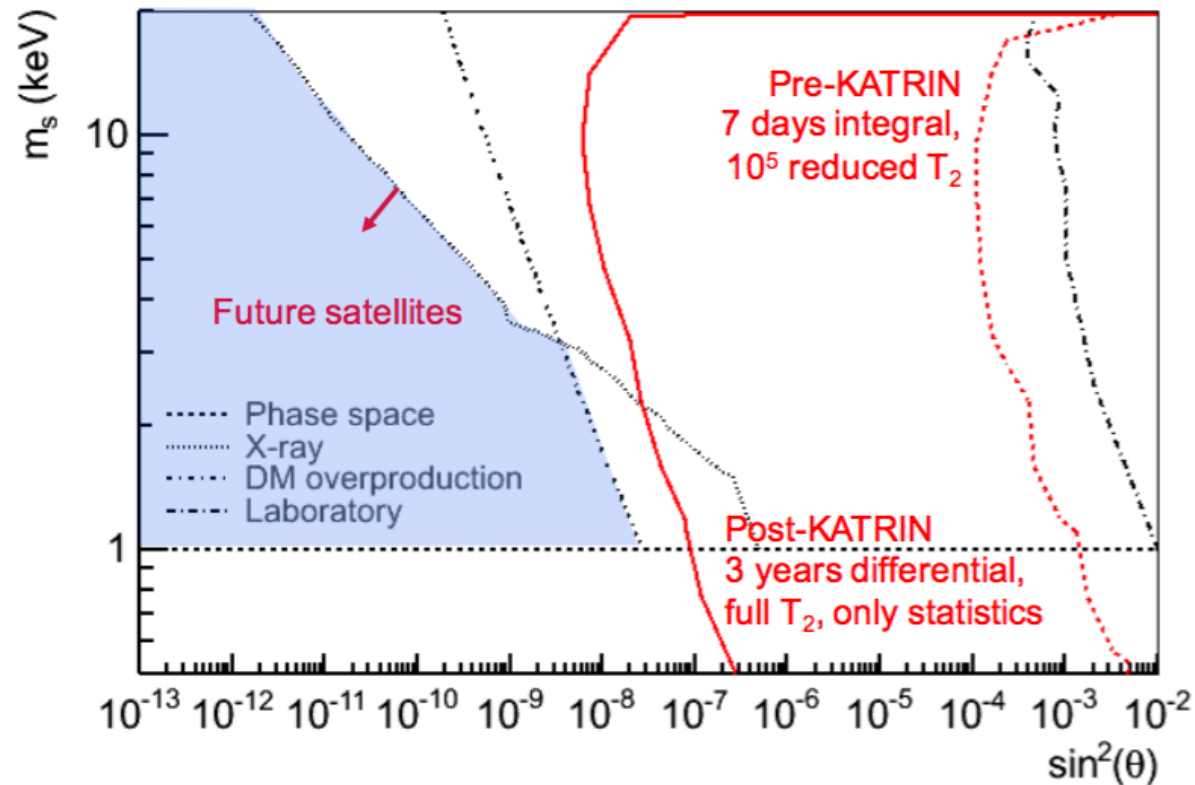
A **Post** KATRIN Measurement

- > requires a new detector system
- > will be performed after KATRIN determined the neutrino mass

