

# **(sub)eV Sterile Neutrinos: an experimental overview**

**Chalonge Meudon Workshop 2016**  
**Meudon, June 15<sup>th</sup> 2016**

**Matthieu Vivier**  
**CEA**

# Established Neutrino Physics

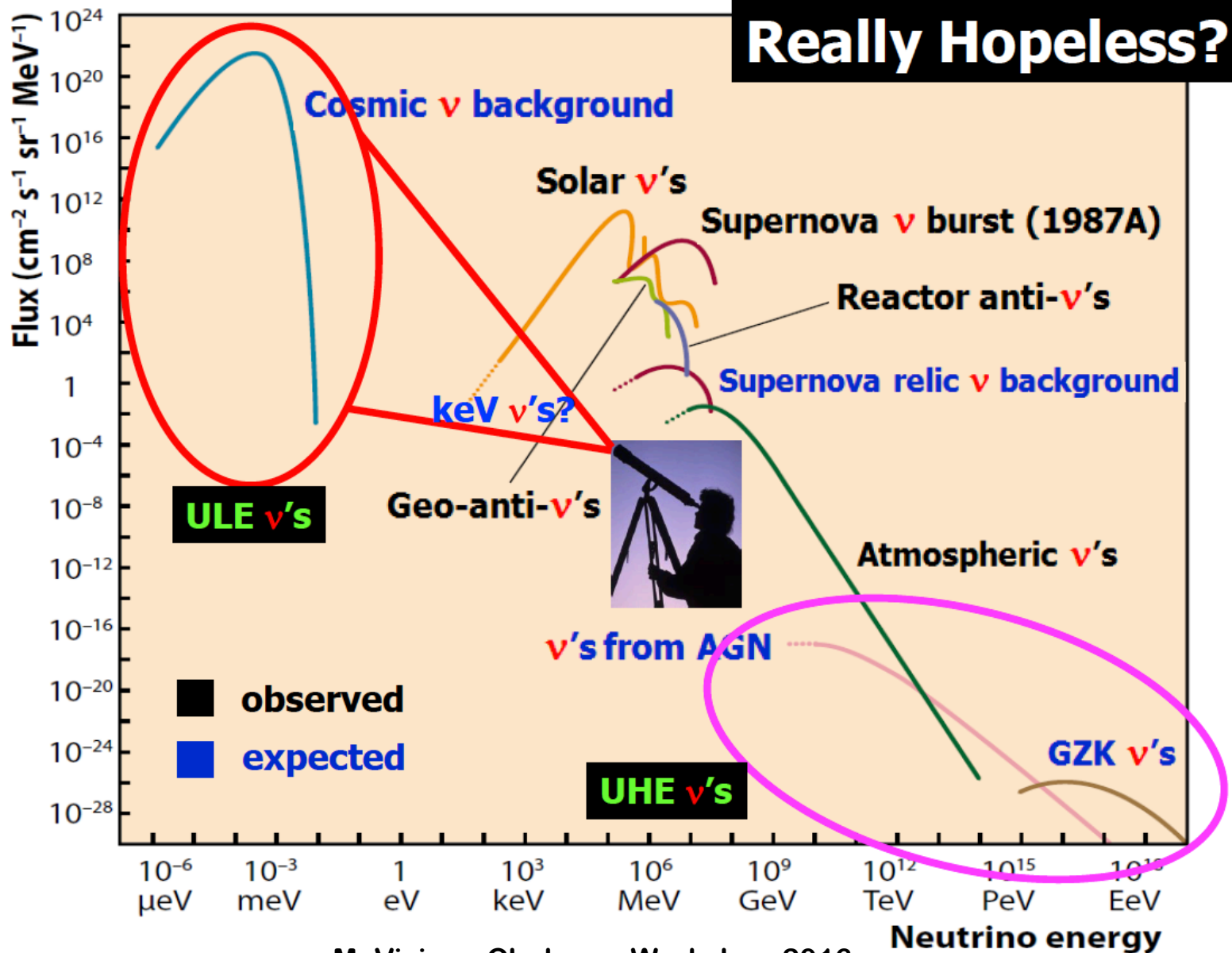
- 3 types, spin  $\frac{1}{2}$ , neutral, left handed,  $\sigma(1 \text{ MeV}) \approx 10^{-45-43} \text{ cm}^2$
- Neutrinos have tiny masses and mix:  $0.04 \text{ eV} < m_\nu < \approx 1 \text{ eV}$
- PMNS matrix  $U$  relates mass & flavor:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{\text{PMNS}} \times \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \Leftrightarrow |\nu_e(t)\rangle = \sum_i U_{ei} |\nu_i(t)\rangle$$

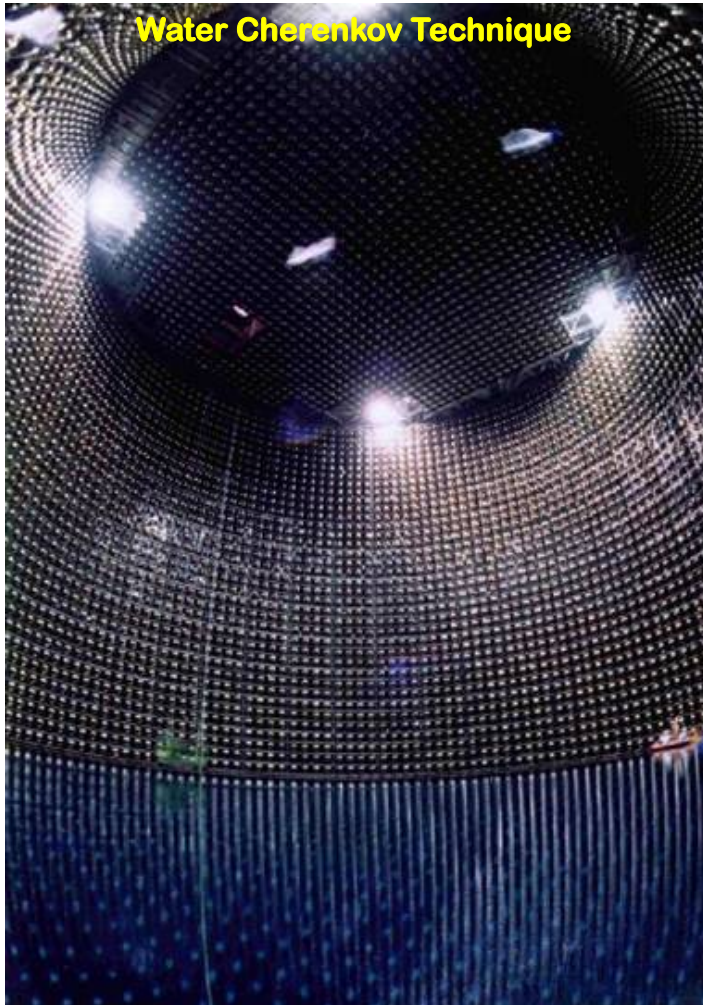
- A compelling evidence of physics Beyond the Standard Model



# Neutrinos in the Universe

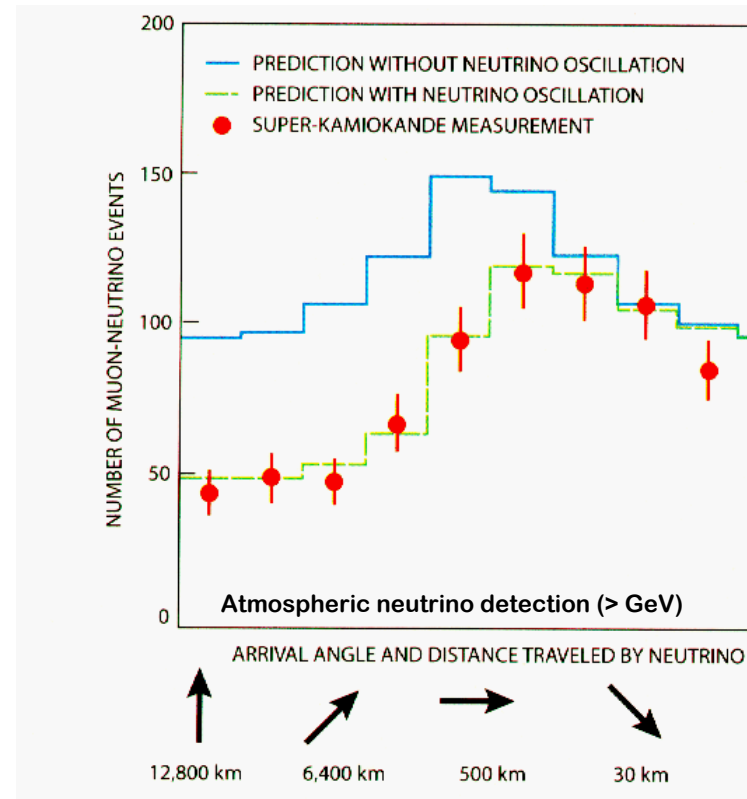


# SuperKamiokande Breakthrough (1998)



50 kt of pure water, 12 000 PMTs  
 Good E-resolution  
 e/μ discrimination at low energy

$$\frac{\Phi^{\text{Atm}}_{\nu_{\mu}}(\text{up})}{\Phi^{\text{Atm}}_{\nu_{\mu}}(\text{down})} = 0.54 \pm 0.04$$



**Neutrino do have mass  
 and they mix (oscillation)**

# Neutrino Oscillation: Established

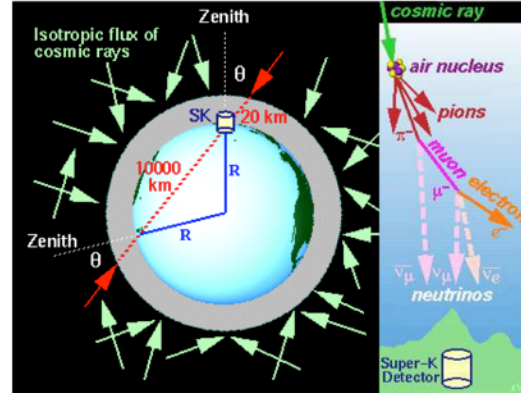
sun



reactors



atmosphere



accelerators



Homestake, SAGE, GALLEX  
SuperK, SNO, Borexino

KamLAND, CHOOZ

SuperKamiokande

K2K, MINOS, T2K

- $\nu_\mu \rightarrow \nu_{\tau,e}$  or  $\text{anti-}\nu_\mu \rightarrow \text{anti-}\nu_{\tau,e}$  :
- $\nu_e \rightarrow \nu_{\mu,\tau}$  :
- $\text{anti-}\nu_e \rightarrow \text{anti-}\nu_{\mu,\tau}$  :
- $\nu_\mu \rightarrow \nu_e$  :

- atmospheric & beam experiments
- solar experiments
- reactor experiments
- beam experiments

$$P(\bar{\nu}_x \rightarrow \bar{\nu}_x) = 1 - \sin^2(2\theta_i) \sin\left(1.27 \frac{\Delta m_i^2 \text{ (eV}^2\text{)} L \text{ (m)}}{E \text{ (MeV)}}\right)$$

2-flavor oscillation formalism

# 3 $\nu$ Oscillation Formalism

## PMNS mixing matrix

$$U = \begin{matrix} \text{Atmospheric} & \text{Cross-Mixing} & \text{Solar} & \text{Majorana CP phases} \\ \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} & \times \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} & \times \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} & \times \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix} \end{matrix}$$

$\theta_{23}$  : “atm.” mixing angle       $\theta_{13}$        $\theta_{12}$  : “solar” mixing angle

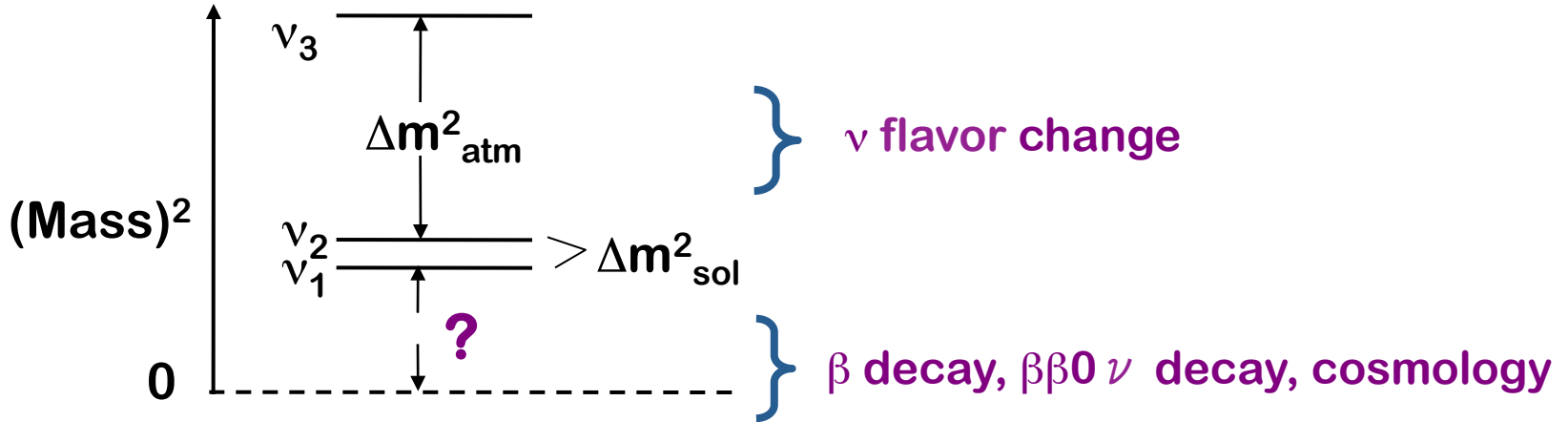
$\delta$  Dirac CP violating phase

2 Majorana phases  
(L violating processes)

- 3 masses  $m_1, m_2, m_3$  :  $\Delta m_{\text{sol}}^2 = m_2^2 - m_1^2$  &  $\Delta m_{\text{atm}}^2 = |m_3^2 - m_1^2|$
- 3-flavour effects are suppressed since:  $\Delta m_{\text{sol}}^2 \ll \Delta m_{\text{atm}}^2$  (1/30) &  $\theta_{13} \ll 1$

# Open questions

- What are the masses of the mass eigenstates  $\nu_i$ ?



- Is the spectral pattern or?  $\bar{\nu}$  behavior in matter,  $\beta\beta 0 \nu$

- Is there any conserved Lepton Number (Dirac or Majorana neutrino)?  $\beta\beta 0 \nu$

- Precise measurements of the leptonic mixing matrix?
  - Do the behavior of  $\nu$  violate CP?
  - Is leptonic  $\bar{CP}$  responsible for the matter-antimatter asymmetry?
- }  $\nu$  flavor change

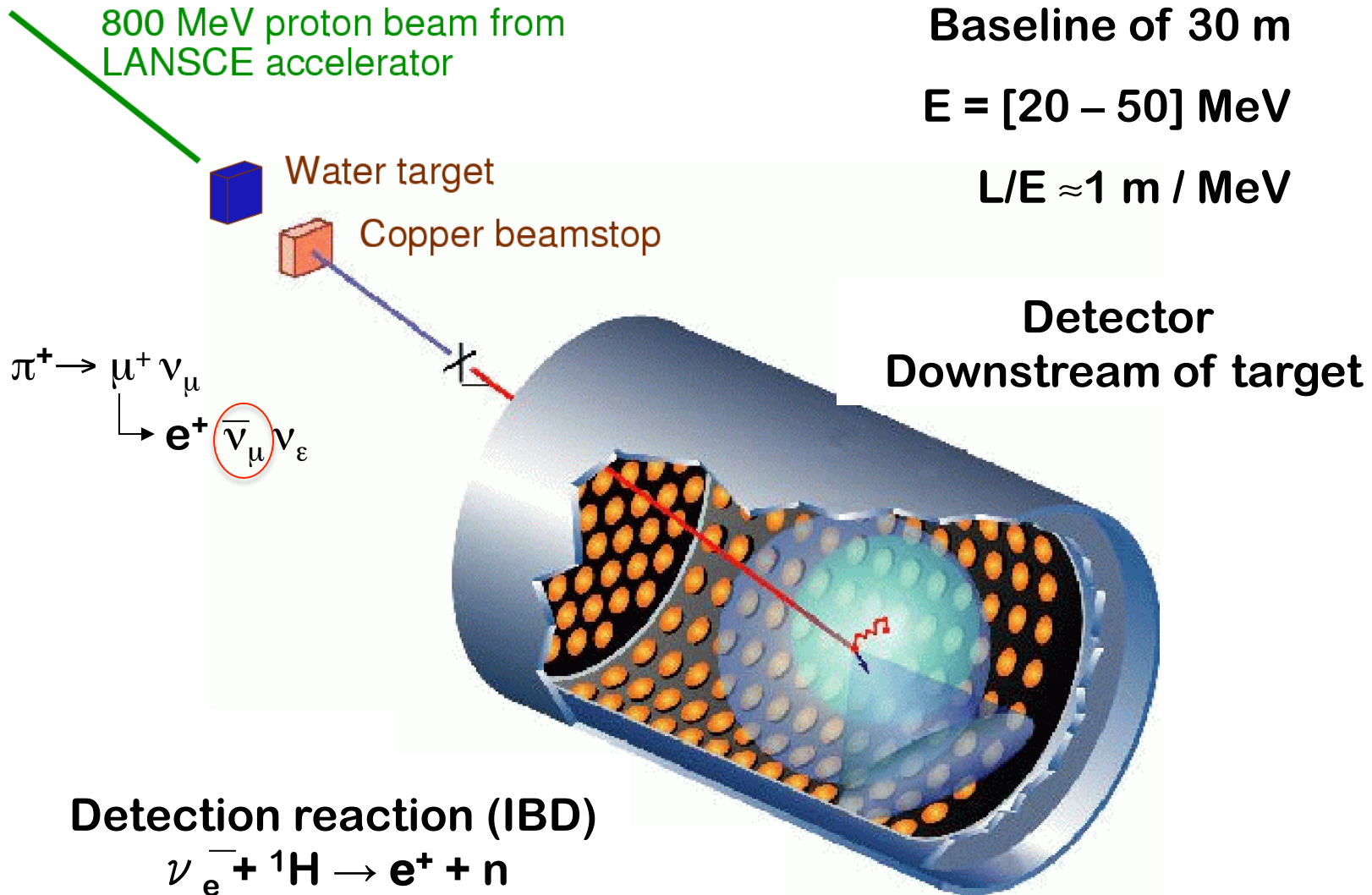
- Are there additional (sterile) neutrino states  $\nu$  flavor change, Cosmology