# Preliminary Sensitivity

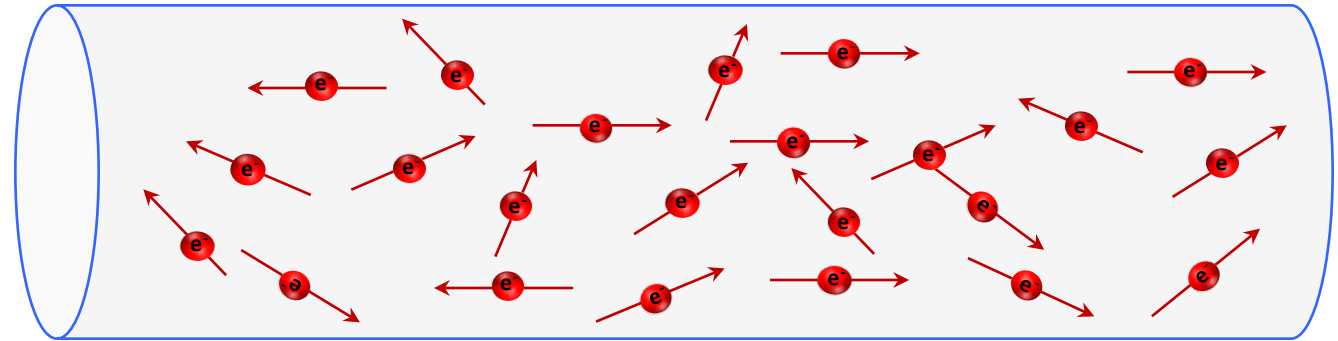
A **Pre** KATRIN Measurement

A **Post** KATRIN Measurement

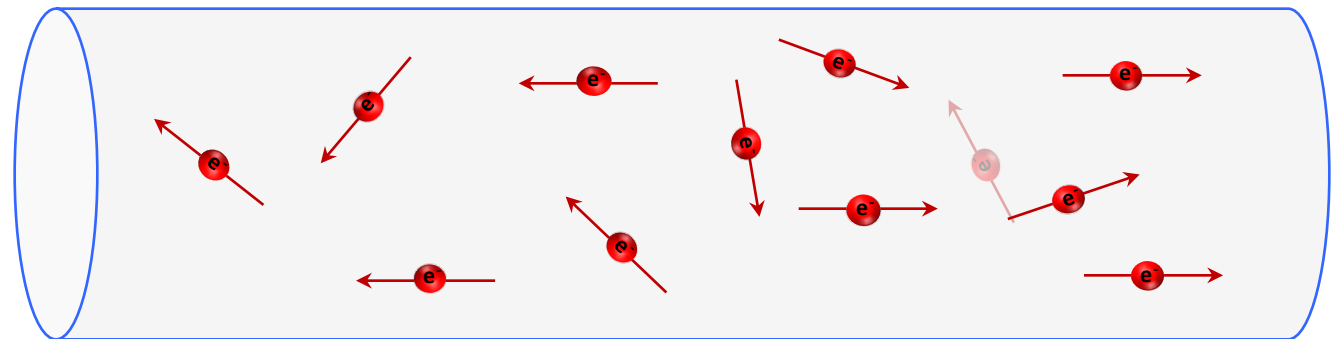


# How to reduce the Rate – Pre-Measurement

Standard  
KATRIN  
source



modified  
source



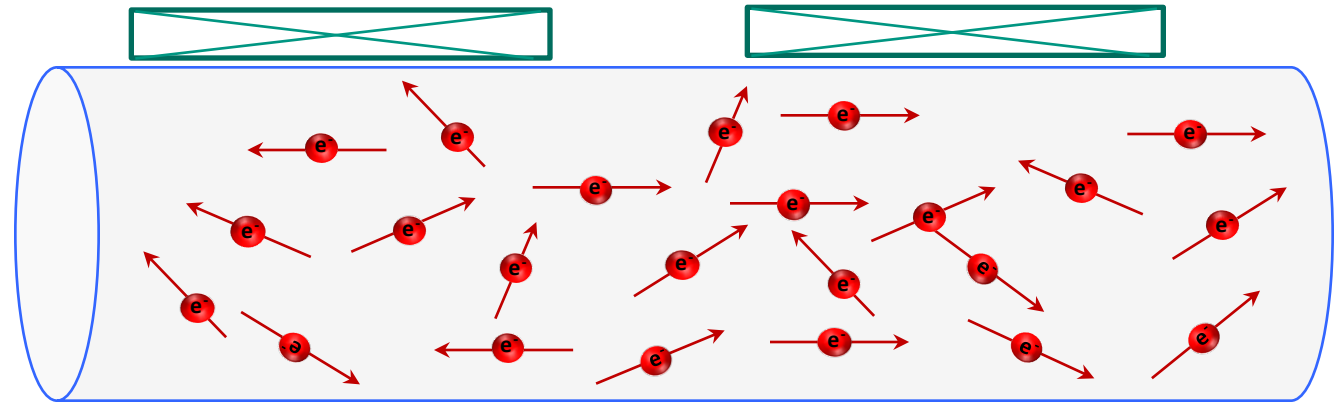
less tritium in  
source



# How to reduce the Rate – Pre-Measurement

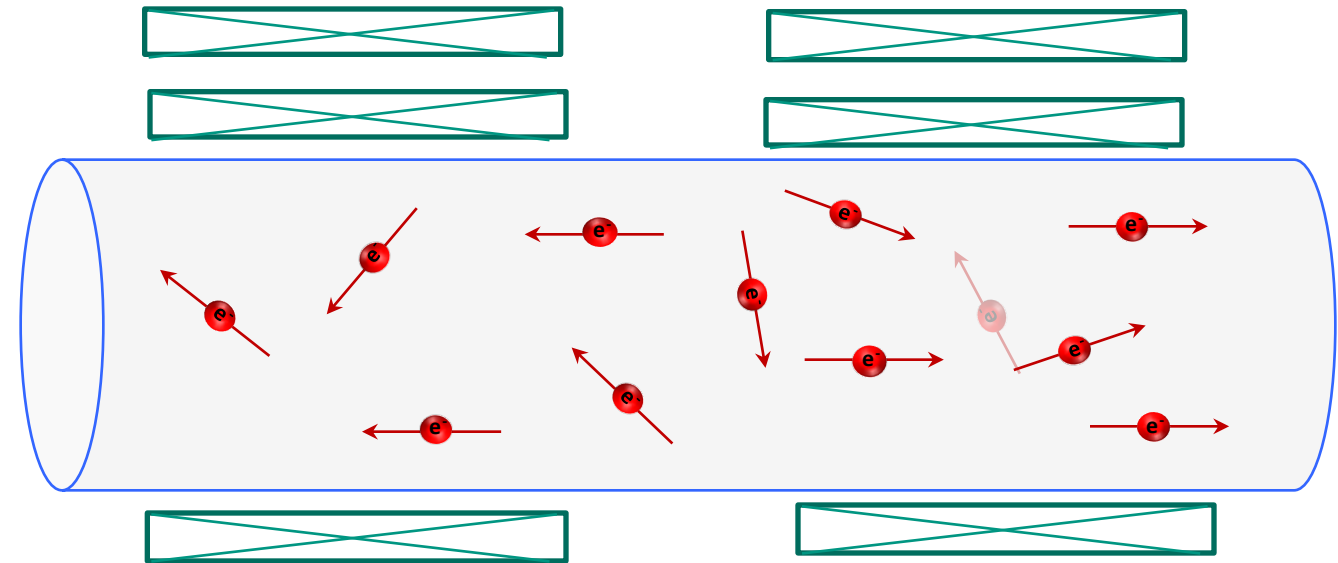
Standard KATRIN

$$B_{source} = 3.6 T$$

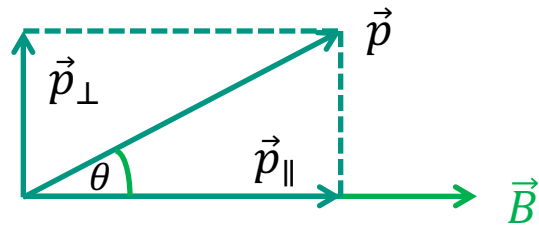


Modified source

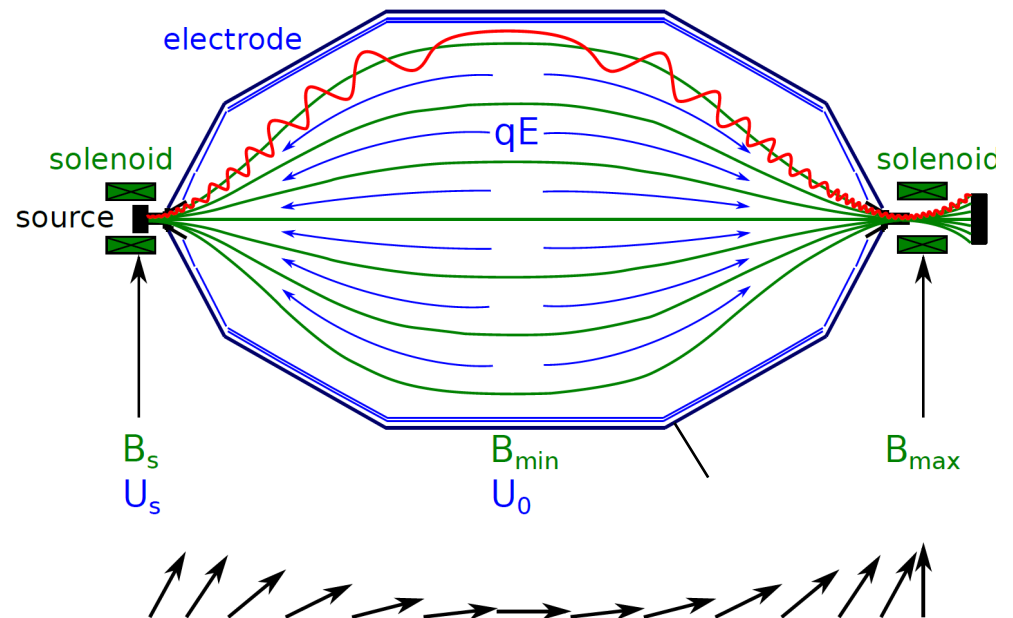
$$B_{source} = 0.03 T$$



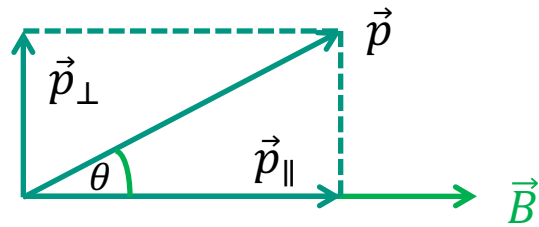
# How to reduce the Rate – Pre-Measurement



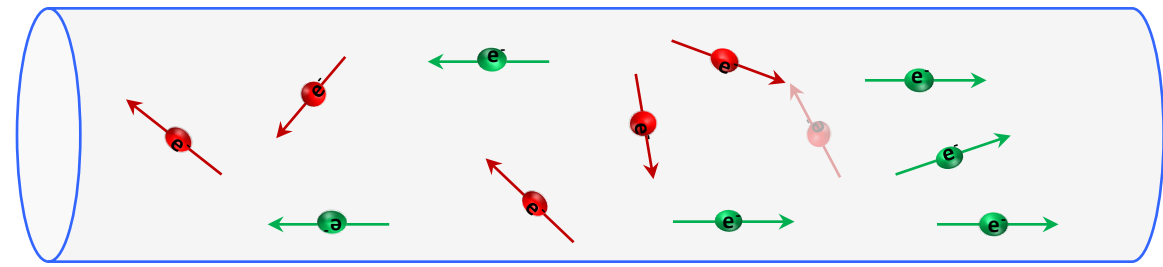
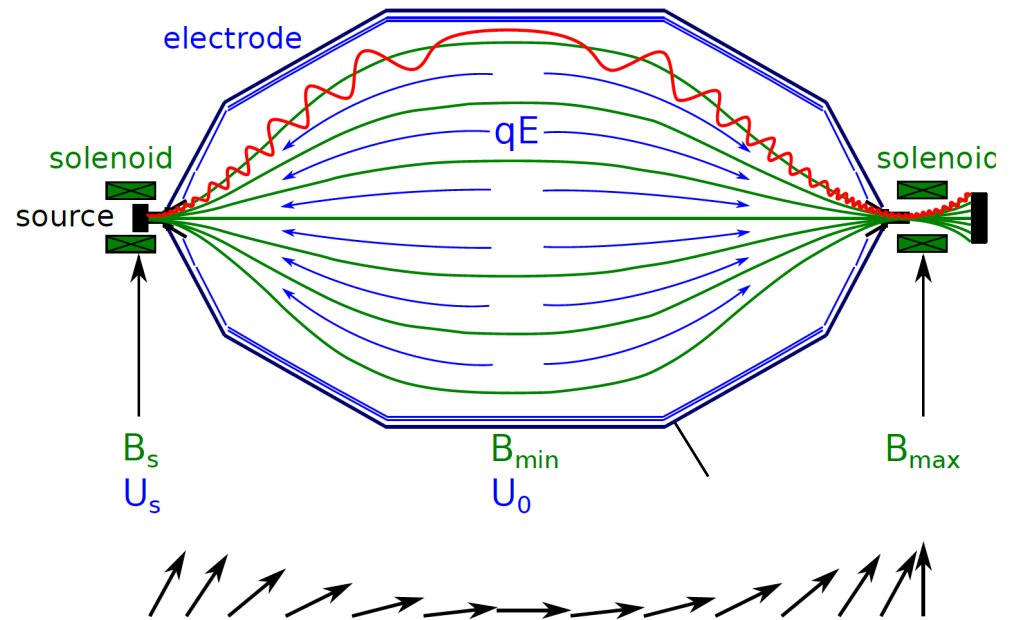
$$\theta_{max} = \sin^{-1} \sqrt{\frac{B_{ini}}{B_{max}}}$$



# How to reduce the Rate – Pre-Measurement

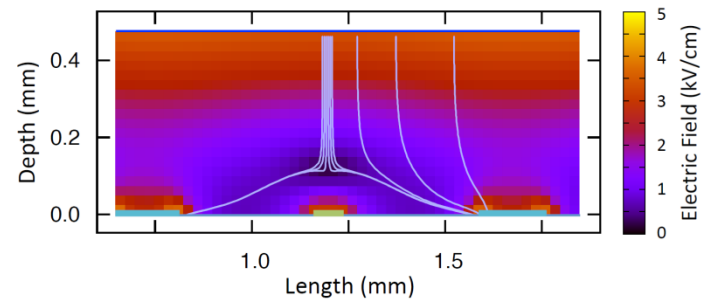
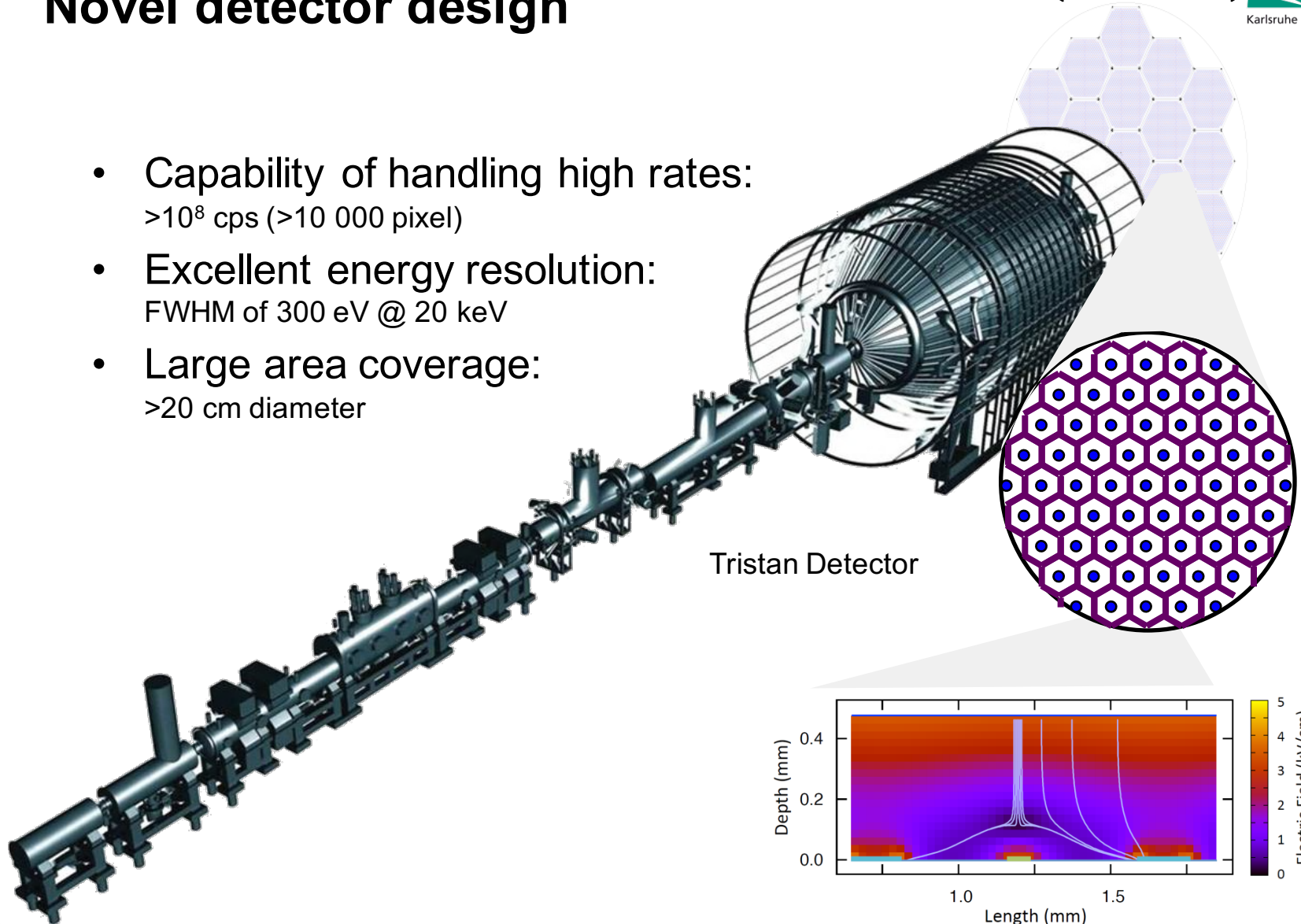


$$\theta_{max} = \sin^{-1} \sqrt{\frac{B_{ini}}{B_{max}}}$$



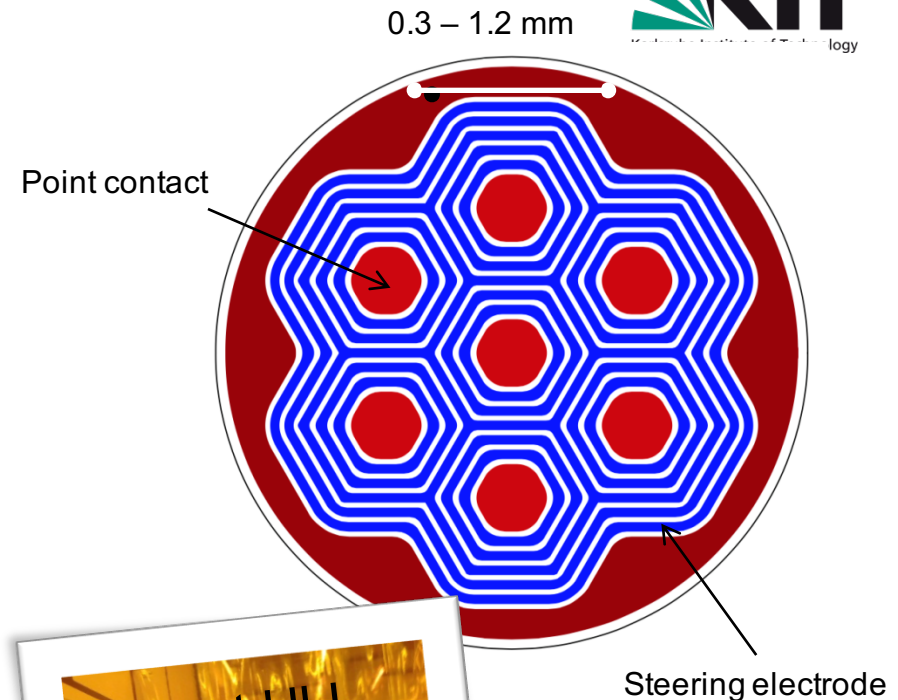
# Novel detector design

- Capability of handling high rates:  
>10<sup>8</sup> cps (>10 000 pixel)
- Excellent energy resolution:  
FWHM of 300 eV @ 20 keV
- Large area coverage:  
>20 cm diameter



# TRISTAN Prototype

- Key design features:
  - Very small point contacts
  - Thin entrance window (~10 nm)
  - Shared steering electrode
- Cooperations with Max-Planck Halbleiterlabor in Munich and Lawrence Berkeley Lab
- First prototype was built in October last year
- Characterize pile-up, backscattering, charge-sharing, etc.

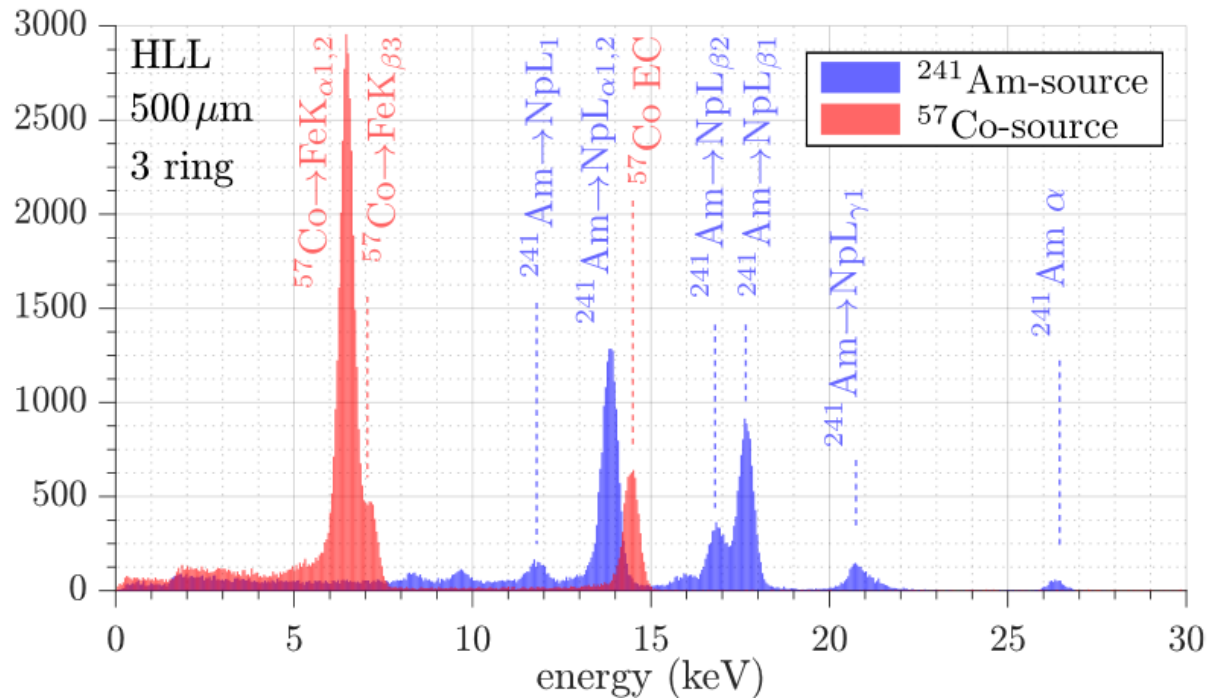


Prototype supported by Research Seed Capital funding of MWK Baden Württemberg

# TRISTAN Prototype

> test setups at KIT & CEA

FWHM of 400 eV @ 6 keV



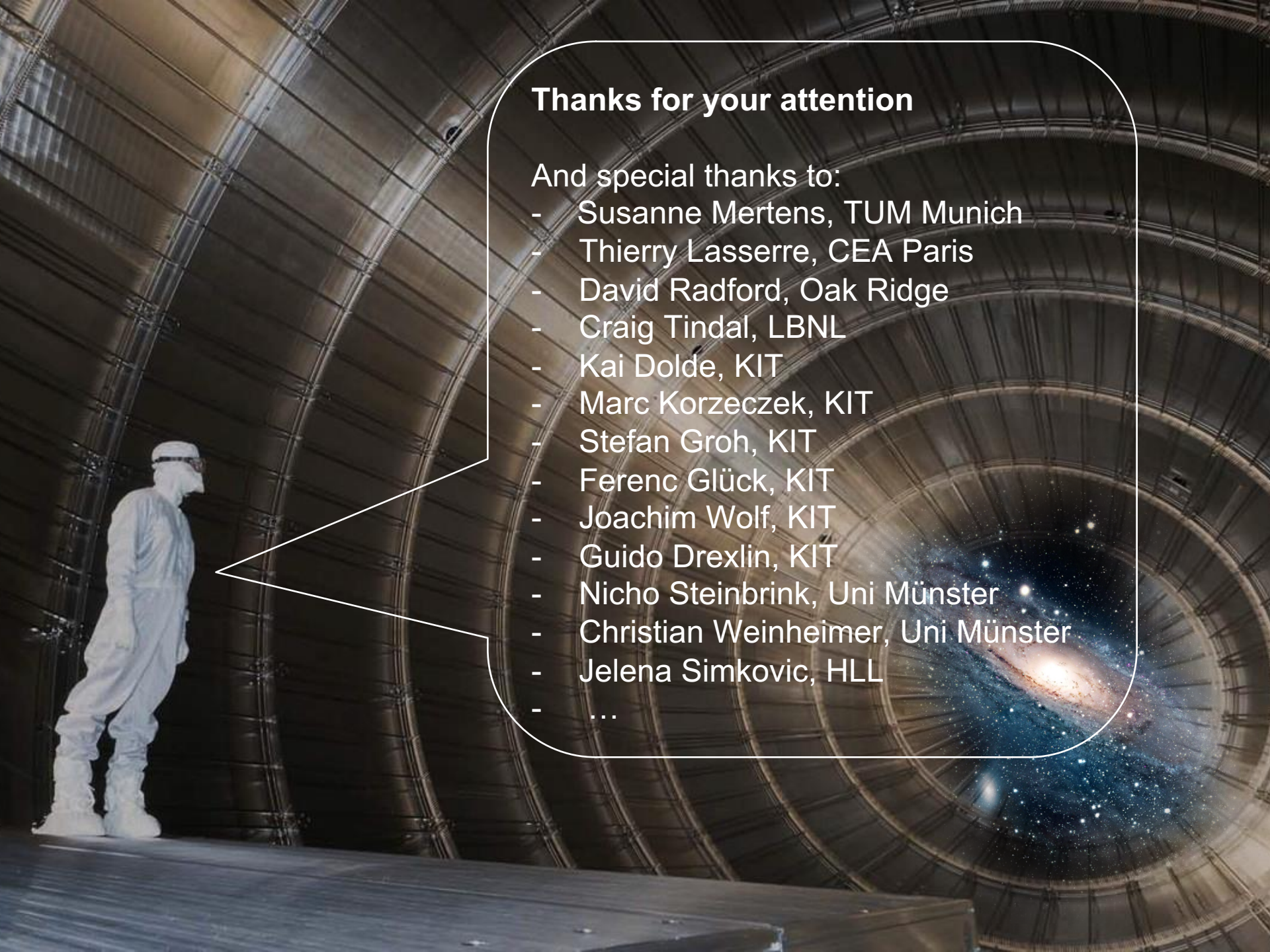
„First Light“ of TRISTAN Prototype measured at CEA