# Anomalies or new $\nu$ -oscillation?



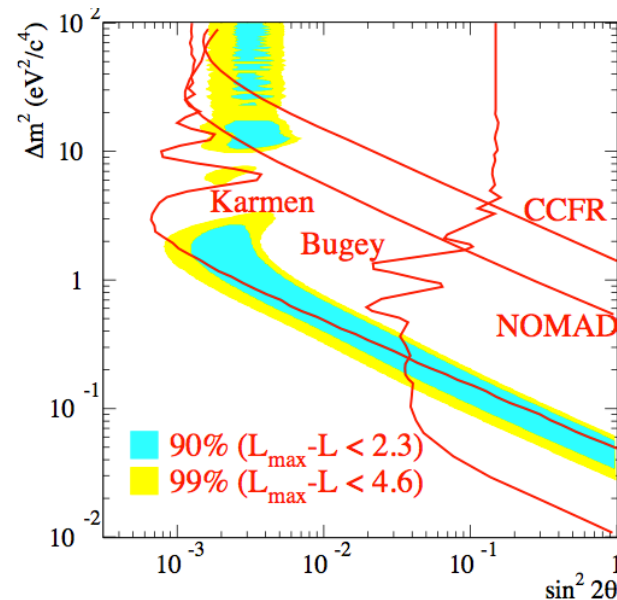
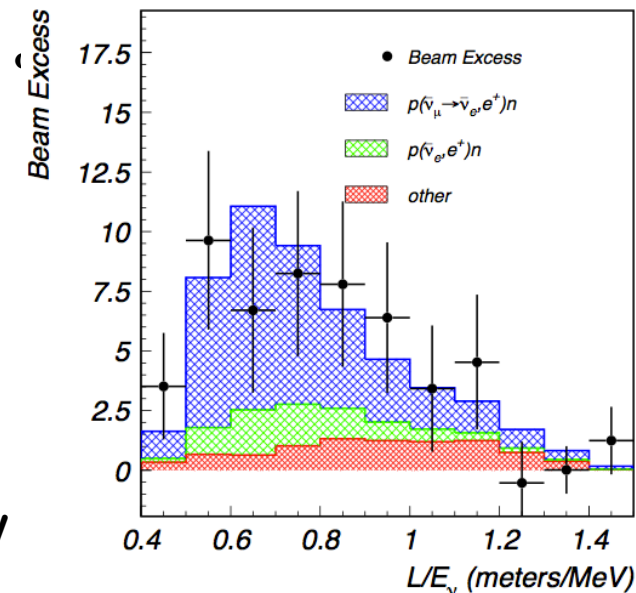
# LSND (stopped $\pi^+$ beam)

Anomaly on the electron antineutrino interaction rate



# LSND Results

- 1<sup>st</sup> results in 1995
  - Channel: anti- $\nu_\mu \rightarrow$  anti- $\nu_e$
  - Detection : anti- $\nu_e + {}^1\text{H} \rightarrow e^+ + n$
  - Baseline: 30 m
  - Energy:  $20 < E \text{ (MeV)} < 50$
- }  $L/E \approx 1\text{m/MeV}$
- Status:
    - anti- $\nu_e$  excess observed  
 $\rightarrow 32.2 \pm 9.4 \pm 2.3 \text{ (3.8}\sigma\text{)}$
    - not confirmed nor ruled out by Karmen
  - $\nu$ -Oscillation interpretation:
    - $\Delta m^2 > 0.1 \text{ eV}^2 \gg \Delta m_{\text{atm}}^2$
    - Require a 4<sup>th</sup> neutrino state

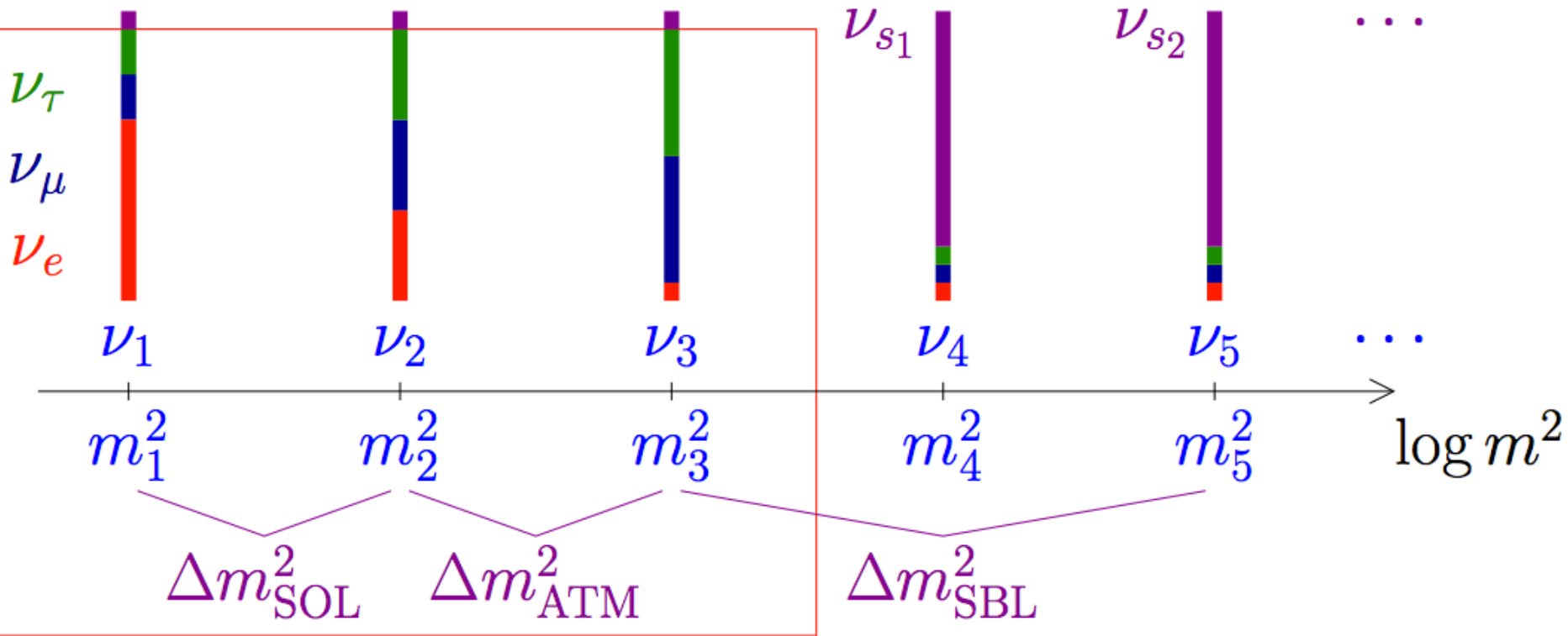




# The (light) sterile neutrino hypothesis

- Generic extension of SM model
- Add a SM singlet fermion
- Mixing with active  $\nu$ 's

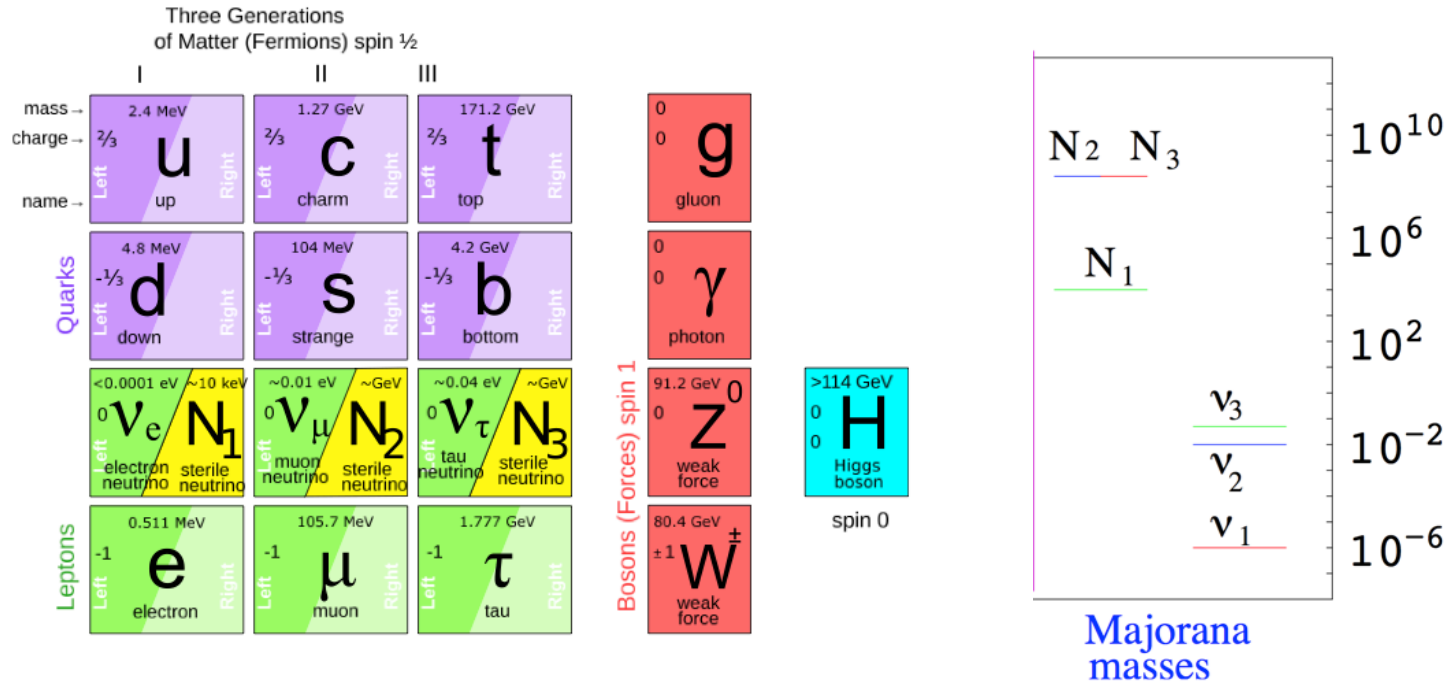
No or tiny SM model interaction



$3\nu$ -mixing

# $\nu$ MSM: a comprehensive model

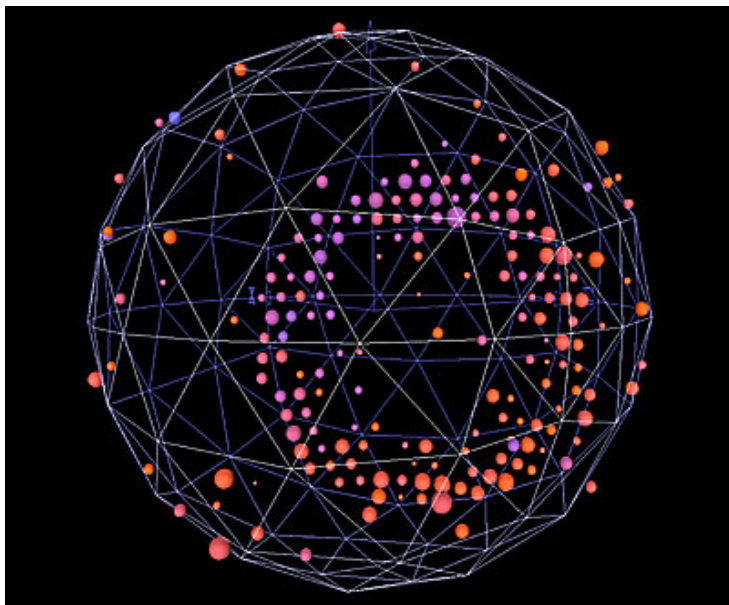
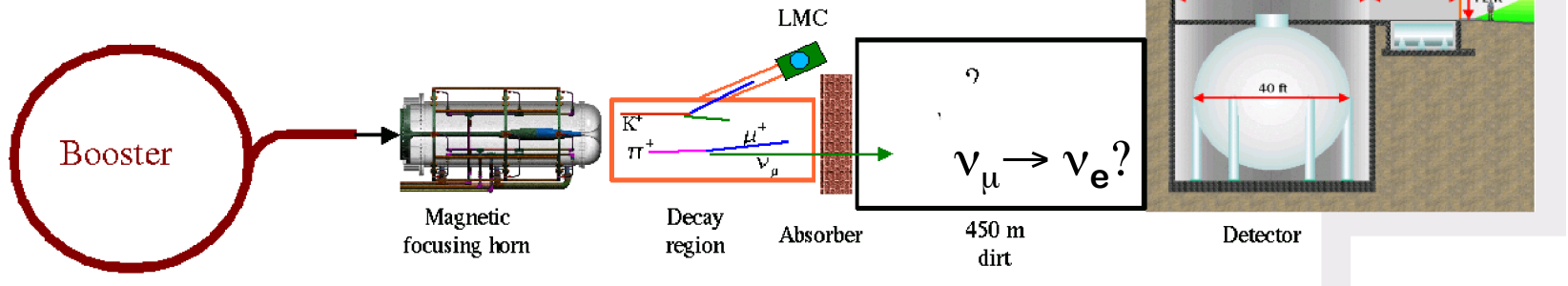
Shaposhnikov, PLB 620, 17 (2005)



- SM + 3 MJ masses, 3 Dirac masses, 6 mixing angles and 6 CP-phases
- Role of  $N_1$  with mass in keV region: **Dark Matter**
- Role of  $N_2, N_3$  with mass in 100 MeV – GeV region: “give” masses to  $\nu$ 's and produce **baryon asymmetry** of the Universe
- No eV scale neutrino...