- The KATRIN Experiment Status
  - All components on-site
  - Alignment and final commissioning at the moment
  - Detailed calibration measurements scheduled for end of this year
  
- The KATRIN Experiment and sterile neutrinos
  - with the KATRIN source: statistical sensitivity  $\sin^2 \theta < 10^{-8}$
  - Study of systematic uncertainties is ongoing (goal is to optimize the experiment to reach a sensitivity of  $\sin^2 \theta < 10^{-6}$ )
  - Two measurement ideas: a Pre- and a Post-measurement
  - Pre-Measurement will take place mid 2017 (preliminary:  $\sin^2 \theta < 10^{-4}$ )
  - A new detector system (TRISTAN) is currently developed at KIT, CEA and Munich

## Thanks for your attention

And special thanks to:

- Susanne Mertens, TUM Munich
- Thierry Lasserre, CEA Paris
- David Radford, Oak Ridge
- Craig Tindal, LBNL
- Kai Dolde, KIT
- Marc Korzeczek, KIT
- Stefan Groh, KIT
- Ferenc Glück, KIT
- Joachim Wolf, KIT
- Guido Drexlin, KIT
- Nicho Steinbrink, Uni Münster
- Christian Weinheimer, Uni Münster
- Jelena Simkovic, HLL
- ...





# The entire KATRIN Collaboration

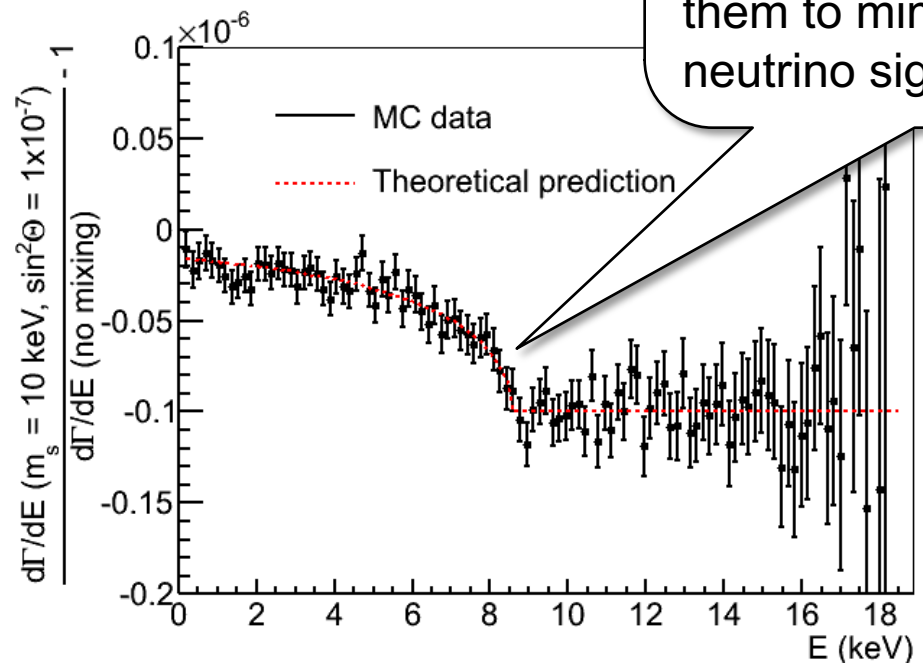


# Back up Slides

# Systematic uncertainties

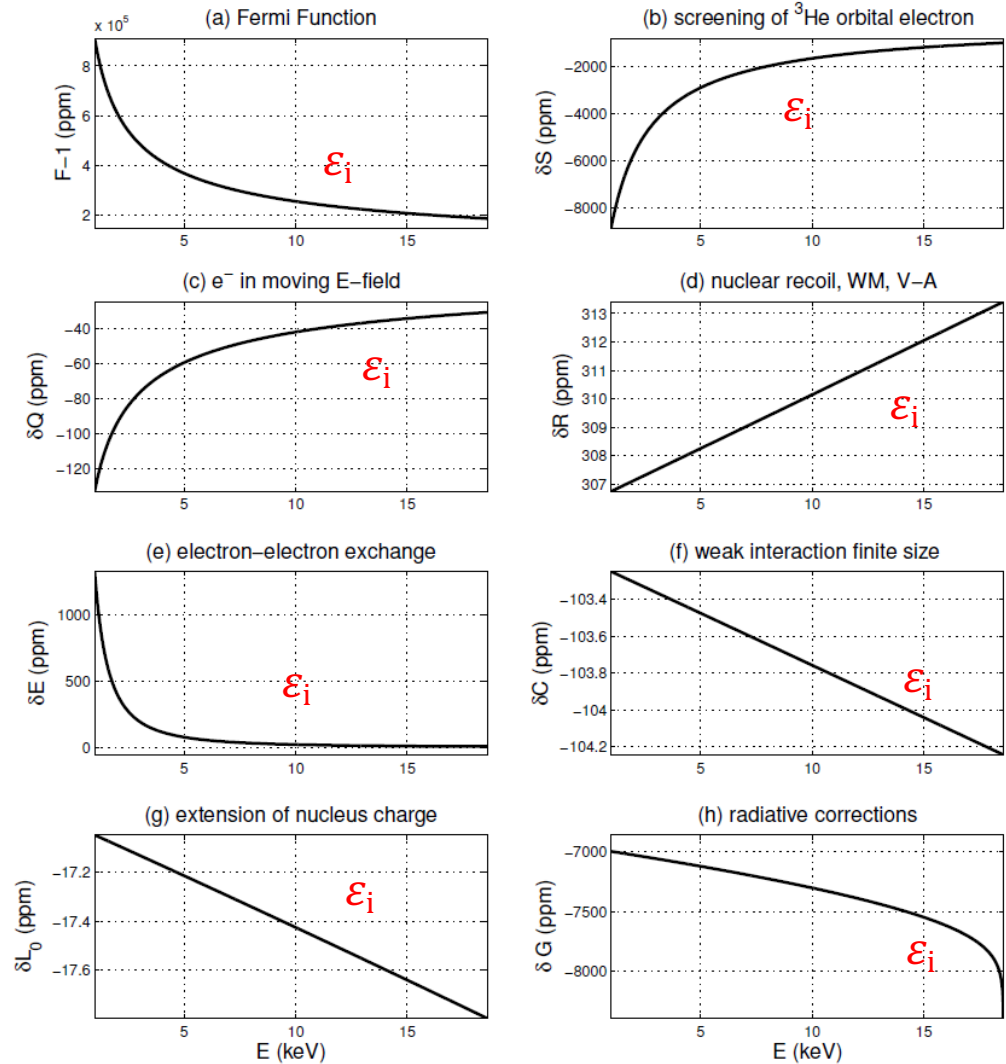
## 1. Knowledge of the theoretical spectrum?

Idea:  
 Parametrize uncertainties  
 and let these parameters  
 free in the fit, to allow  
 them to mimic a keV  
 neutrino signature

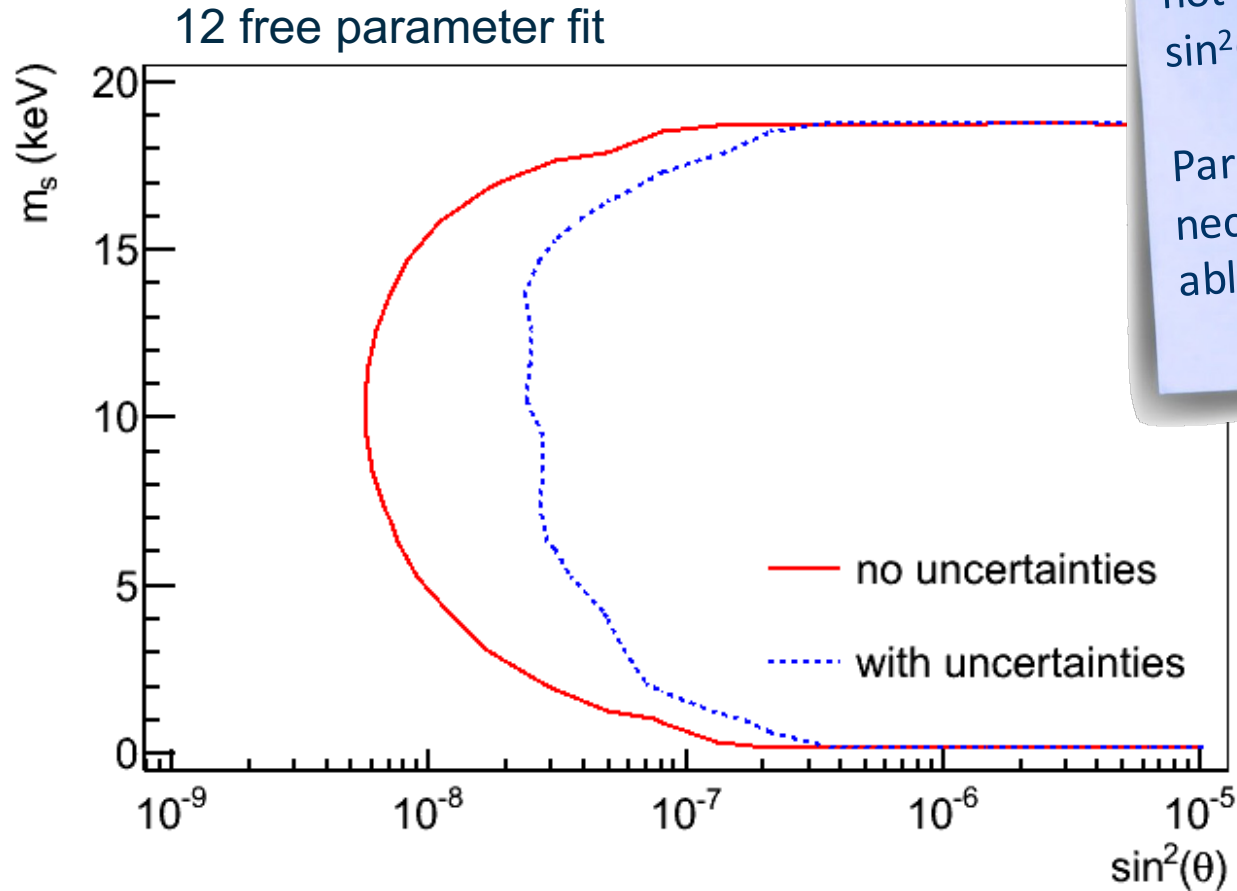


# Theoretical corrections to the $\beta$ -spectrum

Non-negligible  
BUT  
smooth  
in energy



# Spectral Fit Approach



Smooth corrections do not fake a kink signal  
 $\sin^2(\theta) > 10^{-7}$

Parametrization is necessary in order to be able to perform a fit

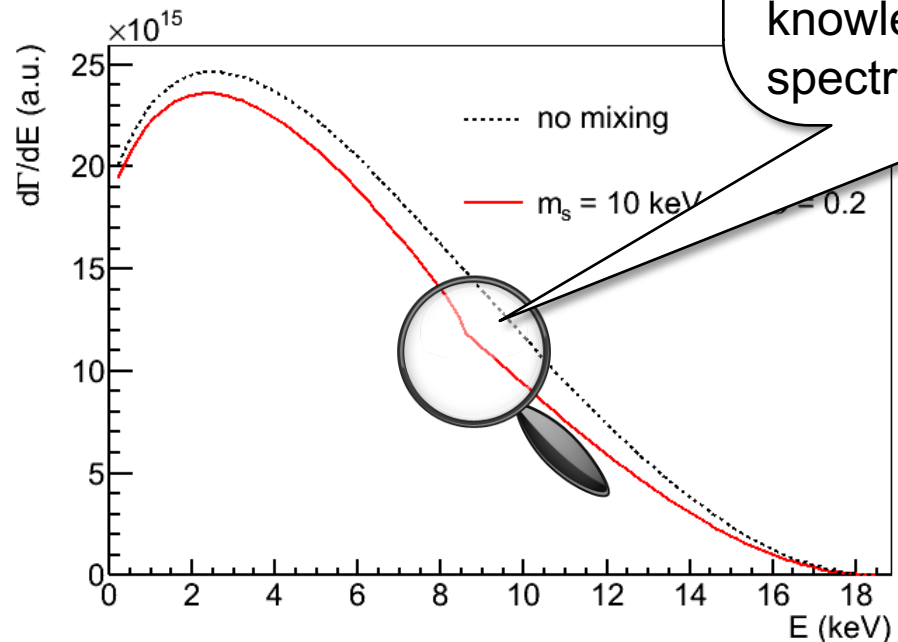
KATRIN source strength, 3-years differential measurement, 90%CL

S. M. *et al.* JCAP 1502 (2015) 02, 020,  
arXiv:1409.0920

# Systematic uncertainties

1. Knowledge of the theoretical spectrum?
2. Is this knowledge necessary at all?

Idea:  
 Use wavelet transformation to detect „kink“ feature in the spectrum, in order to be insensitive to the exact knowledge of the true spectrum

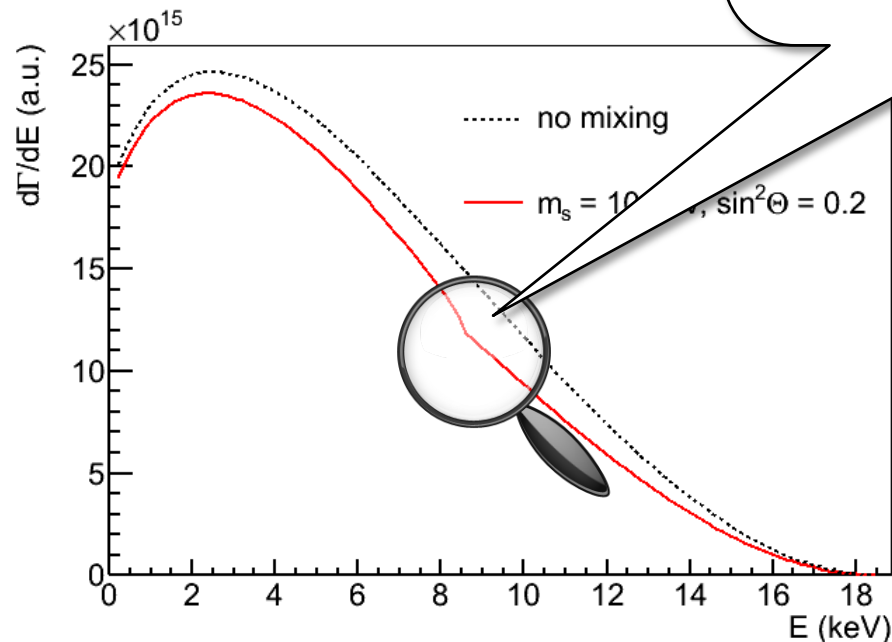


# Systematic uncertainties

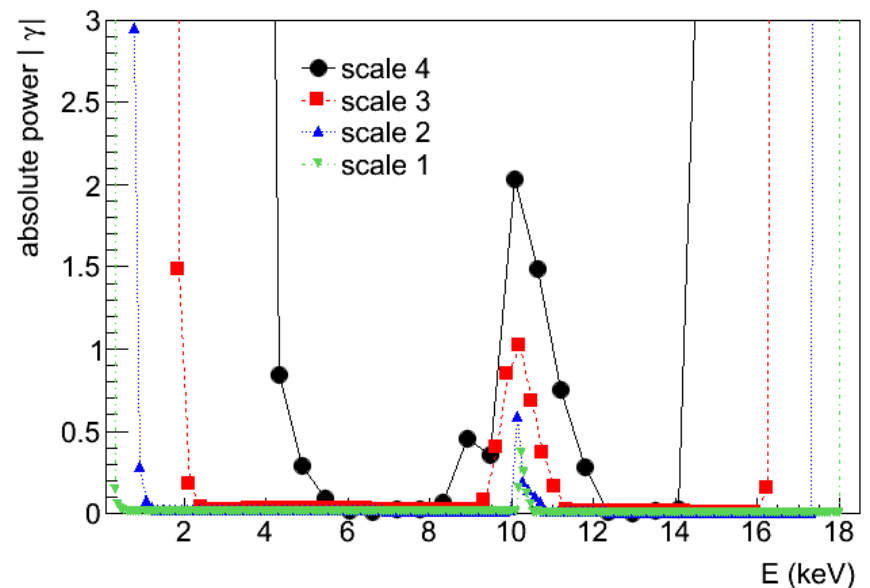
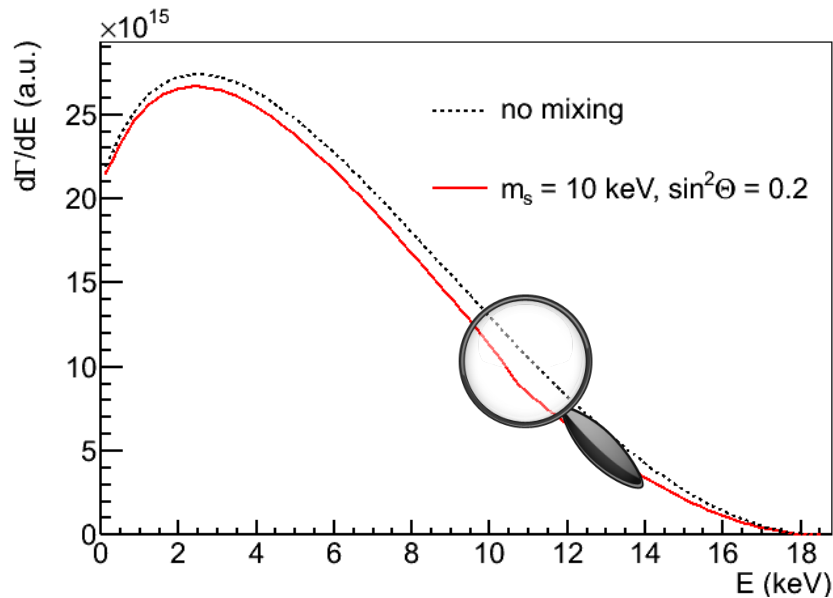
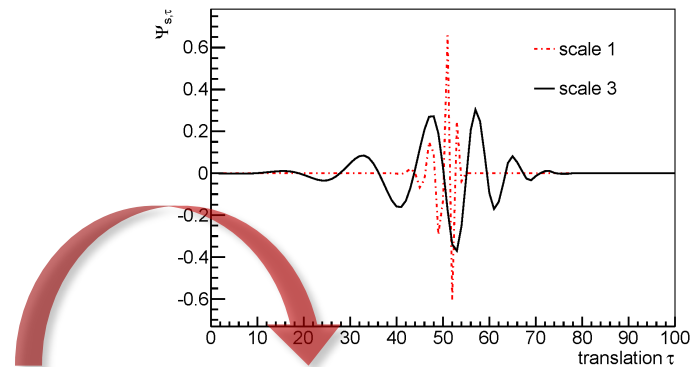
1. Knowledge of the theoretical spectrum?
2. Is this knowledge necessary at all?

Wavelet Transformation:

Which frequency is present in the signal at which energy?