# MiniBooNE (FNAL)

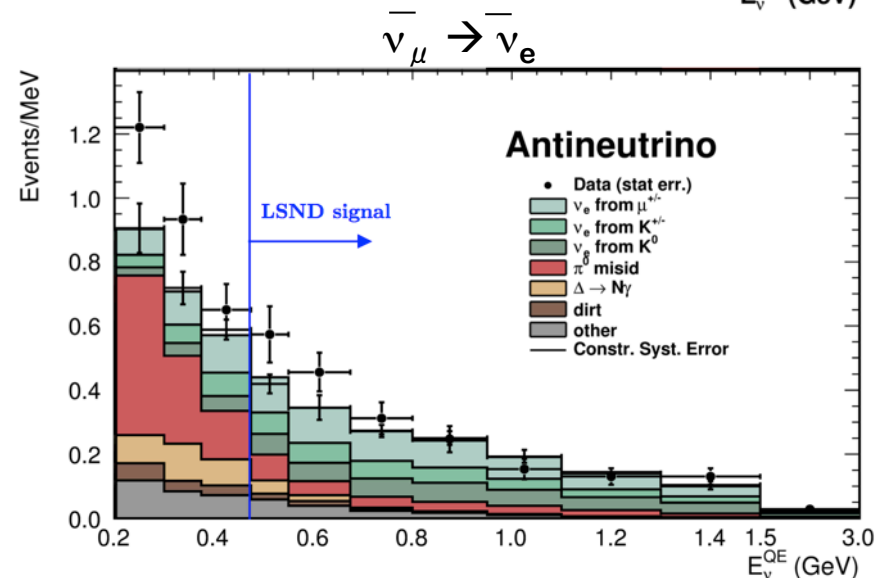
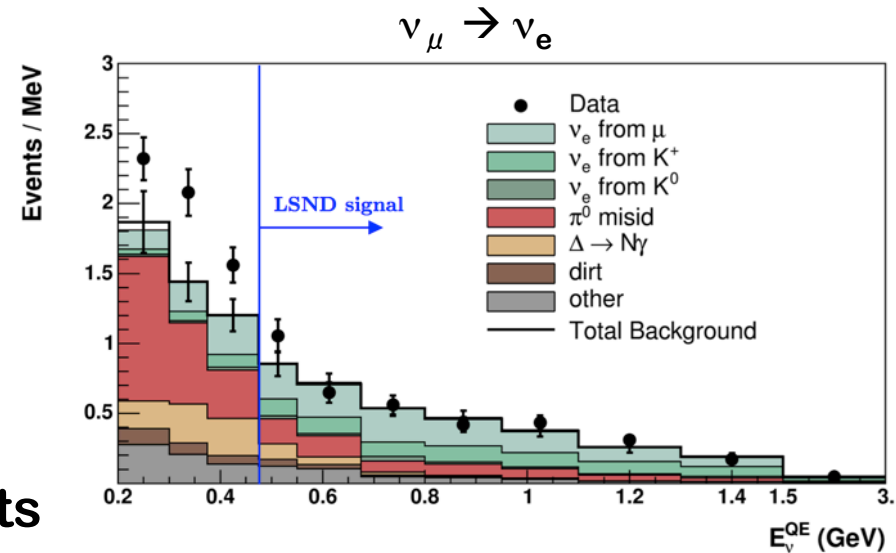
Primary goal: look for  $\nu_e$  appearance in a  $\nu_\mu$  beam  
 Check the LSND with similar L/E



- Beam:  $\pi^+$  ( $\pi^-$ ) decay in flight
- Detection: Cherenkov + scintillation
- L/E  $\approx 1$  m / MeV
  - Baseline: 541 m
  - $200 < E$  (MeV)  $< 3000$
- Statistics:
  - $\nu$  :  $6.46 \times 10^{20}$  POT (2008)
  - $\bar{\nu}$  :  $1.27 \times 10^{20}$  POT (2012)

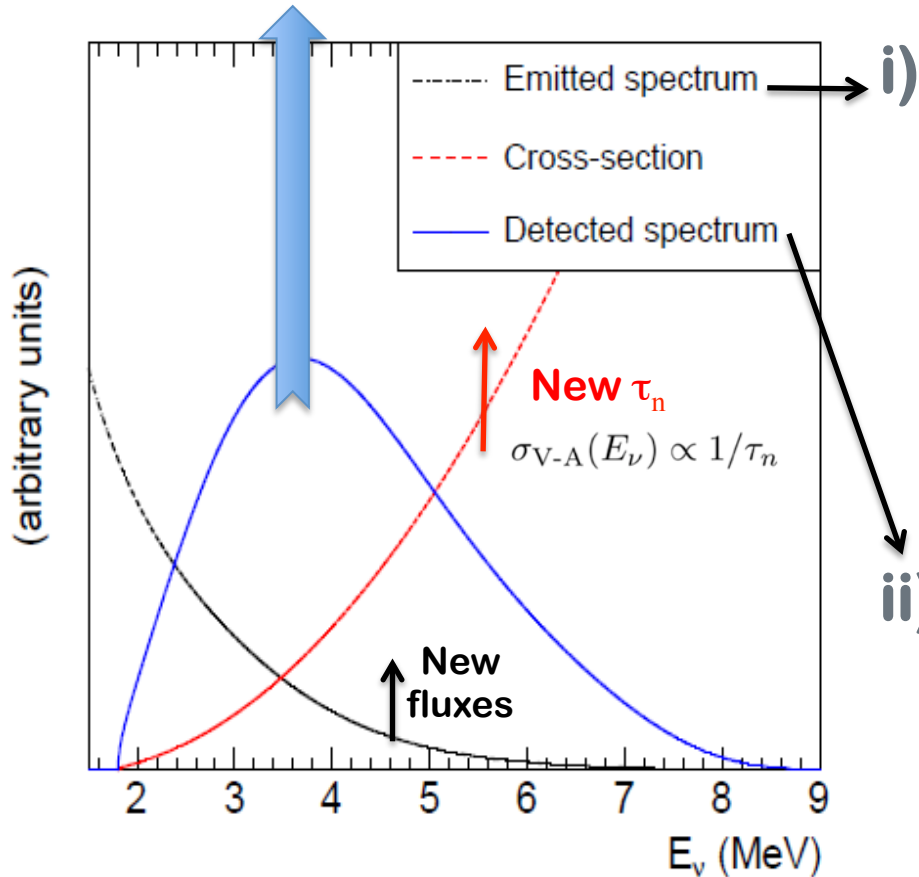
# MiniBooNE Results

- Results published from 2007-12
- **Channel:**  $(\text{anti-})\nu_\mu \rightarrow (\text{anti-})\nu_e$
- **Detection:**  $\nu_e (p)n \rightarrow e p$  (CCQE)
- **Results:**
  - An overall  $3.8 \sigma$  excess of events
  - Mostly at low energy
- **Interpretation:**
  - Backgrounds issue? (to be checked by MicroBooNE)
  - 4<sup>th</sup> neutrino? Or more....
- **MiniBooNE was not conclusive checking the LSND anomaly**



# New Reactor $\nu$ -Fluxes

Increased prediction of  
detected flux by 6.5%



## i) Neutrino Emission:

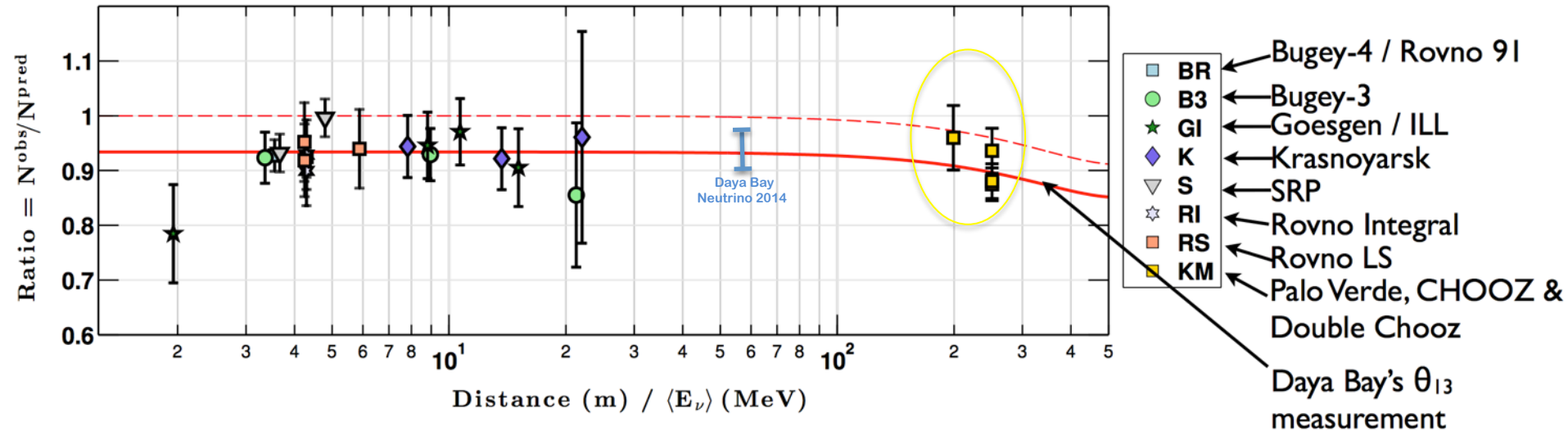
- Improved reactor neutrino spectra  $\rightarrow$  +3.5%
- Accounting for long-lived isotopes in reactors  $\rightarrow$  +1%

## ii) Neutrino Detection:

- Reevaluation of  $\sigma_{IBD} \rightarrow$  +1.5% (evolution of the neutron life time)
- Reanalysis of all SBL experiments

# The Reactor Anomaly

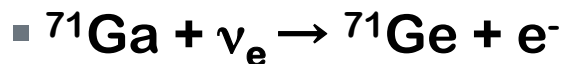
## 2014 Reactor Anomaly Update (new)



- All known nuclear corrections to  $\beta - \nu$  spectra
- Refined treatment of experimental correlations
- Updated neutron mean life ( $\tau_n = 881.5$  s)
- km-scale baselines (Chooz, DC, PV)
  - correcting for  $\theta_{13}$  deficit from Daya Bay's measured value
- **2014 result:  $\mu = 0.938 \pm 0.023$ ,  $2.7\sigma$  deviation from unity**

# The Gallium Neutrino Anomaly

- **Test of solar neutrino radiochemical detectors GALLEX and SAGE**

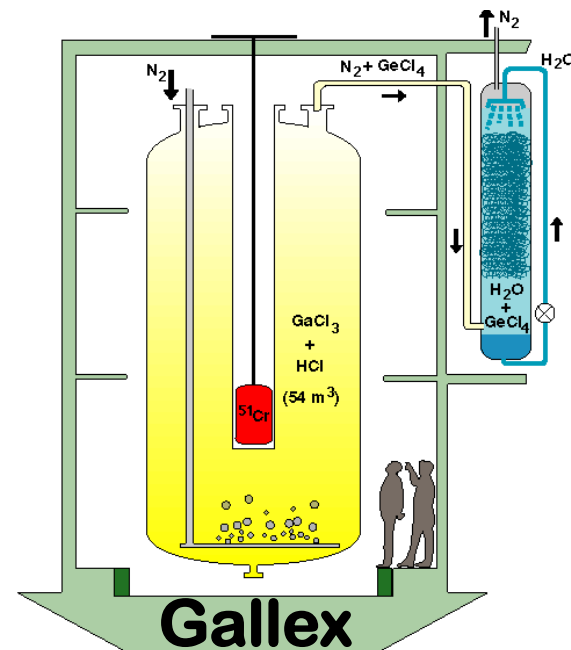
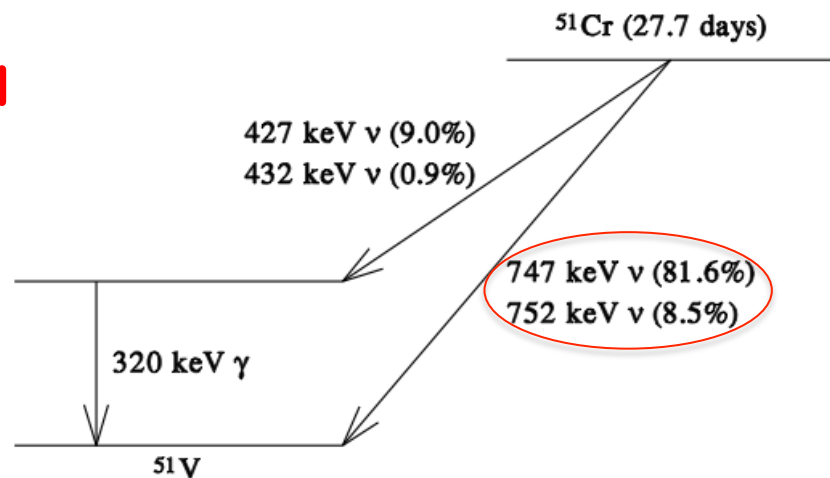


- **4 calibration runs with 19 – 62 PBq Electron Capture  $\nu_e$  emitters**

- Gallex,  $\langle L \rangle = 1.9$  m
    - $^{51}\text{Cr}$ , 750 keV
  - Sage,  $\langle L \rangle = 0.6$  m
    - $^{51}\text{Cr}$  &  $^{37}\text{Ar}$  (810 keV)

- **Deficit observed**

- $3\sigma$  anomaly
  - Supported by new  $^{71}\text{Ga}$  ( $^3\text{He}$ ,  $^3\text{H}$ )  $^{71}\text{Ge}$  cross section measurement

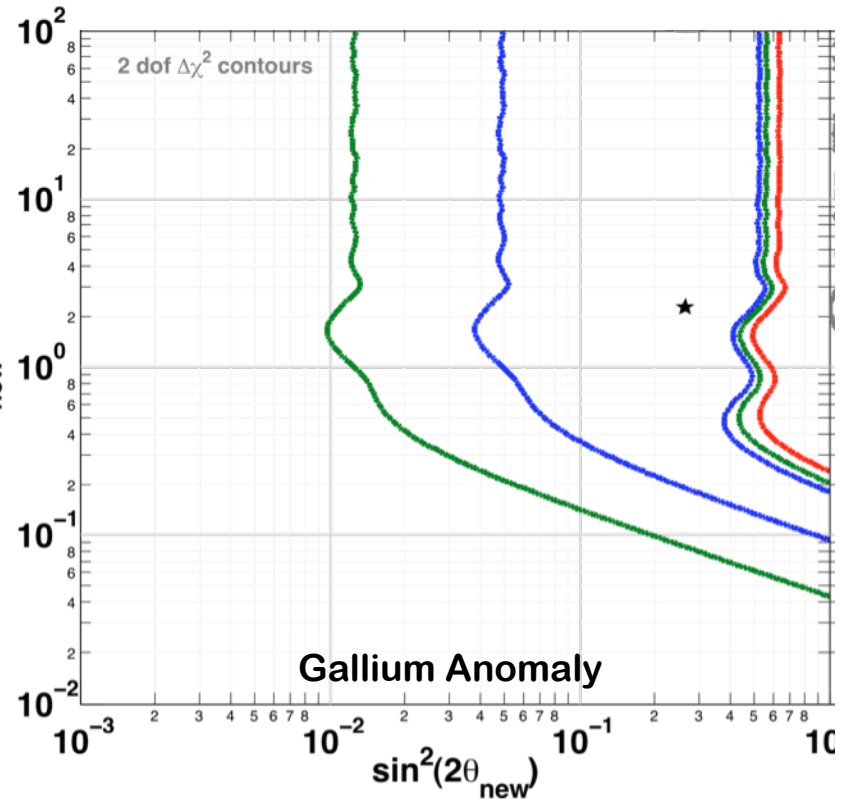
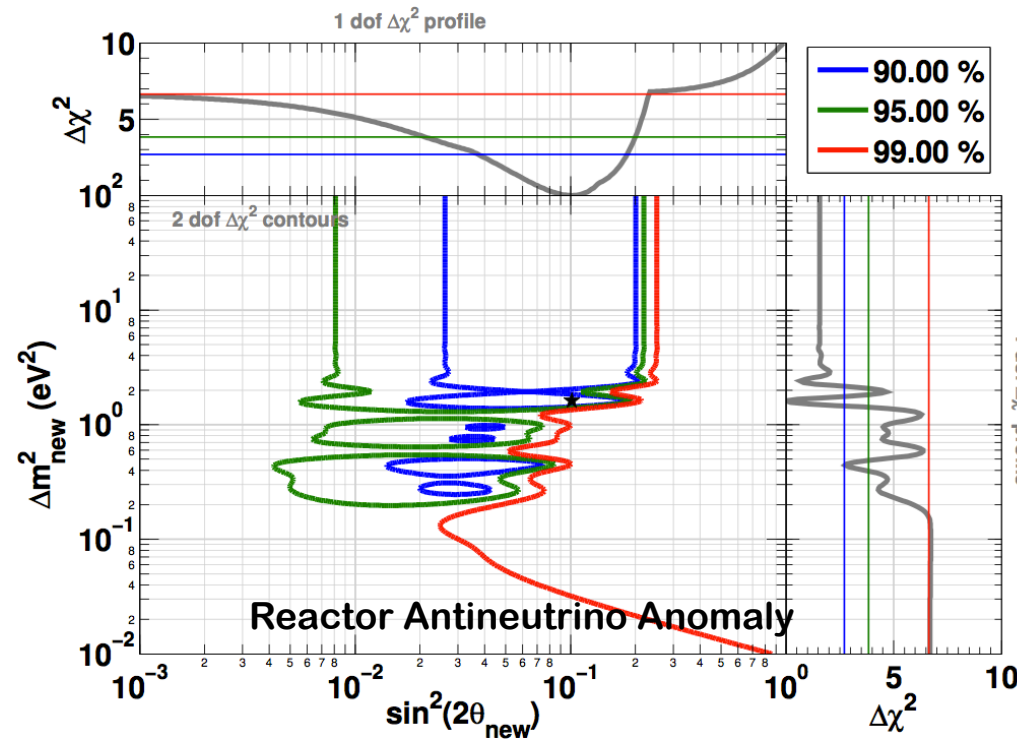




# Sterile Neutrinos

Fit to  $\bar{\nu}_e$  and  $\nu_e$  disappearance hypothesis (3+1 model)

$$\begin{pmatrix} \nu_e \\ \nu_s \end{pmatrix} = \begin{pmatrix} \cos \theta_{\text{new}} & \sin \theta_{\text{new}} \\ -\sin \theta_{\text{new}} & \cos \theta_{\text{new}} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_{\text{new}} \end{pmatrix}, P_{ee} = 1 - \sin^2(2\theta_{\text{new}}) \sin^2\left(\frac{\Delta m_{\text{new}}^2 L}{E}\right)$$



**No-oscillation hypothesis disfavored at >99.9% C.L.**



# Interpreting data as $\nu$ -oscillation

# Anomalous & Regular Results

Anomalous	Source	Type	Signal	Channel	Significance
LSND	Meson Decay-at-Rest	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	<u>Total Rate</u> , Energy	CC	3.8 $\sigma$
MiniBooNE	Meson Decay-in-Flight	$\nu_\mu \rightarrow \nu_e$	<u>Total Rate</u> , Energy	CC	3.8 $\sigma$
Gallium	Electron Capture	$\nu_e$ dis.	<u>Total Rate</u>	CC	2.7-3.0 $\sigma$
Reactor	Beta-decay	$\nu_e$ dis.	<u>Total Rate</u> , Energy	CC	2.7 $\sigma$