# Wavelet Approach



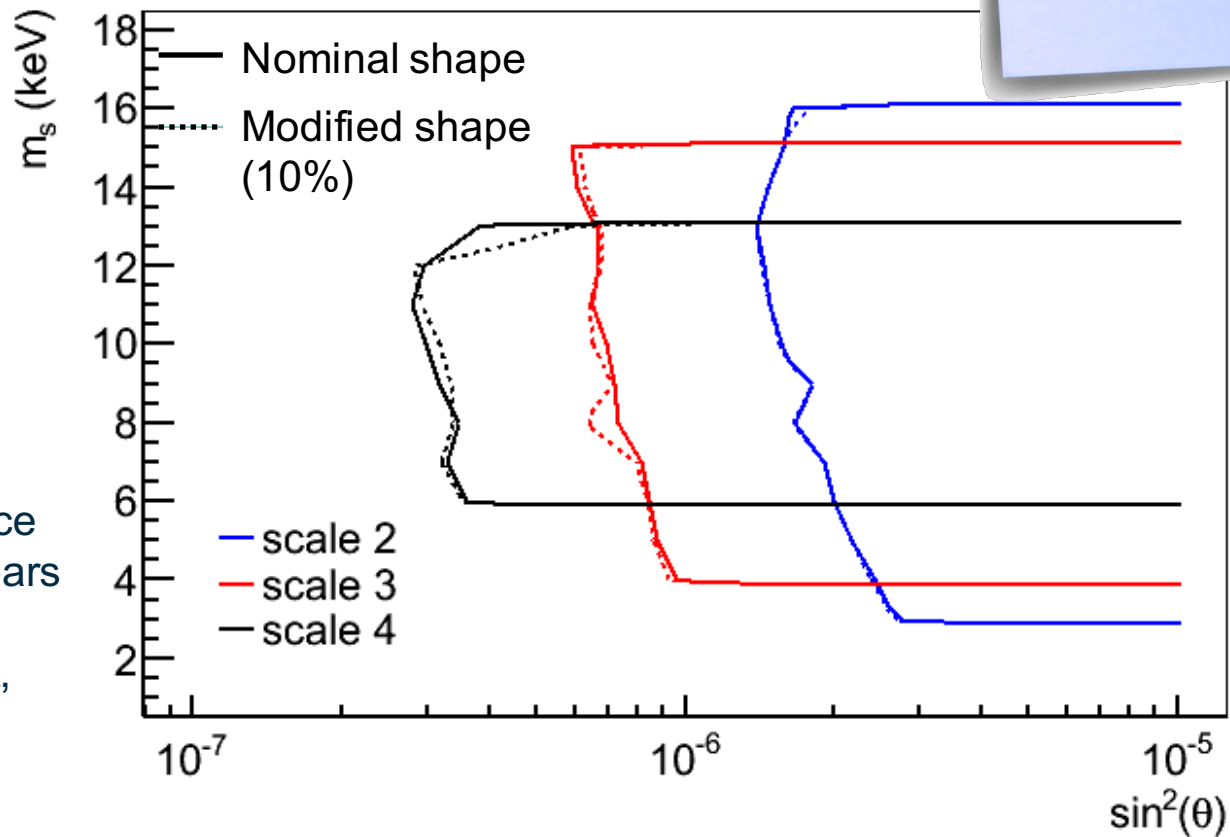
S. Mertens et. al.  
 arXiv:1410.7684



# Wavelet Approach

Wavelet approach is independent of exact shape of tritium spectrum

KATRIN source strength, 3-years differential measurement, 90%CL

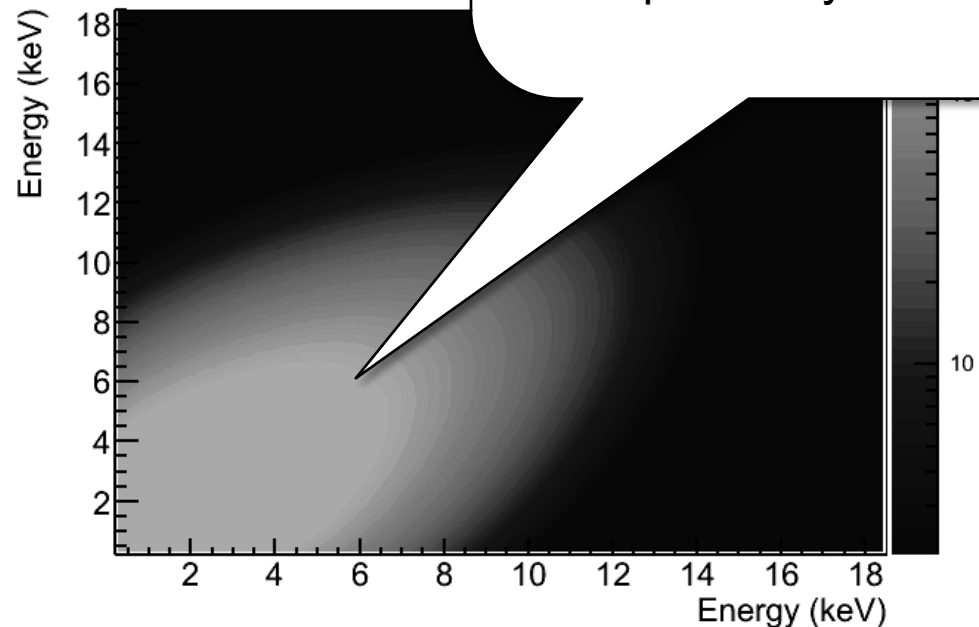


# Systematic uncertainties

1. Knowledge of the theoretical spectrum?
2. Is this knowledge necessary at all?
3. How to deal with Systematics?

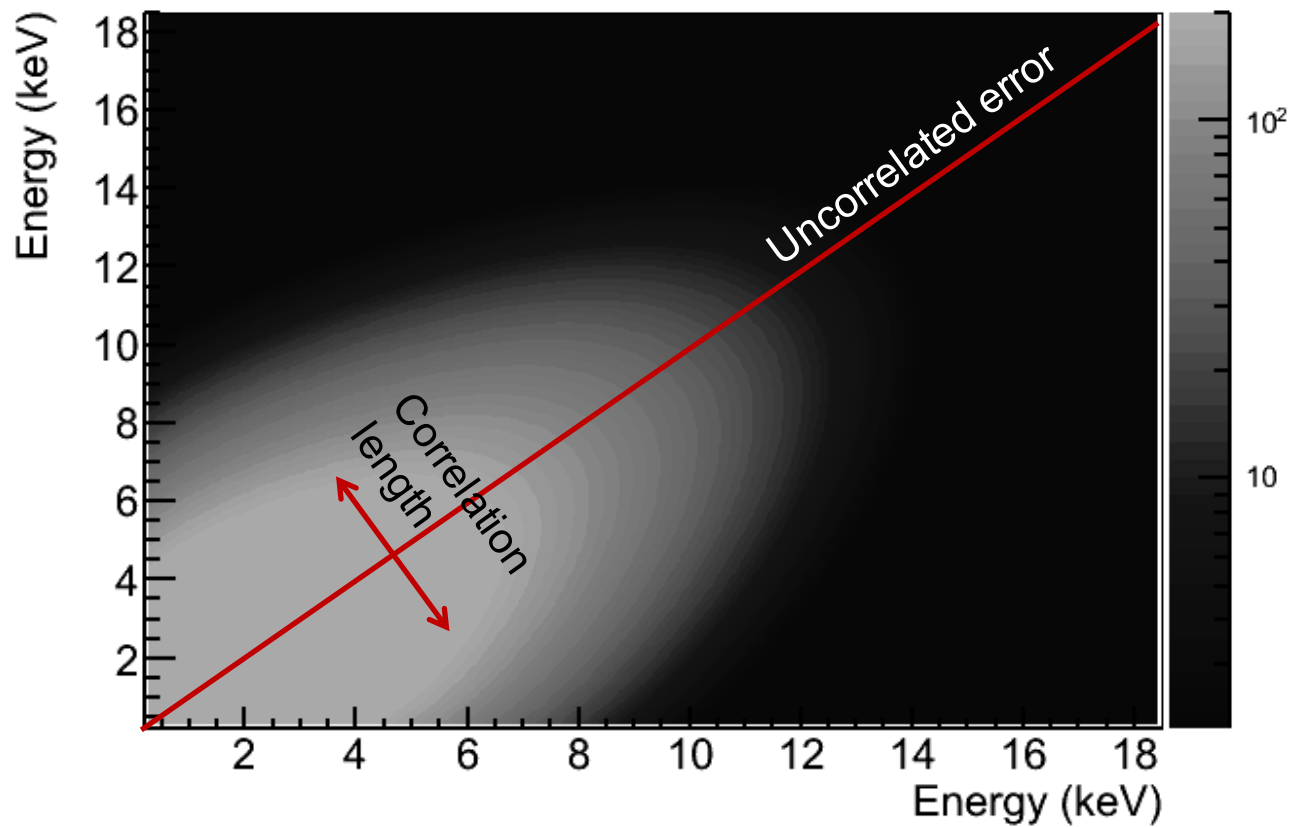
Idea:

Construct realistic covariance matrix to investigate experimental uncertainties in a conceptual way.