Regular	Source	Type	Signal	Channel
KARMEN Icarus/Opera	Meson Decay -at-Rest & Flight	$\nu_\mu \rightarrow \nu_e$	<u>Total Rate</u> , Energy	CC
CDHS/Minos/ MiniBooNE	Meson Decay-in-Flight	$\nu_\mu \rightarrow \nu_\mu$	<u>Total Rate</u> , Energy	CC
Minos	Meson Decay-in-Flight	$\nu_\mu \rightarrow \nu_s$	<u>Total Rate</u>	CC
T2K	Meson Decay-in-Flight	$\nu_e \rightarrow \nu_s$	<u>Total Rate</u> , Energy	CC

# Sterile- $\nu$ Phenomenology (3+1)

- $\bar{\nu}_e$  disappearance (Reactor, Gallium, ...)

$$\square \quad P_{ee} = 1 - \sin^2 2\theta_{ee} \sin^2 \frac{\Delta m_{41}^2}{4E} \quad \& \quad \sin^2 2\theta_{ee} = |U_{e4}|^2 (1 - |U_{e4}|^2)$$

- $\bar{\nu}_\mu$  disappearance (CDHS, MiniBOONE, Minos,...)

$$\square \quad P_{\mu\mu} = 1 - \sin^2 2\theta_{\mu\mu} \sin^2 \frac{\Delta m_{41}^2}{4E} \quad \& \quad \sin^2 2\theta_{\mu\mu} = |U_{\mu4}|^2 (1 - |U_{\mu4}|^2)$$

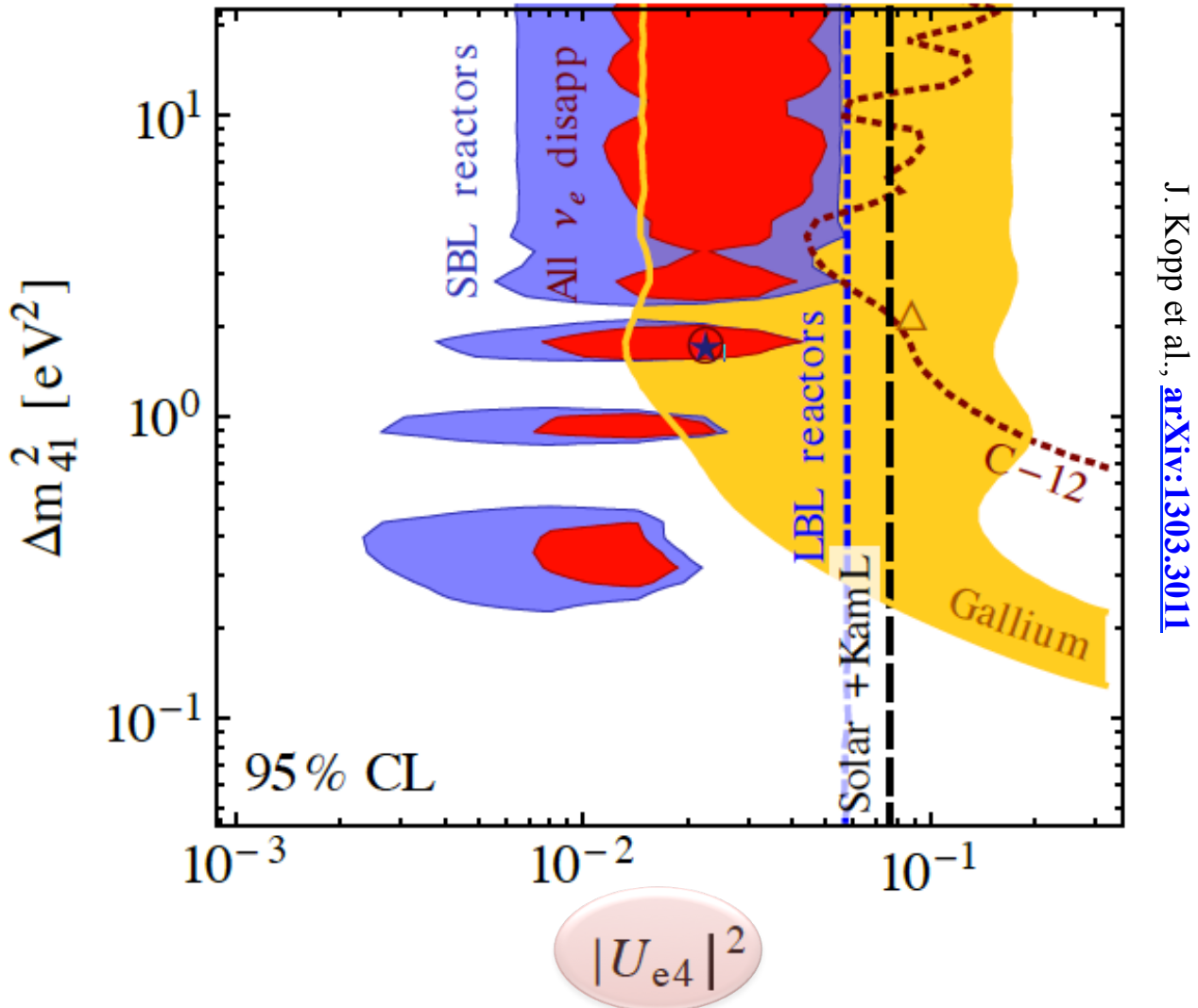
- $\bar{\nu}_e$  appearance (LSND, Karmen, MiniBooNE, Opera, Icarus...)

$$\square \quad P_{\mu e} = 4 \sin^2 2\theta_{\mu e} \sin^2 \frac{\Delta m_{41}^2}{4E} \quad \& \quad \sin^2 2\theta_{\mu e} \approx \frac{1}{4} \sin^2 2\theta_{ee} \sin^2 2\theta_{\mu\mu}$$

$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance requires  $\bar{\nu}_\mu$  &  $\bar{\nu}_e$  disappearance

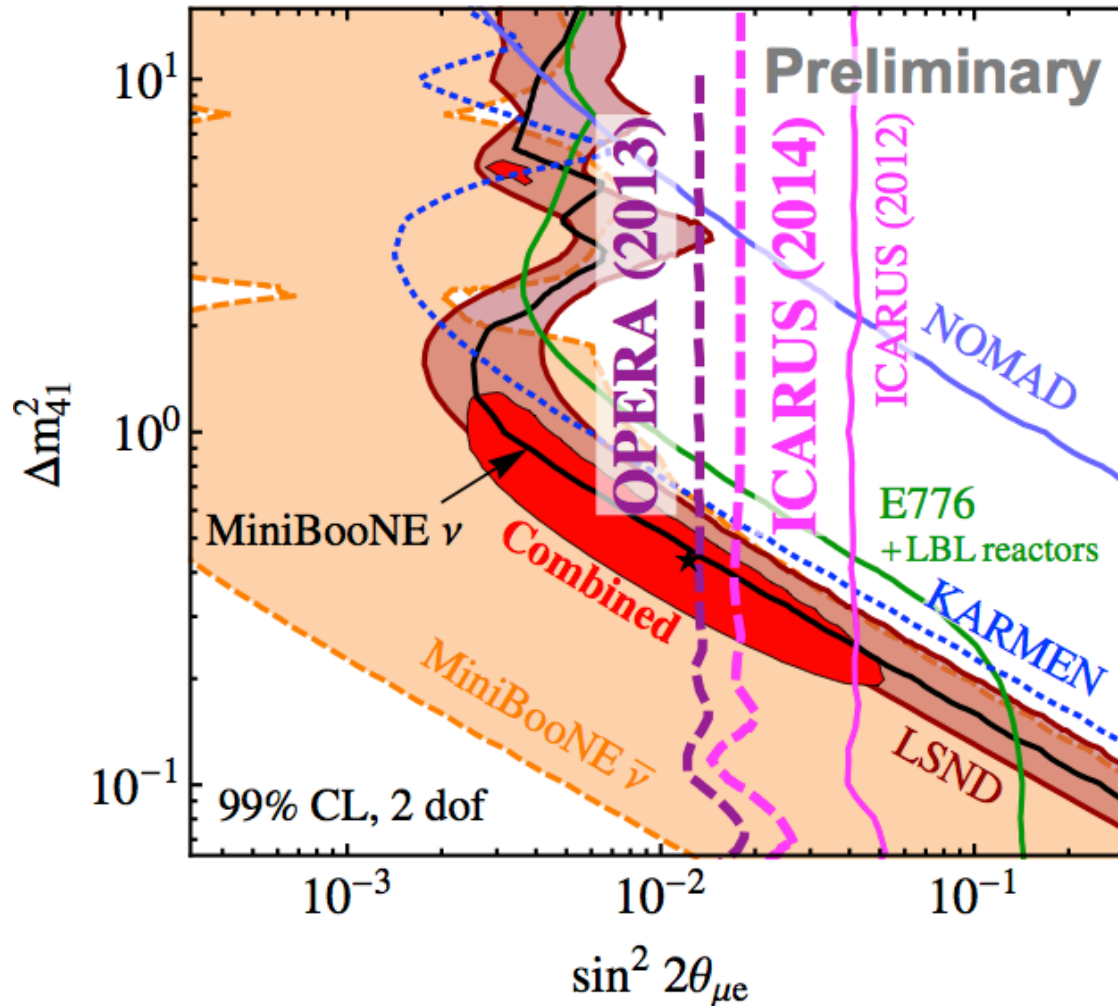
# $\bar{\nu}_e$ disappearance (3+1 scenario)

Data consistent with  $\bar{\nu}_e$  disappearance with  $L/E \approx 1$  m/MeV



# $\bar{\nu}_e$ appearance (3+1 scenario)

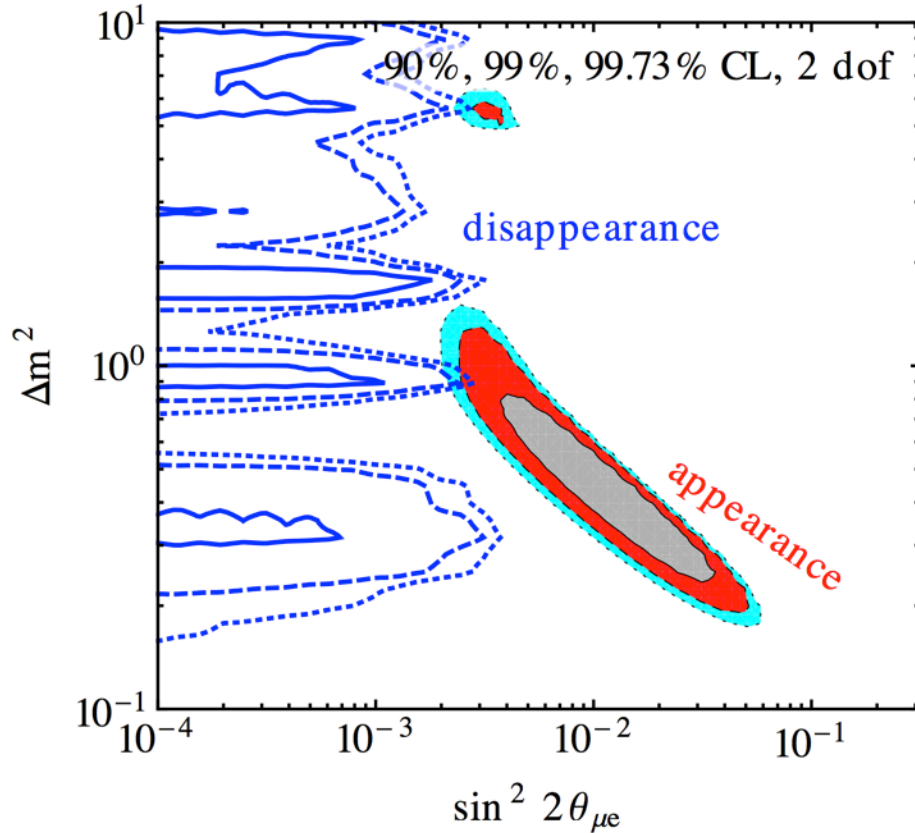
Consistent solution for  $\bar{\nu}_e$  appearance with  $L/E \approx 1$  m/MeV



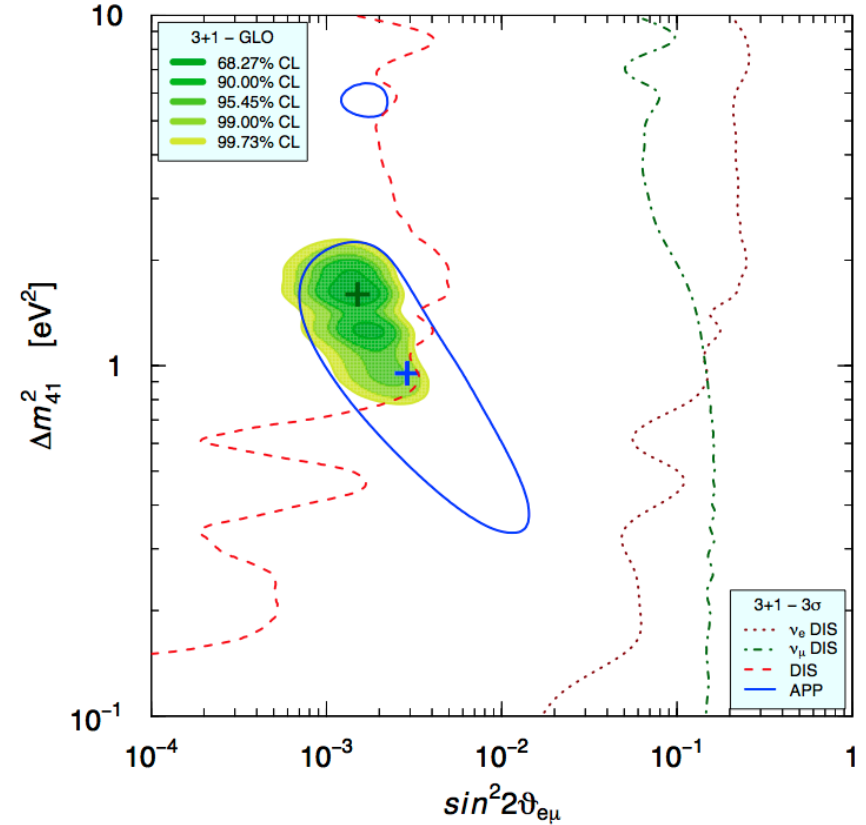
J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

# Appearance VS Disappearance

J. Kopp et al., arXiv:1303.3011



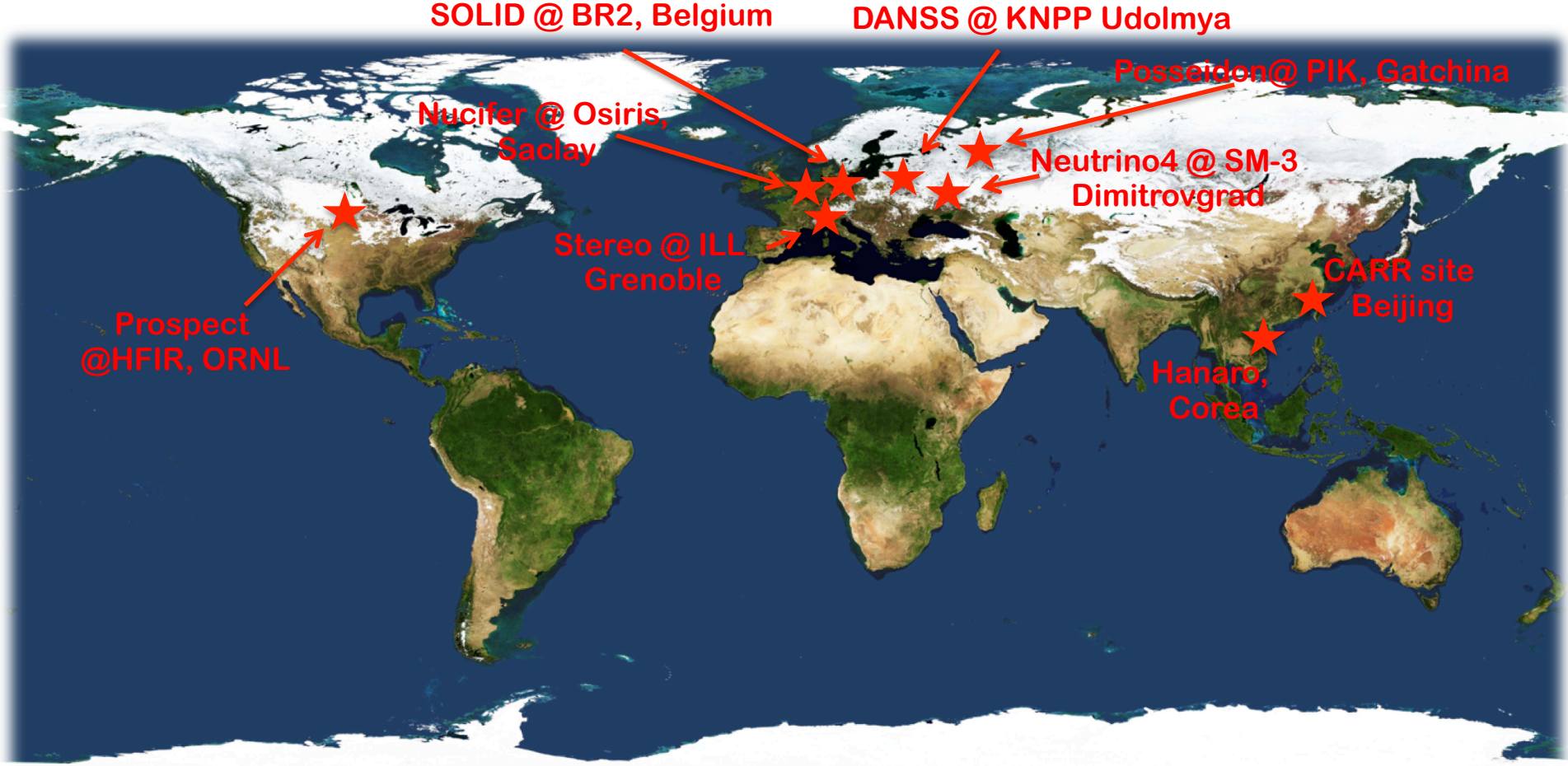
C. Giunti et al., arXiv:1308.5288



**Tension between  $\bar{\nu}_e/\nu_e$  appearance/disappearance and  $\bar{\nu}_\mu/\nu_\mu$  disappearance (3+1 & 3+2 models)**

# Experimental Prospects

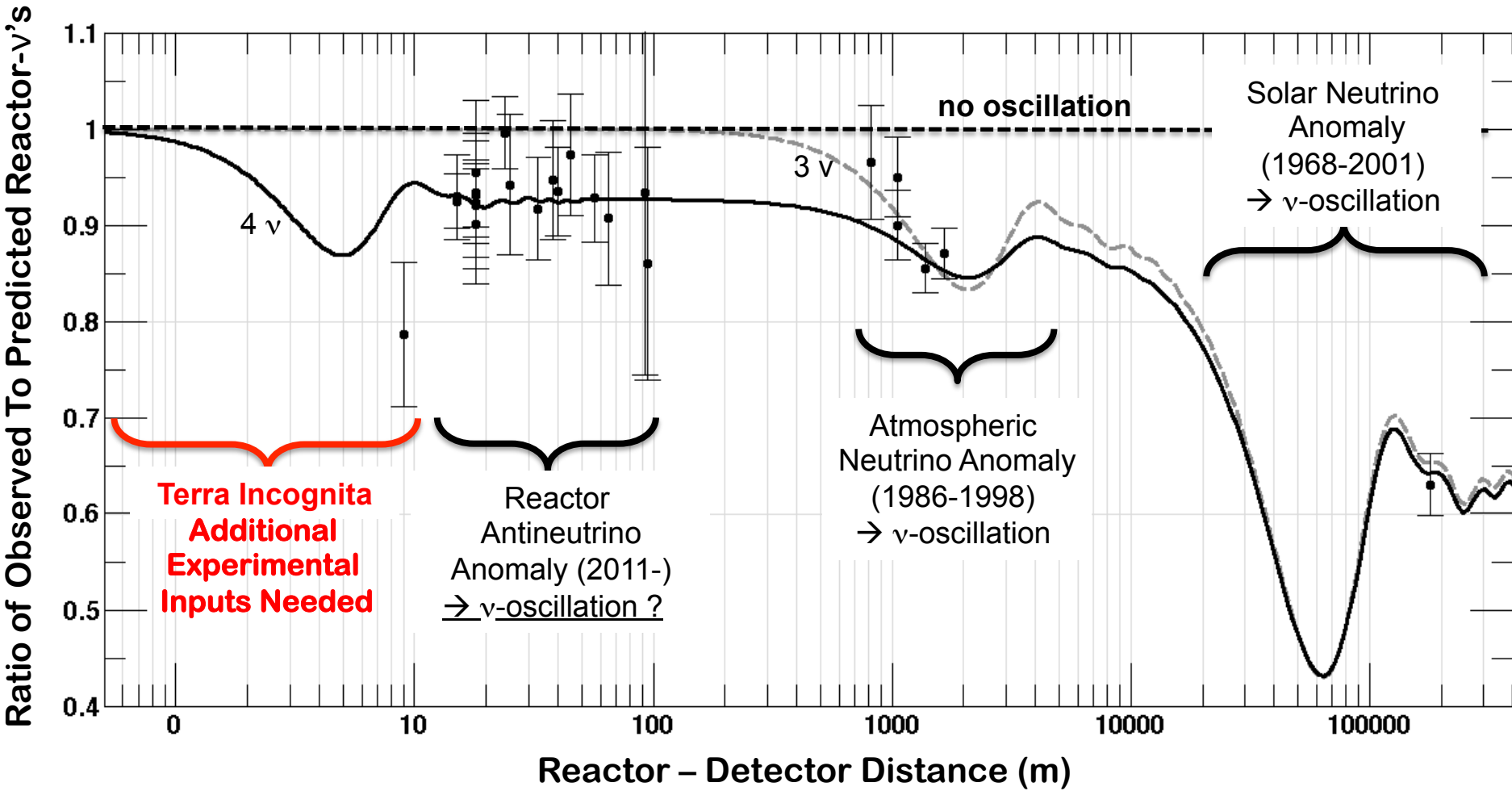
# Experimental Prospect: @ Nuclear Reactor



**Test of both reactor & gallium anomalies**



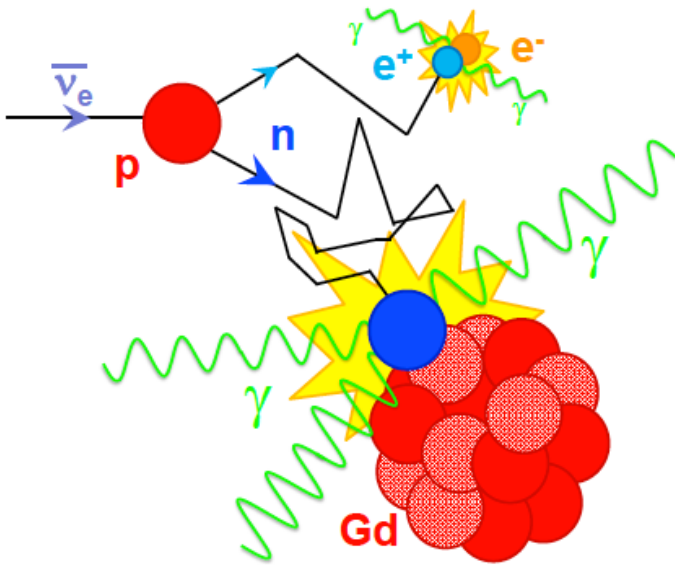
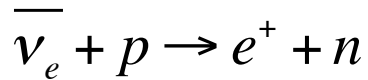
# Experimental Artifact or New Physics?



# Testing $(\bar{\nu}_e)$ disappearance anomalies

- Need robust test, beyond the current mean deviation from reactor predicted rate
- **Input from sterile neutrino fits**
  - $\Delta m^2 \approx 0.1-10 \text{ eV}^2 \rightarrow L_{\text{osc}}(\text{m}) = 2.5 \frac{E(\text{MeV})}{\Delta m^2(\text{eV}^2)} \approx 2-10 \text{ m}$
  - $\sin^2(2\theta_{ee}) \approx 0.01-0.15$
- **Experimental specifications**
  - Compact source, <1 meter scale
  - Good vertex and energy resolutions
  - High statistics (few % stat. uncertainty)
  - Few % syst. uncertainty  $\rightarrow$  low backgrounds
- **Search for a new oscillation pattern in E & L completed by normalization information**

## Inverse Beta Decay



**Selective coincidence**  
**e<sup>+</sup> prompt signal & n-capture**

## Background rejection

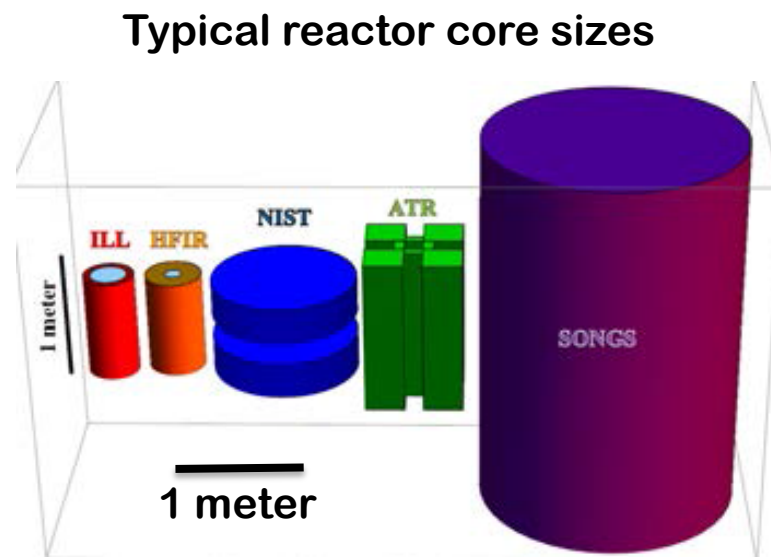
- **Accidental  $\gamma - \gamma$ /neutron coincidence**

- Shielding
- Segmentation
- Neutron discrimination

- **Fast-n correlated background**

- Rejection of recoil protons with PSD
- Cosmic rays induced:
  - Reactor OFF data
  - Overburden > few m.w.e.
- Reactor induced neutrons: a killer!
  - must be negligible

- **Compact reactor core**
  - No oscillation smearing
- **High statistics (few 100 evts/day/t)**
  - High Power (10-3000 MW)
  - Short baselines (5-50 m)
- **Fuel**
  - Highly enriched  $^{235}\text{U}$  preferable ( $^{235}\text{U}$  fission spectrum)
- **Reactor ON/OFF periods**
  - Moderate overburden compensated by accurate measurement of the cosmogenic background
- **But challenging reactor-induced backgrounds ( $\gamma$  and  $n$ )**
  - Need site characterization & specific shieldings



# Reactor v Proposals

Experiment Type	Projects	$P_{Th}$	$M_{det}$	L	Depth
Mature Gd-doped LS detector Technology	Nucifer (FRA)	70 MW	0.7 tons	7 m	Few mwe
	Stéréo (FRA)	50 MW	2 tons	[8-11] m	10 mwe
	Neutrino 4 (RU)	100 MW	2 tons	[6-12] m	Surf.
Highly segmented detector for background reduction	DANSS (RU)	1 GW	1 ton	[10-12] m	50 mwe
	SoLid (UK)	45-80 MW	3 tons	8 m	10 m
Enhanced neutron Tagging					
	Hanaro (KO)	30 MW	0.5 t	6 m	Few mwe
2 detector complex or Moving detector	Prospect	85 MW	-	7m & 18m	Surf.
	China project			-	
	DANSS/Neutrino4			Movable detector	

# Nucifer @ OSIRIS (Gd-LS)

Designed for non proliferation studies

## Osiris research reactor

- Saclay
- 70 MW, 20%  $^{235}\text{U}$

## Detector

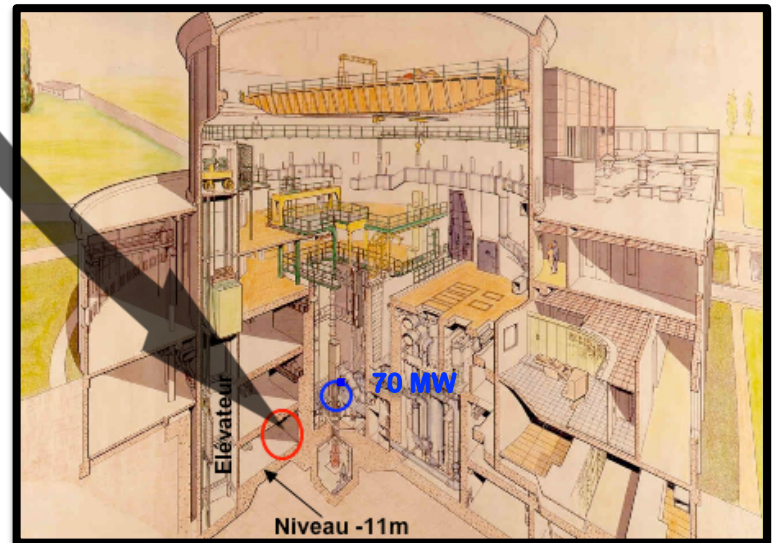
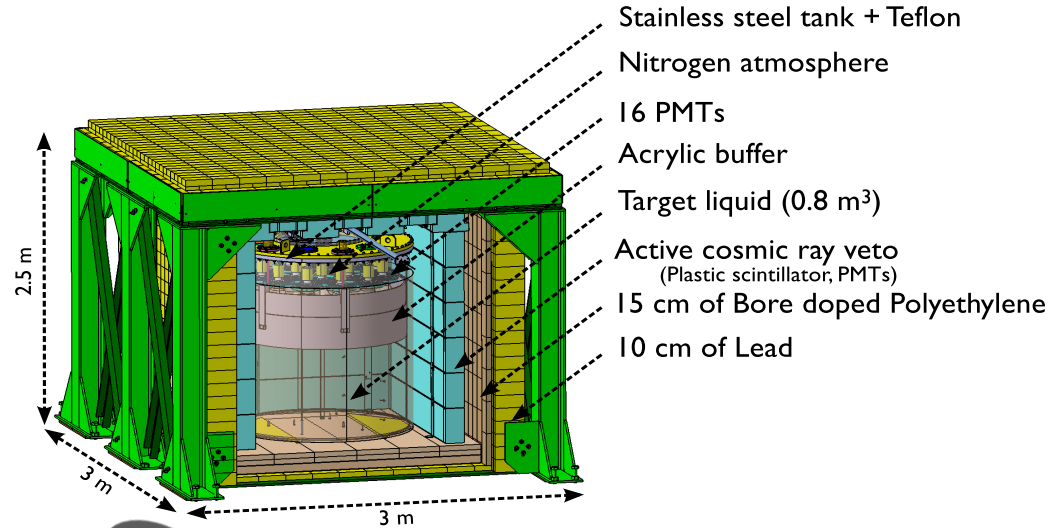
- 850 kg Gd-loaded LS
- Shallow depth (few mwe)

## Sensitivity to Sterile- $\nu$ :

- Compact core:  $60 \times 60 \times 60 \text{ cm}^3$
- Short baseline: 7 m
- Simple design (counting exp.)
- Challenging reactor acc. Bkgs