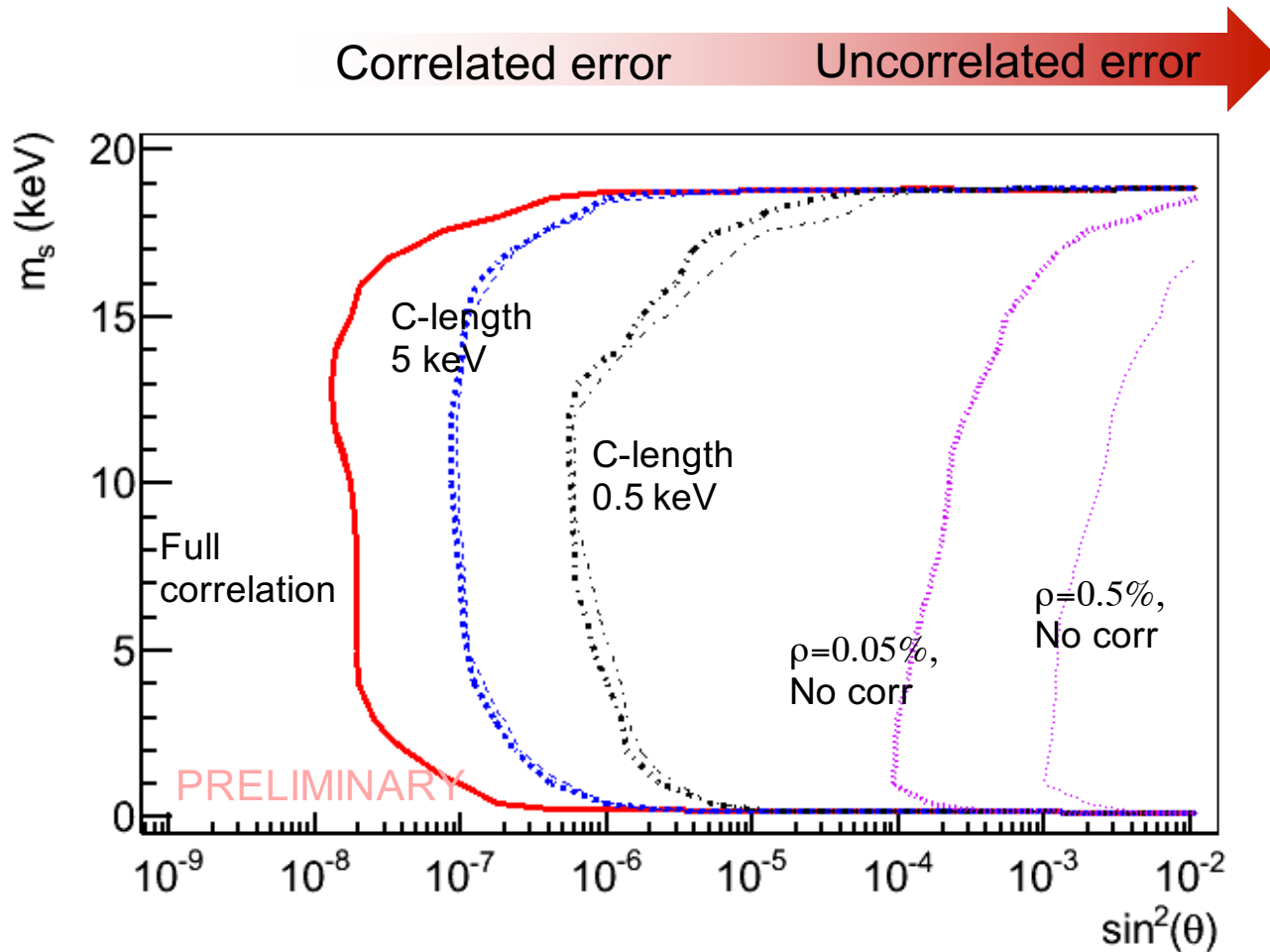


# Experimental uncertainties

## Covariance matrix



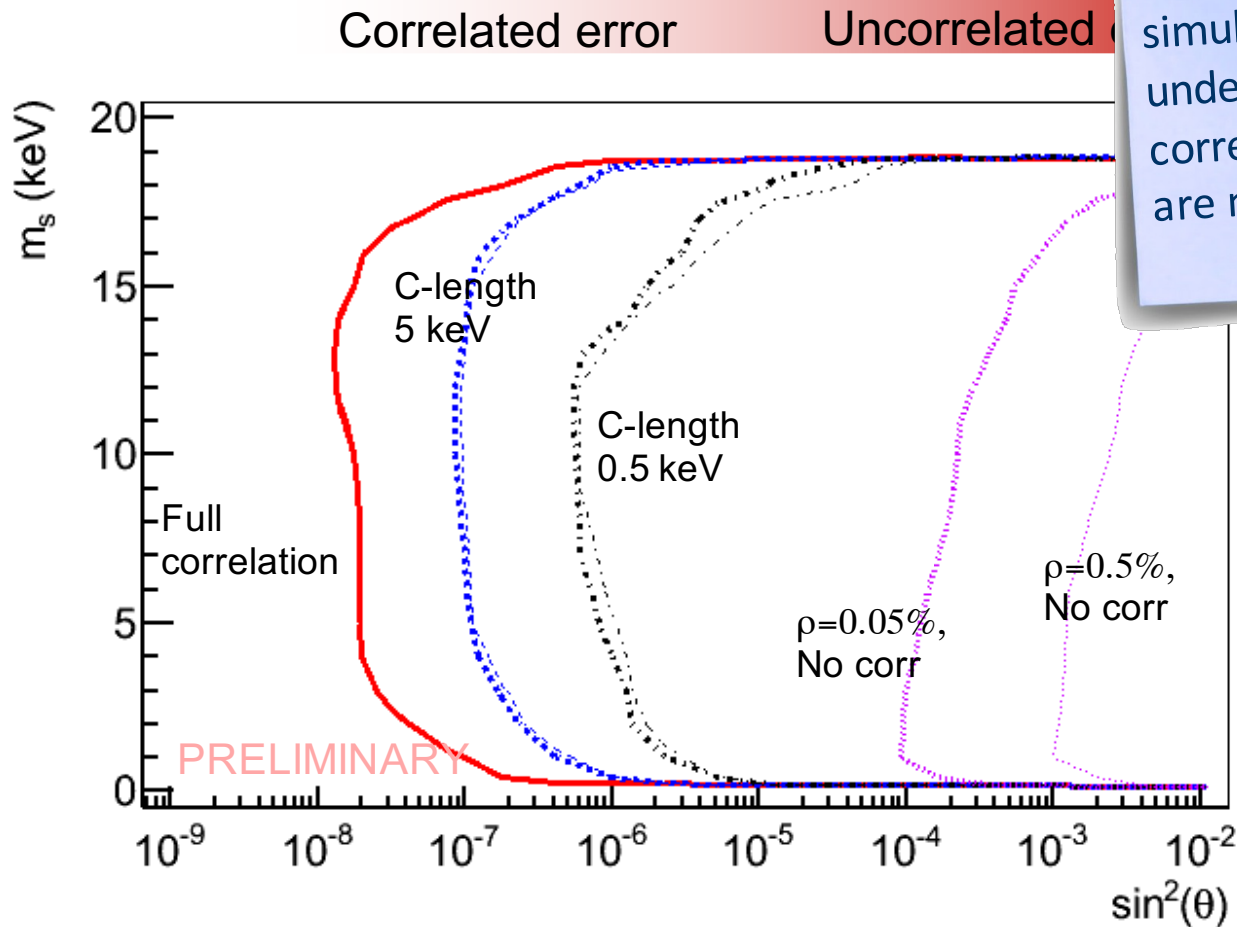
# Experimental uncertainties



KATRIN source strength, 3 years differential measurement, 90% CL

# Experimental uncertainties

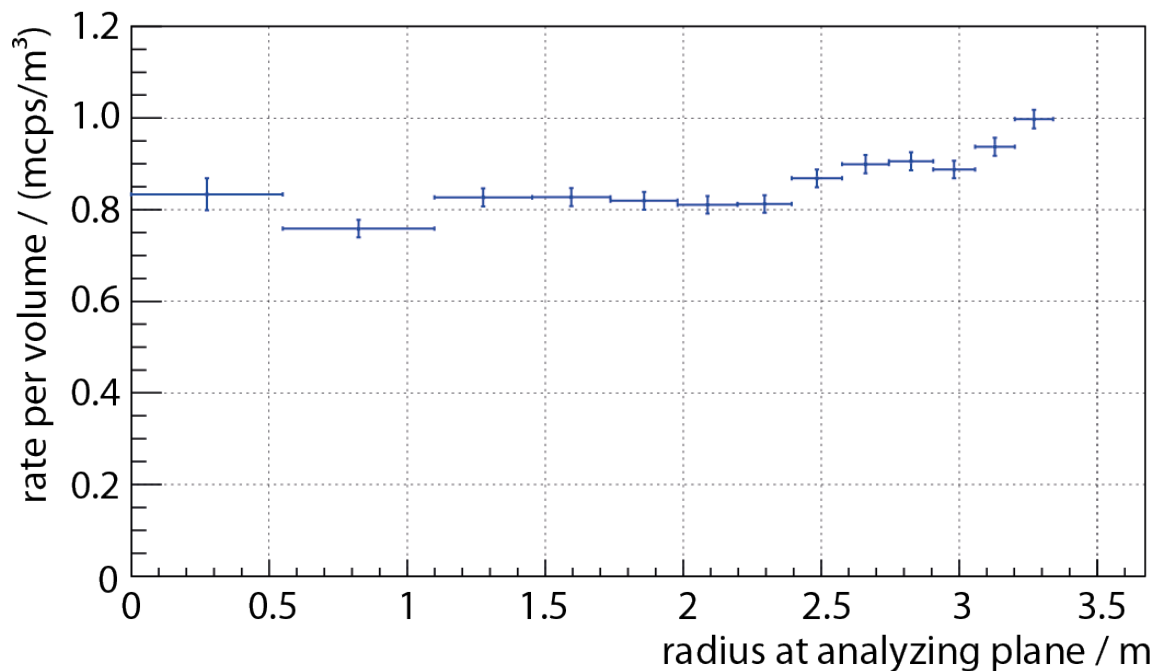
Precise calibrations, simulations, and understanding of correlations are needed