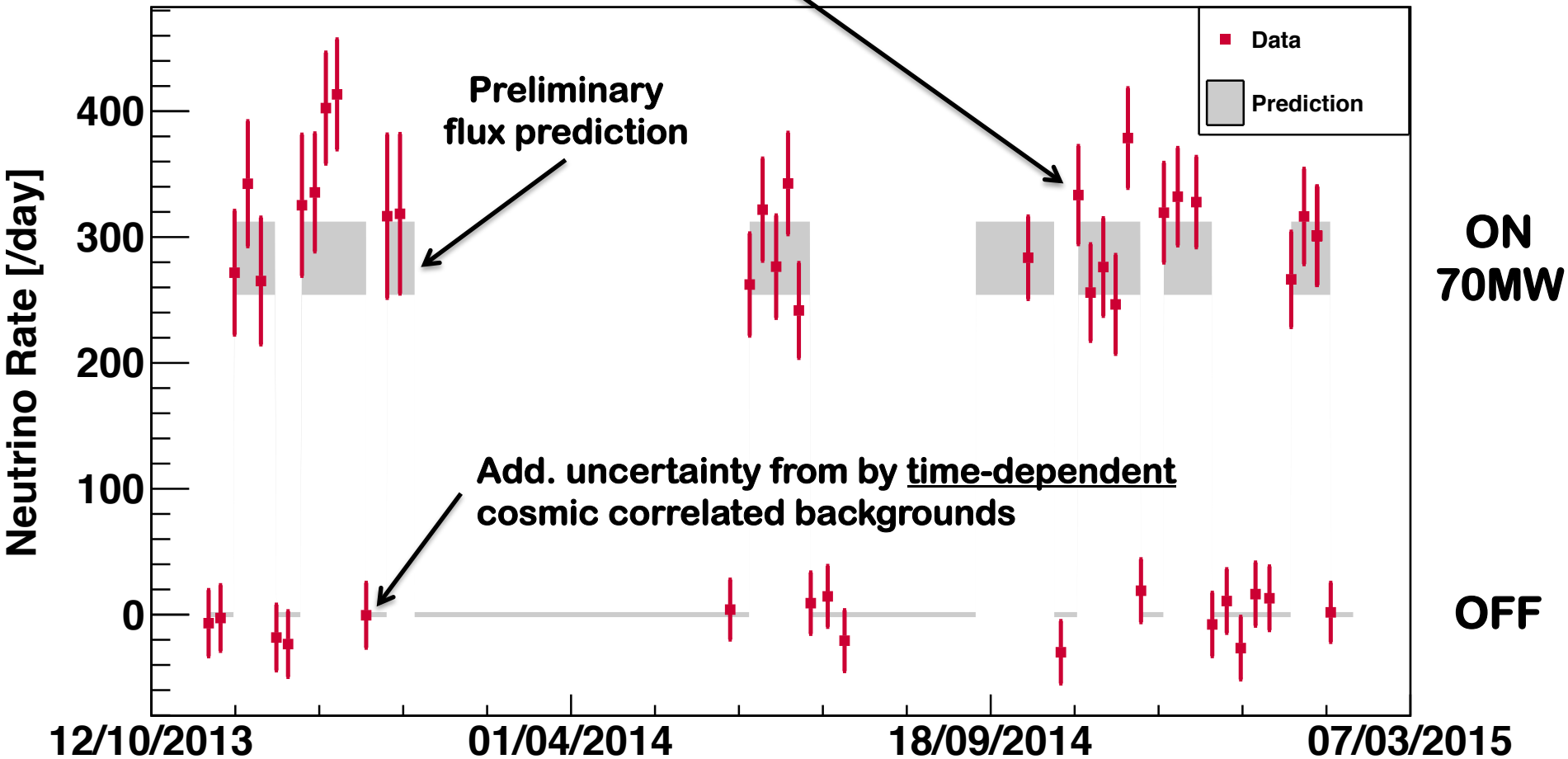
## Data taking ongoing

- $\approx 300 \text{ v/day}$
- Until end of 2015



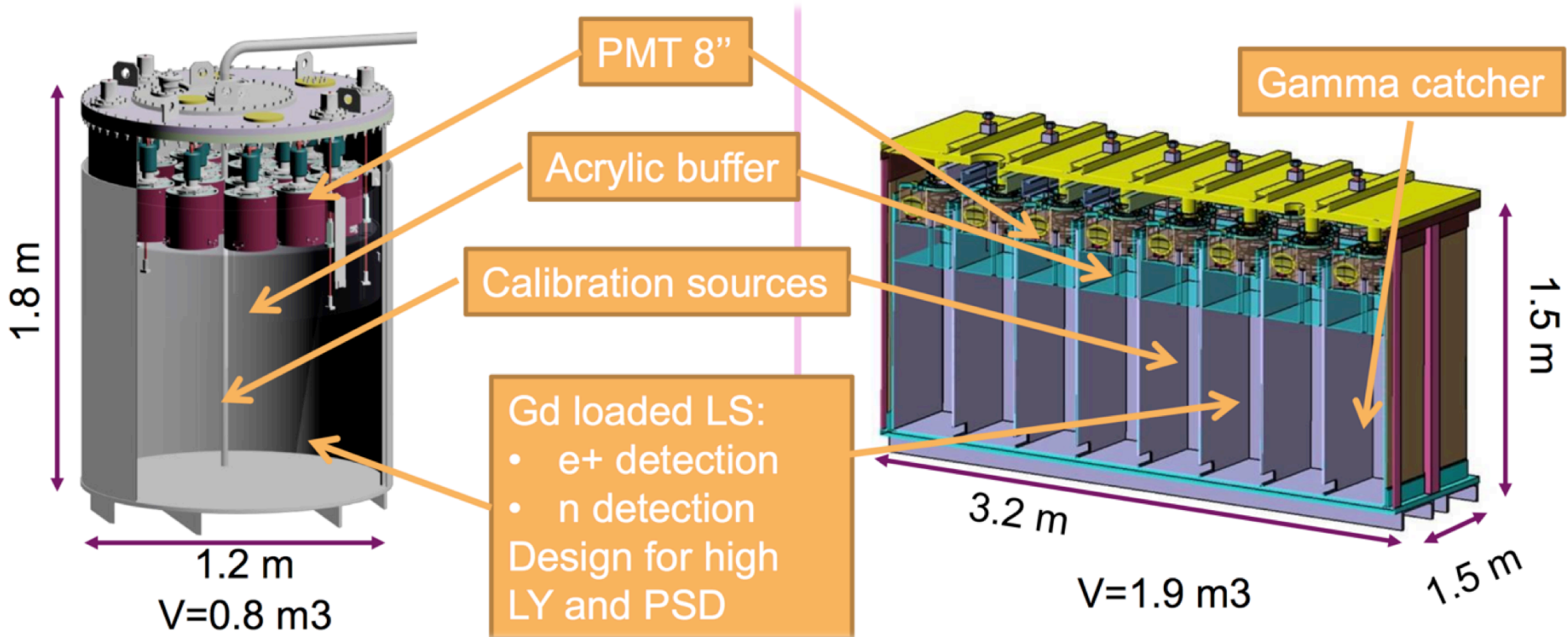
# First Nucifer Results

Uncertainty dominated by subtraction of reactor induced accidental backgrounds – S/N=1/10





# Nucifer / Stéréo Design



- Rate only analysis
- RMS/peak = 30% for 2 MeV e<sup>+</sup>

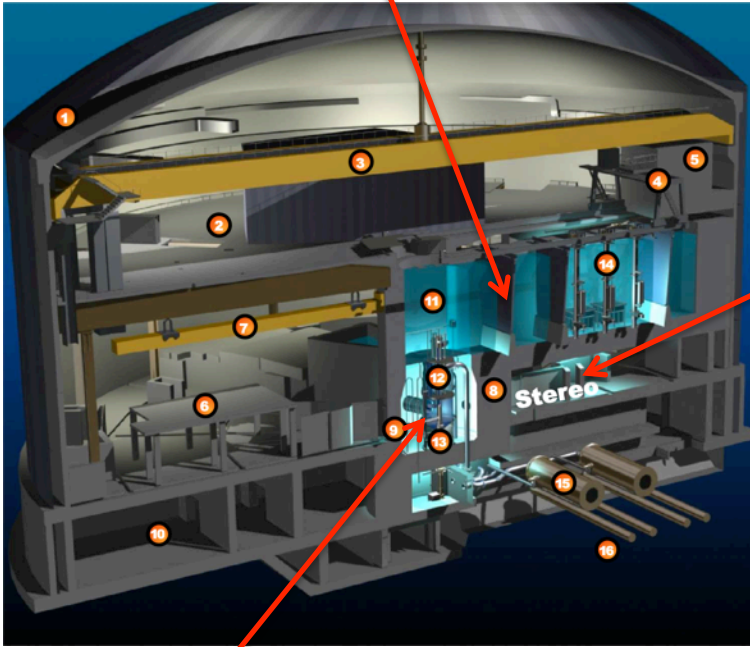
- Shape and rate analysis
- RMS/peak = 11.5% for 2 MeV e<sup>+</sup>



# Stéréo @ ILL (Gd-LS)

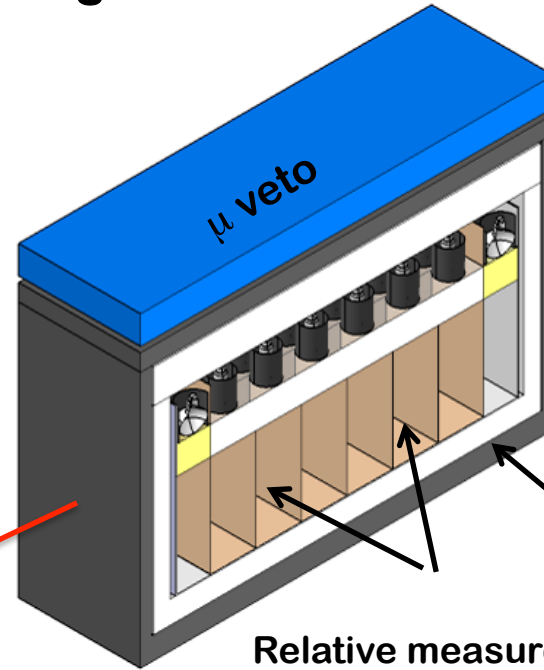
## Start Data Taking in June 2016

factor 4 attenuation of  $\mu$  vertical flux from water pool (15 m.w.e)

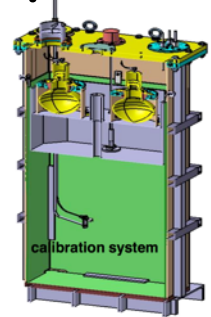


50 MW core  
h=80cm,  $\Phi=40$ cm

[9-11] m  
baseline range

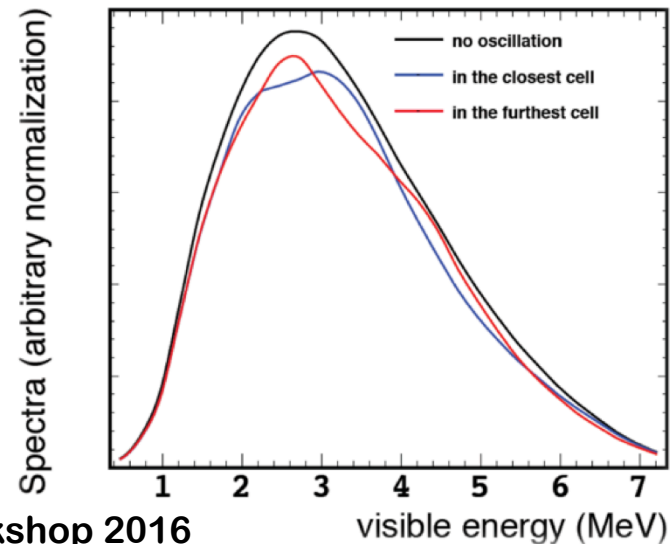


proto

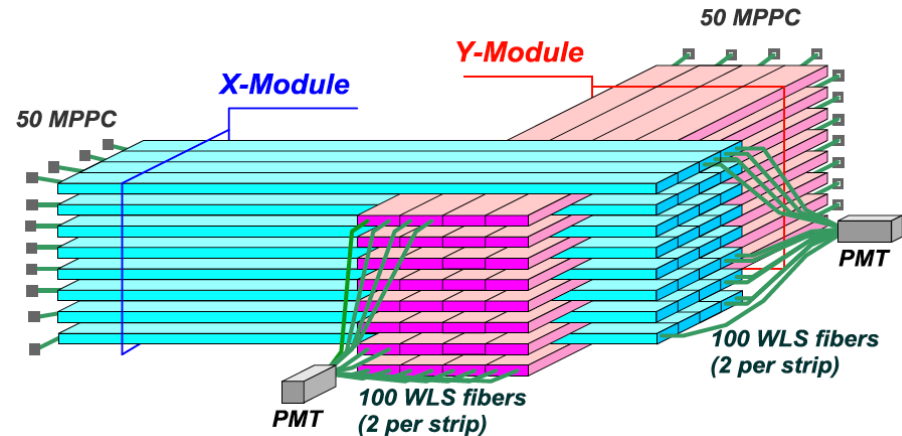
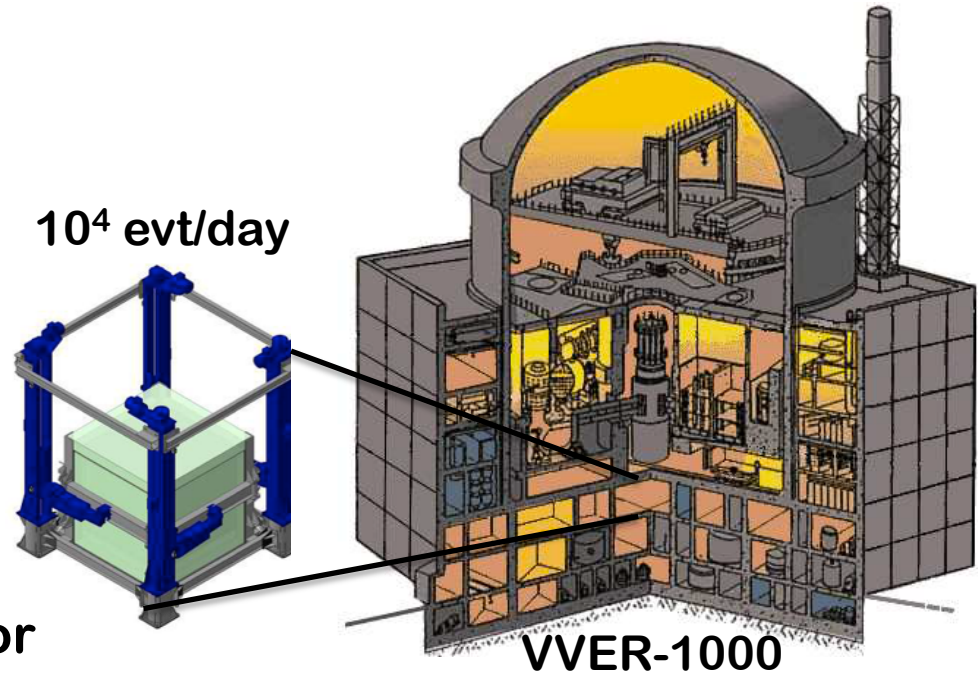


70 tons  $\gamma$  and n shielding

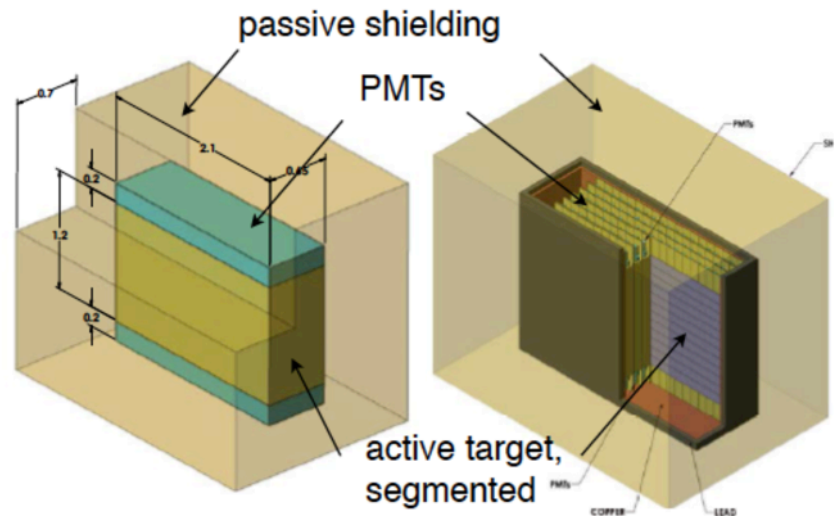
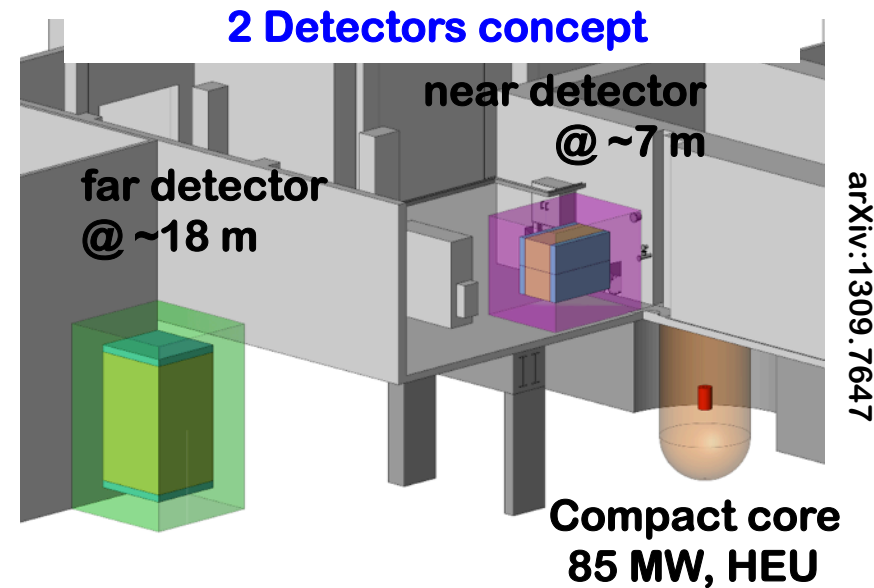
Relative measurement in 6 cells



- 1 GW extended core
- Good overburden
- Vertical motion of the detector (9.7-12.2 m)
- Highly segmented detector  
→ background rejection
- Plastic strips with Gd-loaded interlayer, WLS fibers readout
- Started in 2015

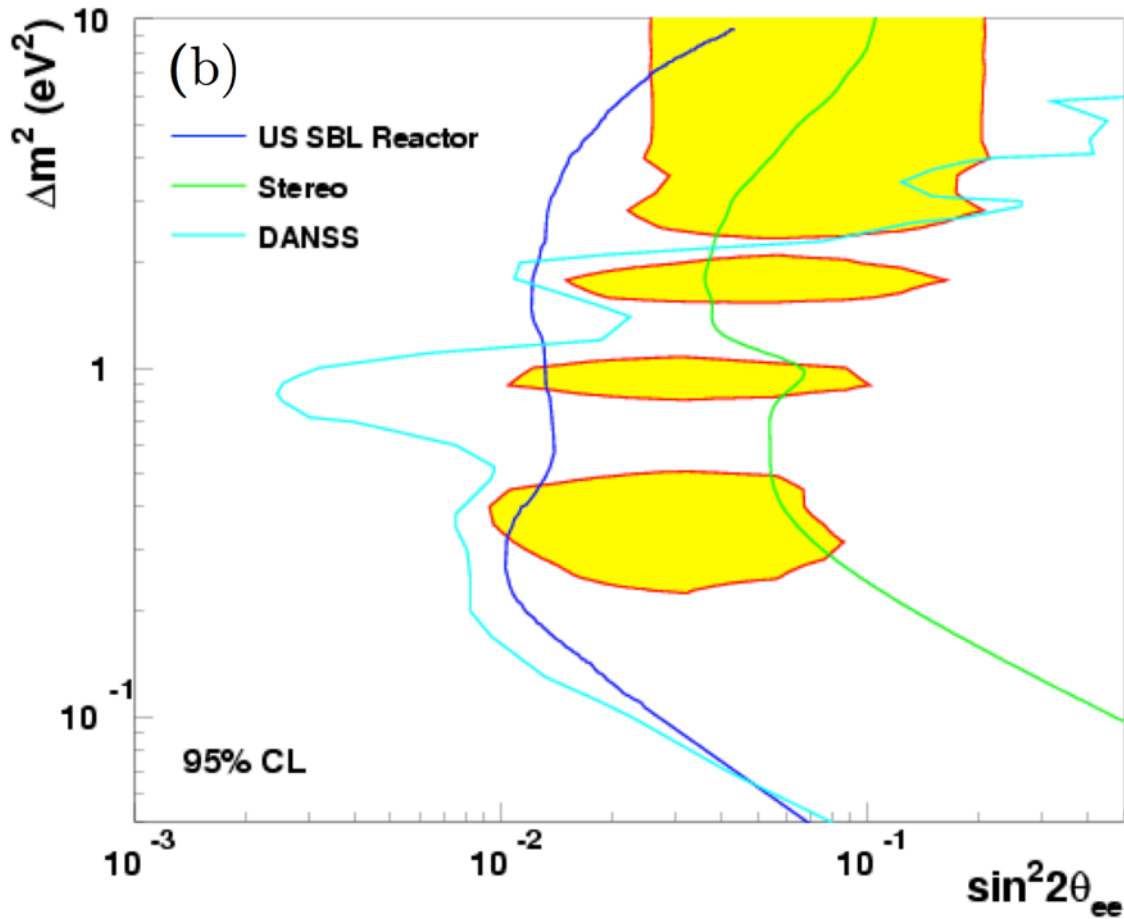


- Reactor sites
  - HFIR @ ORNL – 85 MW
- 7-18 m baselines
- Surface location
- Detector
  - Segmented
  - $^6\text{Li}$ -doped (+PSD)
- Status:
  - Site characterization
  - R&D ongoing
  - Construction start in 2015



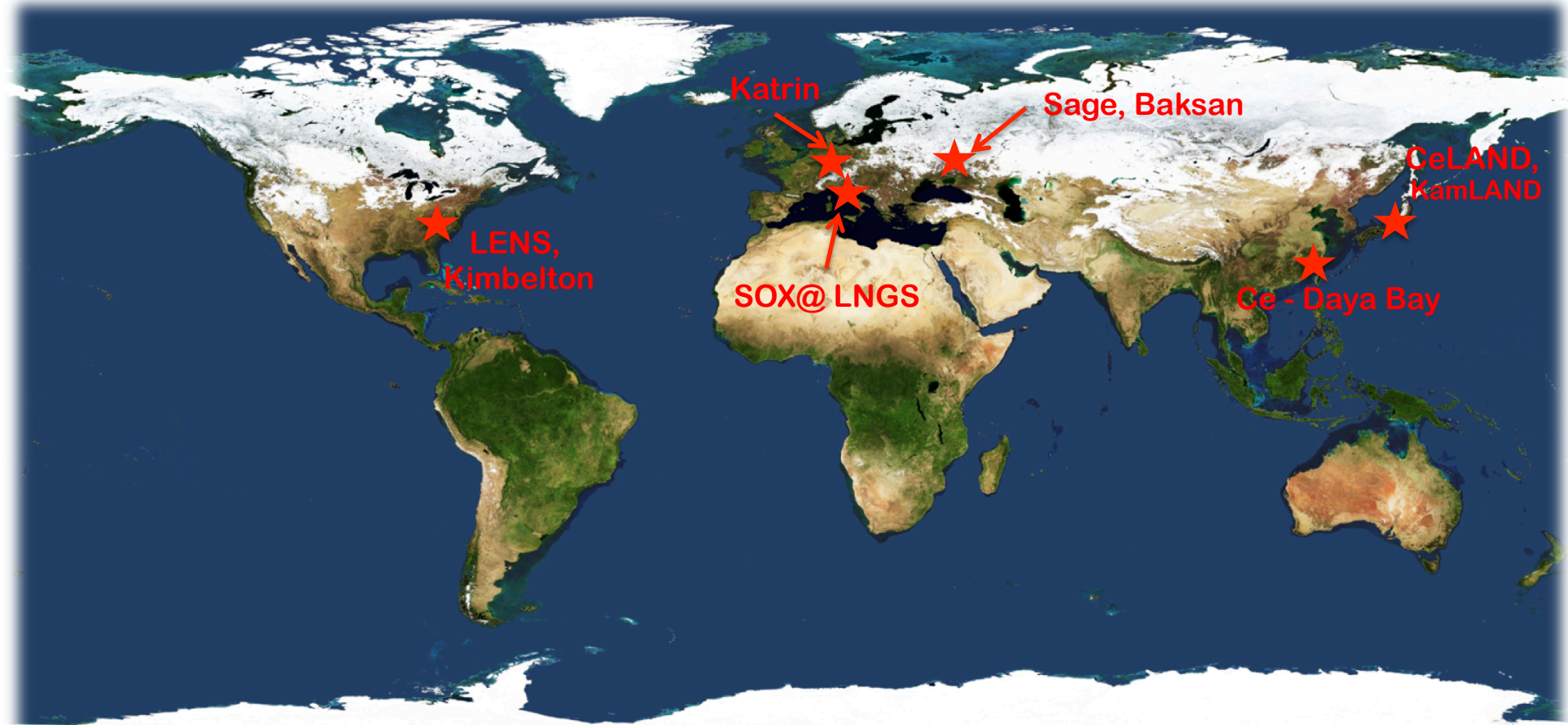
# Reactor Experiment Sensitivity

All current projects have the sensitivity to test the reactor anomaly space of parameters,  $\Delta m^2 > 0.1 \text{ eV}^2$ ,  $\sin^2 2\theta_{ee} > 0.05$





# Experimental Program: @ Neutrino Generator



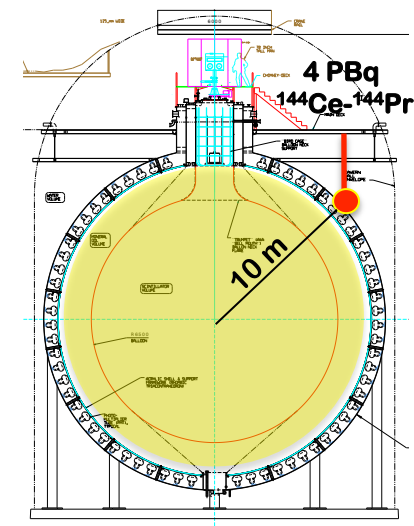
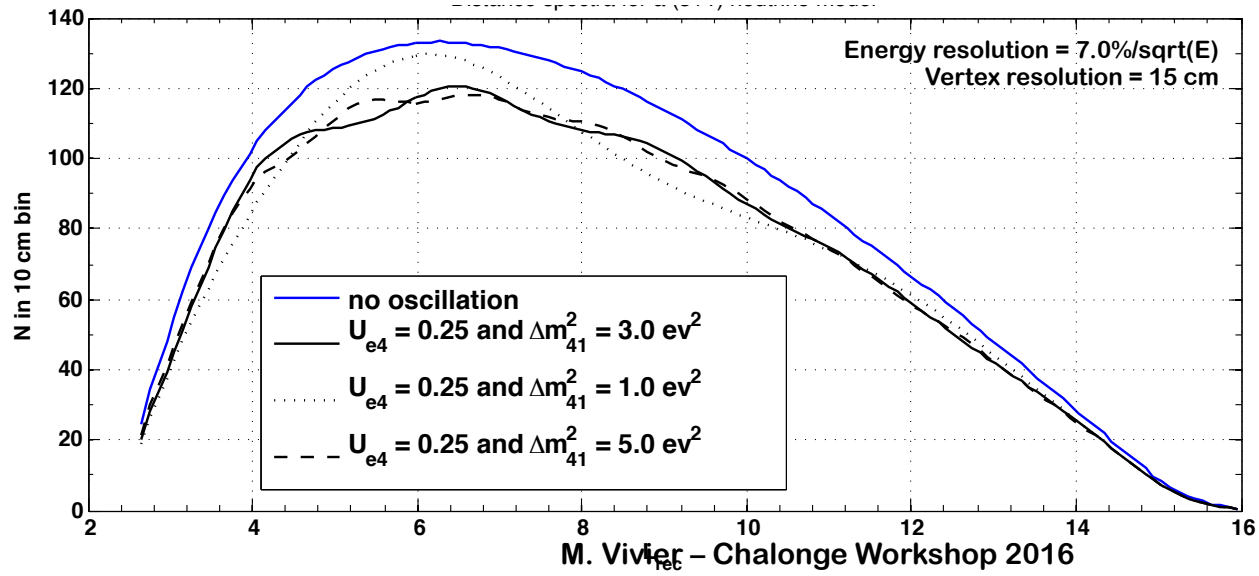
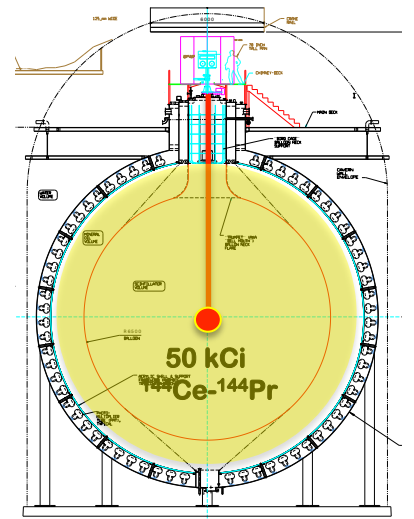
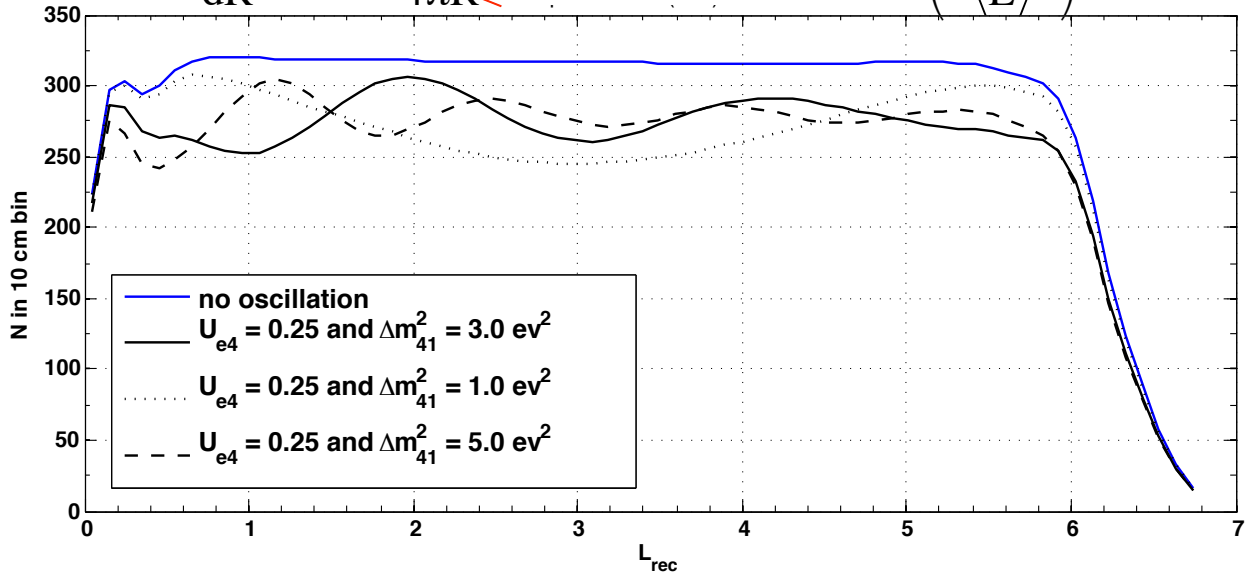
**Test of both reactor & gallium anomalies**

# $\nu$ Generator Proposals

Type	Detection	Background	Isotope	Production	Activity	Projects
$\nu_e$	$\nu_e e \rightarrow \nu_e e$ 5% $E_{res}$ 15cm $R_{res}$  or Radio-chemical	Detector Radioactivity  Solar $\nu$ (irreducible)  $\nu$ generator impurities	$^{51}\text{Cr}$ 0.75 MeV $t_{1/2}=26\text{d}$	$n_{th}$ irradiation in Reactor	>110 PBq	Sage LENS
					>370 PBq	SOX-Cr (SNO+)
			$^{37}\text{Ar}$ 0.8 MeV $t_{1/2}=35\text{d}$	$n_{fast}$ irradiation in Reactor (breeder)	>37 PBq	-
$\bar{\nu}_e$	$\bar{\nu}_e p \rightarrow e^+ n$ $E_{th}=1.8\text{ MeV}$  ( $e^+, n$ )  5% $E_{res}$ 15cm $R_{res}$	reactor $\nu$ , geo $\nu$ ,  $\nu$ generator impurities	$^{144}\text{Ce}$ $E < 3\text{ MeV}$ $t_{1/2}=285\text{d}$	spent nuclear fuel reprocessing + REE extraction	3.7 PBq	CeLAND Ce-SOX
					18.5 PBq	Daya-Bay
			$^{90}\text{Sr}$ $^{106}\text{Rh}$		-	-
	$^3\text{H} \rightarrow \text{He} e^- \bar{\nu}_e$ EC/ $\beta$ -decay	Kink search	$^3\text{H}$ $E < 18\text{ keV}$	Irradiation in reactors	110 GBq	KATRIN (Mare/Echo)

# Search for $\bar{\nu}_e \rightarrow \bar{\nu}_s$ with $^{51}\text{Cr}/^{144}\text{Ce}$

$$\frac{dN}{dR}(R,t) \propto \frac{A(t)}{4\pi R^2} \times \langle \sigma \rangle \times N_p \times \cancel{4\pi R^2} \times P_{ee} \left( \frac{\Delta m^2 R}{\langle E \rangle} \right)$$





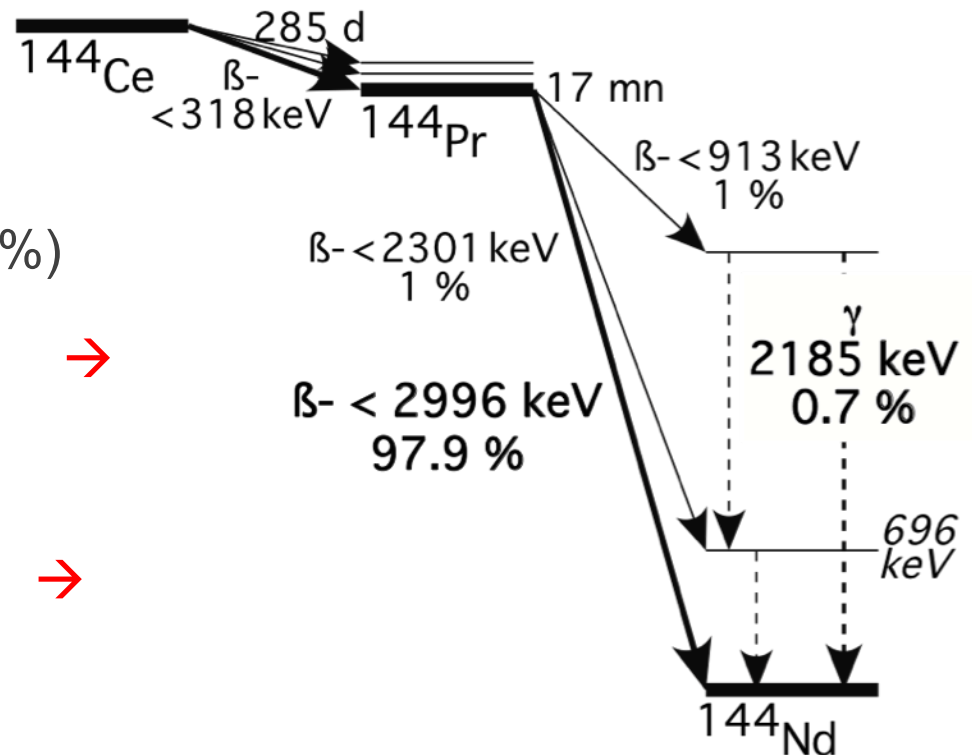
# $^{144}\text{Ce}-^{144}\text{Pr} \bar{\nu}$ generator

erc

- 1<sup>st</sup> Trick:  $\bar{\nu}_e$  source detected via  $\bar{\nu}_e + p \rightarrow e^+ + n$  (Thr=1.8 MeV)
  - High IBD cross section  $\rightarrow$  few PBq activity (3-5 PBq)
  - $(e^+, n)$  detected in coincidence  $\rightarrow$  Strong background reduction