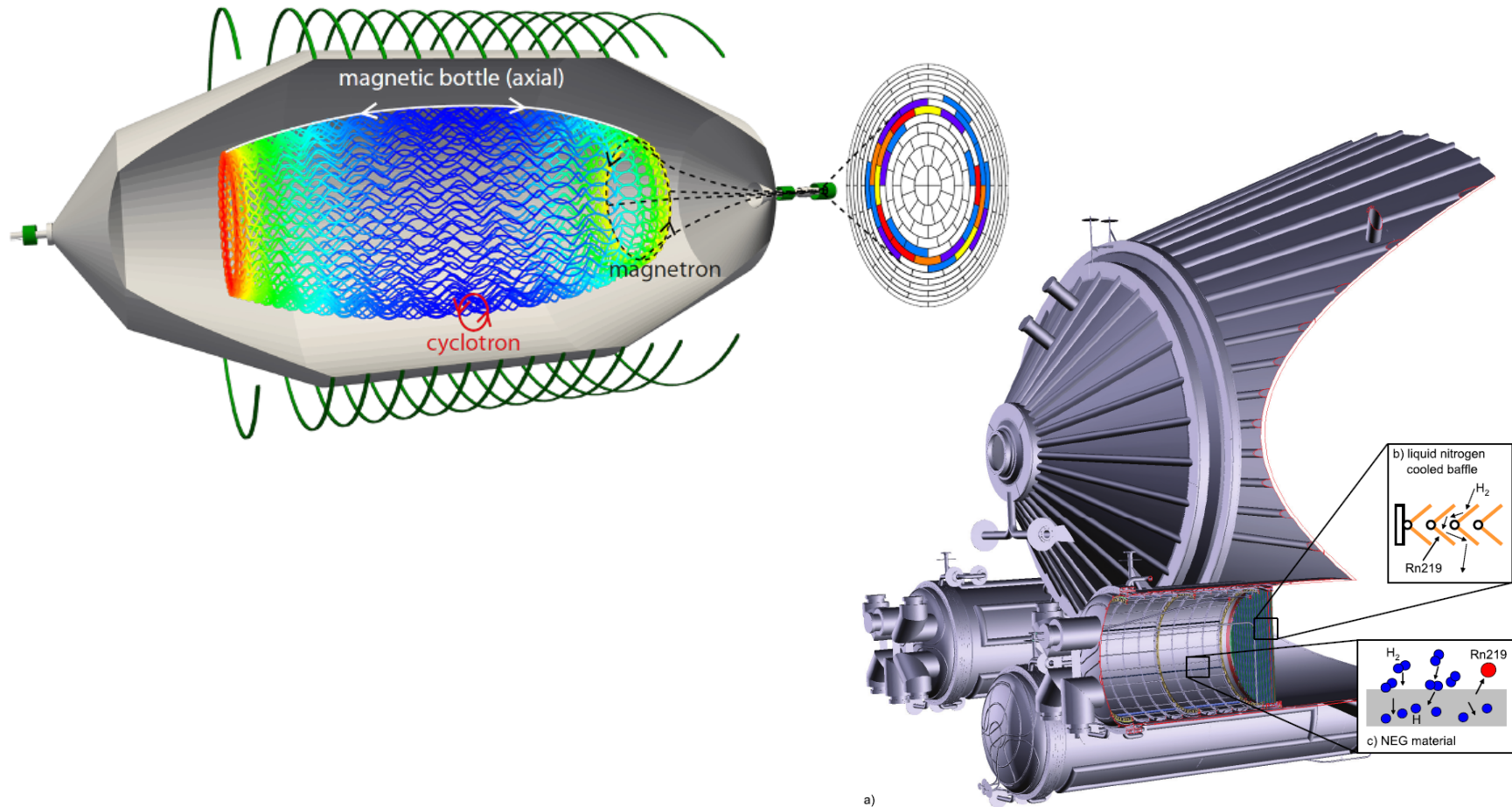
KATRIN source strength, 3 years differential measurement, 90% CL

# KATRIN Background

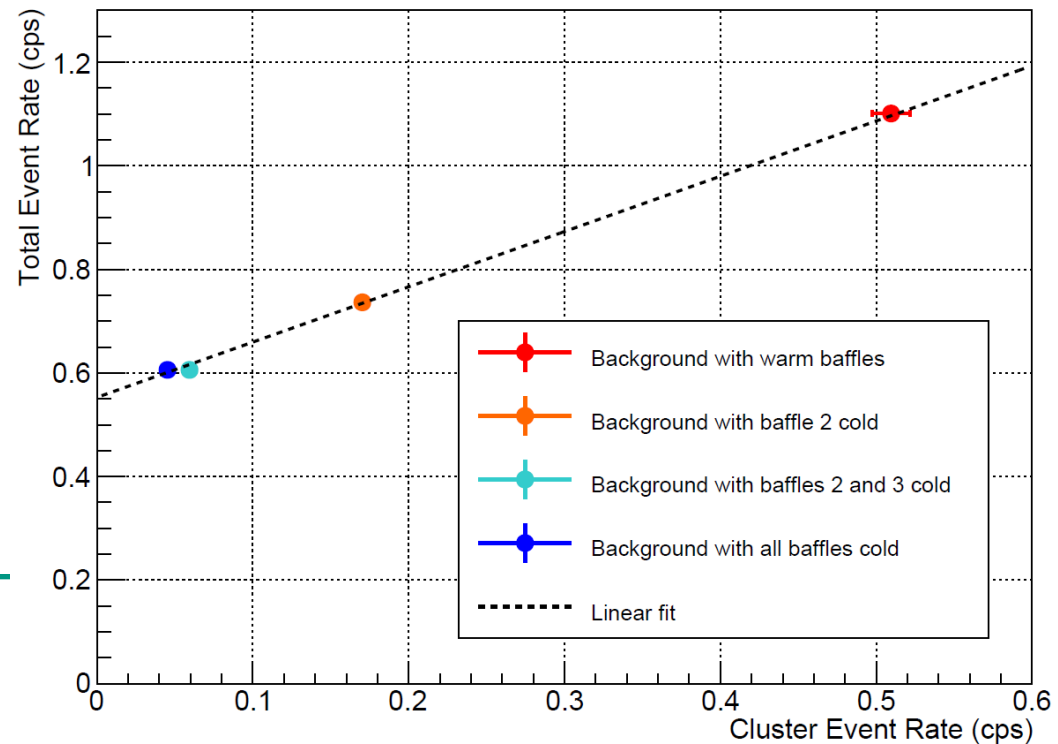
- Background rate in ROI **477 +/- 3 mcps** (10 mcps required)
- Settings: vessel = -18.5kV, IE = -100V, PAE = +10 kV and “5G” magnetic field setting



# Radon induced background



# Effect of cold baffle on Radon background



$$B_{\text{total}} = S_{\text{Rn}} + C_{\text{Rn}} + R$$

$$S_{\text{Rn}} = \alpha \cdot C_{\text{Rn}}$$

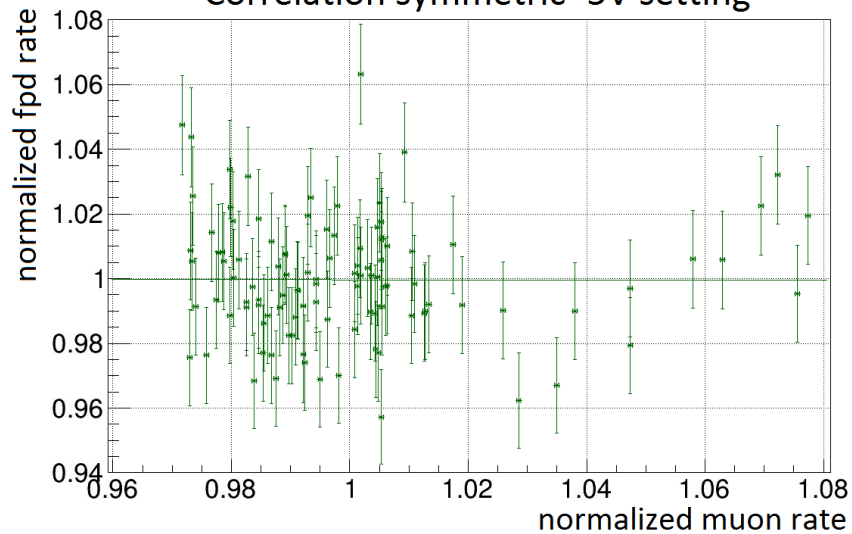
$$B_{\text{total}} = (\alpha + 1) \cdot C_{\text{Rn}} + R$$

$B_{\text{total}}$  : Total background rate.  
 $S_{\text{Rn}}$  : Radon-induced single event rate.  
 $C_{\text{Rn}}$  : Event rate in Radon-induced clusters.  
 $R$  : Non-Radon-induced background rate.



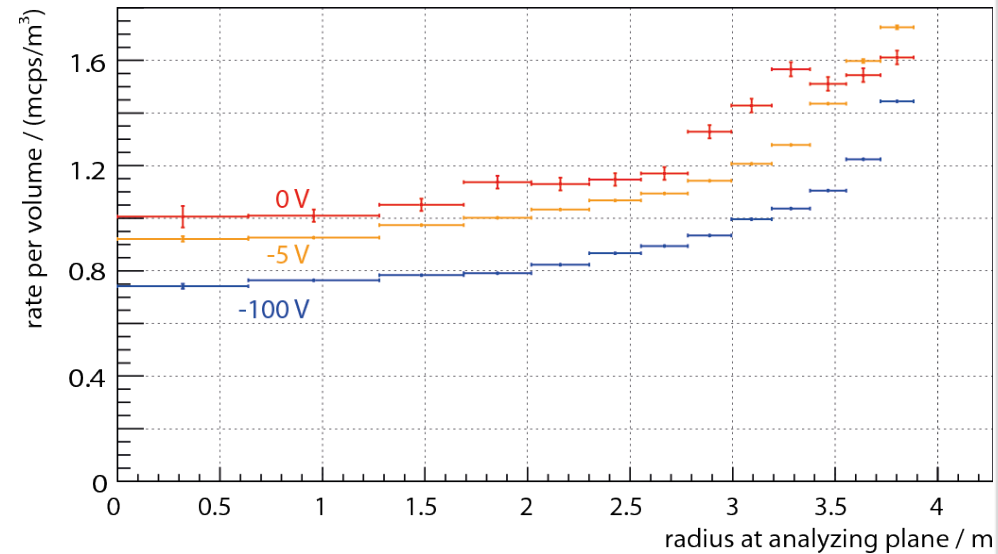
# Cosmic induced backgrounds

Correlation symmetric -5V setting



slope  $\alpha$ :  $-0.001 \pm 0.064$

correlation factor:  $-0.01 \pm 0.1$



# KATRIN Spectrometer Status

Beginning of 2015 measurement phase completed

- Spectrometer works as MAC-E Filter
- Liquid nitrogen cooled baffles eliminate Radon-induced background with an efficiency of  $\varepsilon = (97 \pm 2)\%$
- Remaining background is still under investigation
- Excellent HV stability