## 2<sup>nd</sup> Trick: $^{144}\text{Ce}-^{144}\text{Pr}$

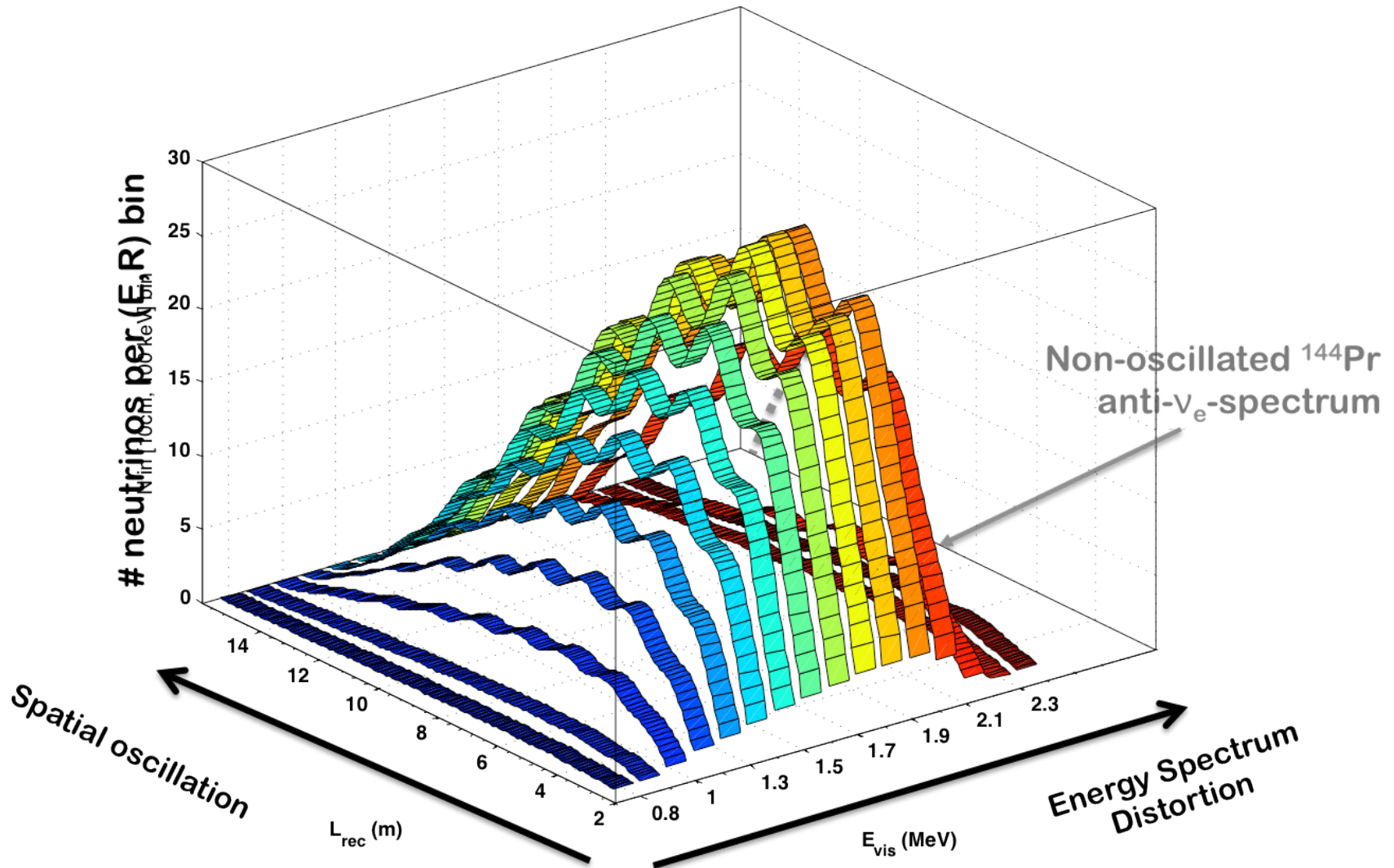
- Abundant fission product (5%)
- $^{144}\text{Ce}$ : long-lived & low- $Q_\beta$   $\rightarrow$  Enough time to produce, transport, use
- $^{144}\text{Pr}$ : short-lived & high- $Q_\beta$   $\rightarrow$   $\bar{\nu}_e$ -emitter above threshold



# $^{144}\text{Ce}$ - $^{144}\text{Pr}$ Signal

100 kCi  $^{144}\text{Ce}$ - $^{144}\text{Pr}$  – 8.3 m from detector center – 1.5 year

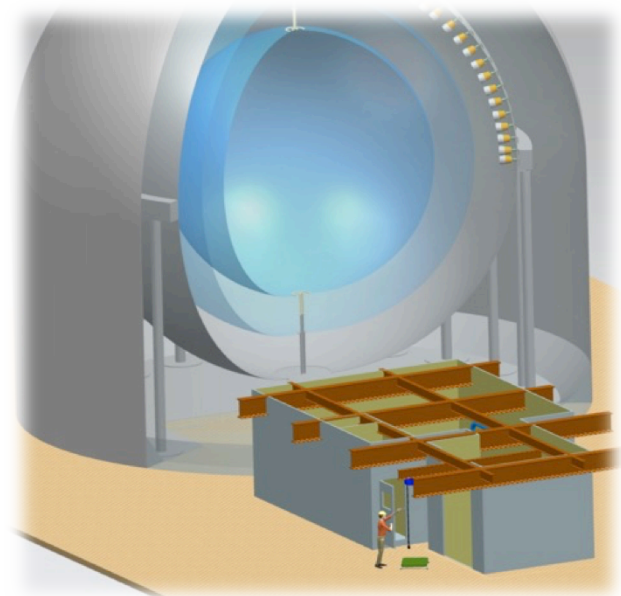
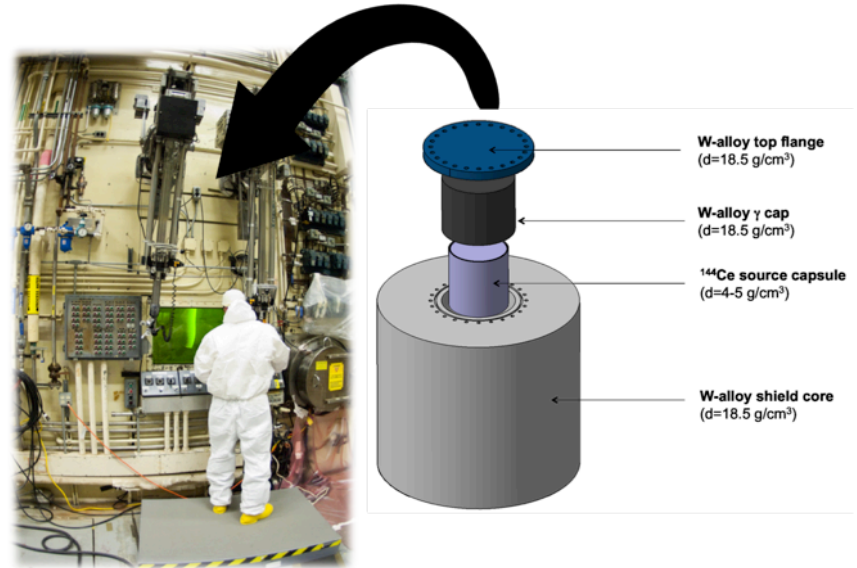
2-D reconstructed spectrum for  $U_{e4} = 0.25$  and  $\Delta m_{41}^2 = 3.0 \text{ eV}^2$



# $^{144}\text{Ce}$ - $^{144}\text{Pr}$ : CeSOX in BX

erc

- 4 PBq of  $^{144}\text{Ce}$ - $^{144}\text{Pr}$  ( $\text{CeO}_2$ )
- **Production at Mayak Facility (RU) in 2015/16 (1 y)**
  - Standard SNF reprocessing
  - Ce extraction through displacement chromatography
- **Need 19 cm tungsten-shield**
  - Manufacturing ongoing
- **Borexino getting ready**
  - Tunnel below the detector
  - 8.25 m from center
- **Deployment in 9/2016**
  - 1.5 y data taking
  - 10000/1.5 y interactions expected

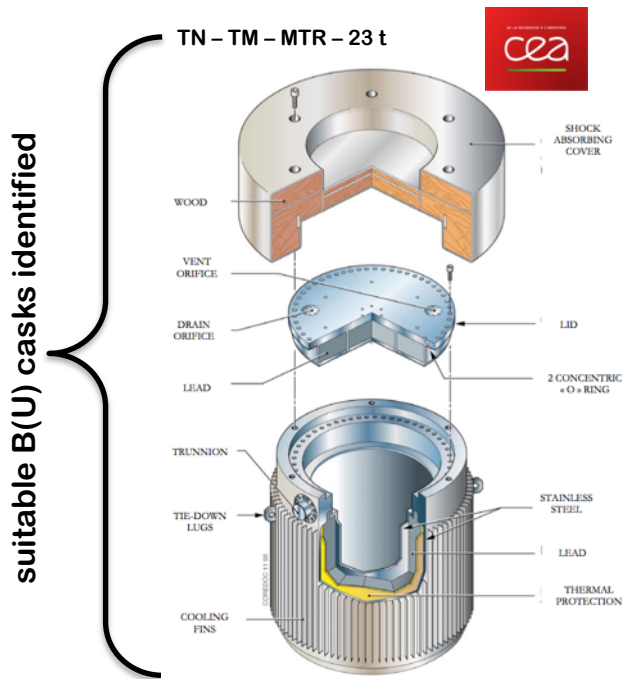


# CeSOX: a Challenging Logistic



## IAEA rules on Safe Transportation of Radioactive Material

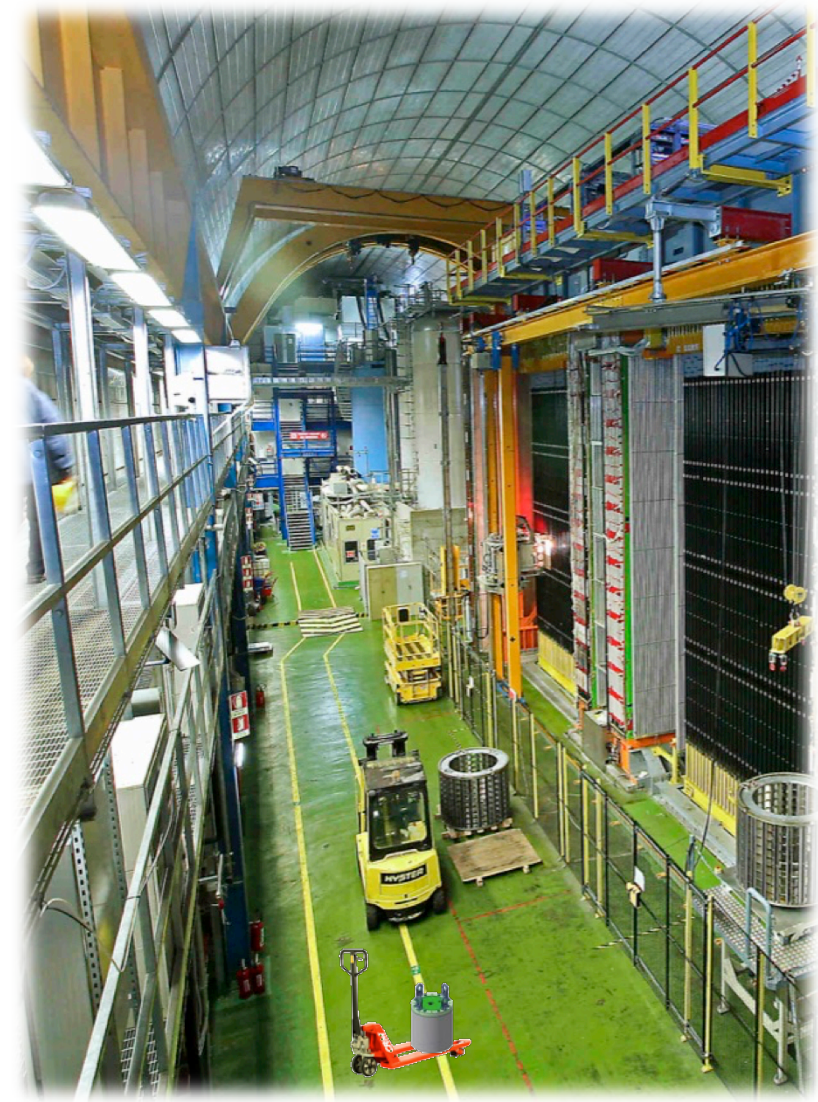
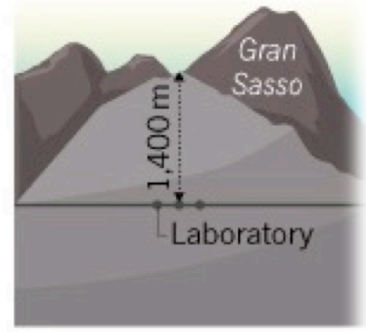
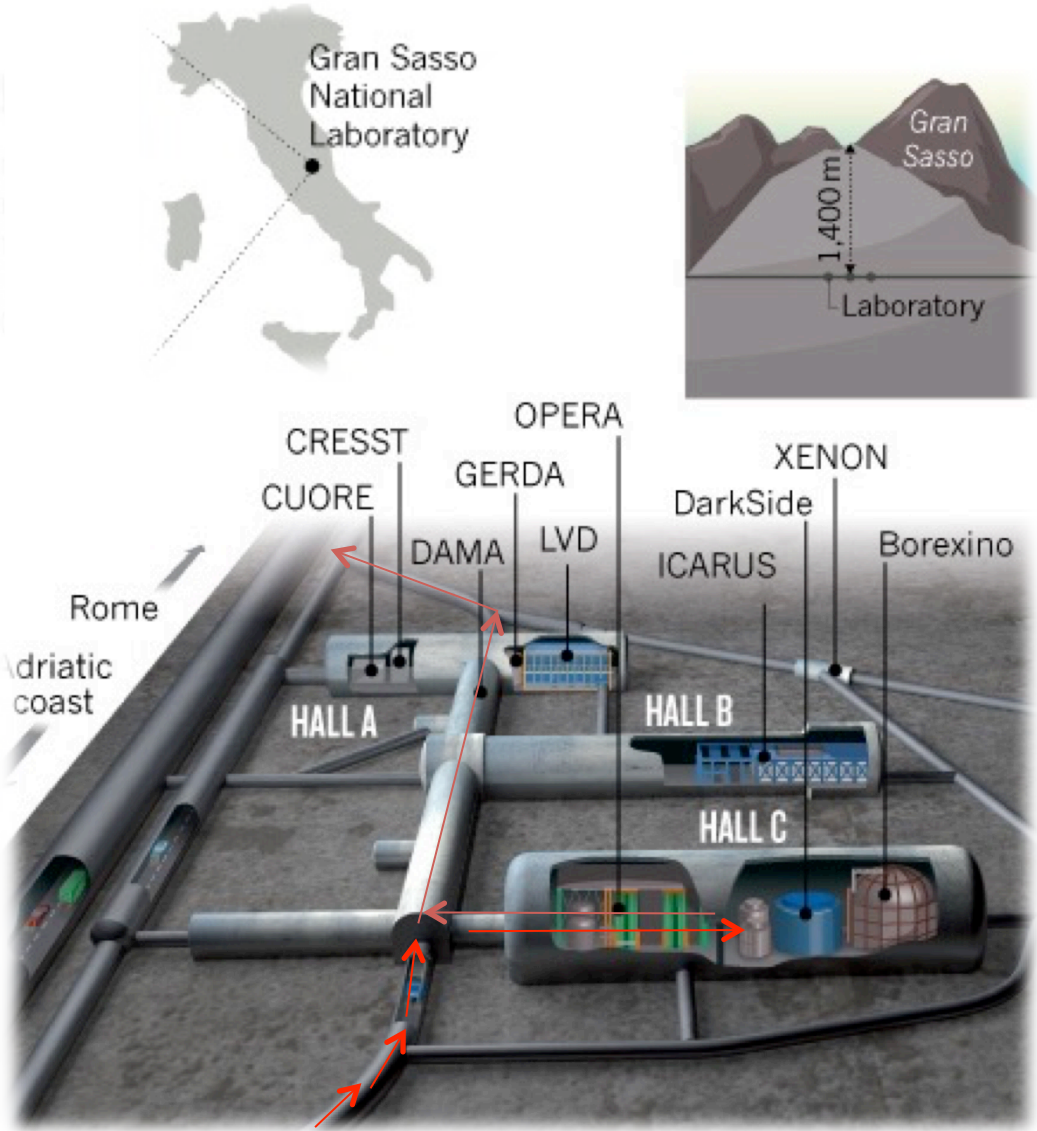
### A) Suitable certified transport container: licensing ongoing



### B) Route: Mayak – St Petersburg – Le Havre – Gran Sasso (duration of the trip: ≈3 weeks)



# Arrival at LNGS

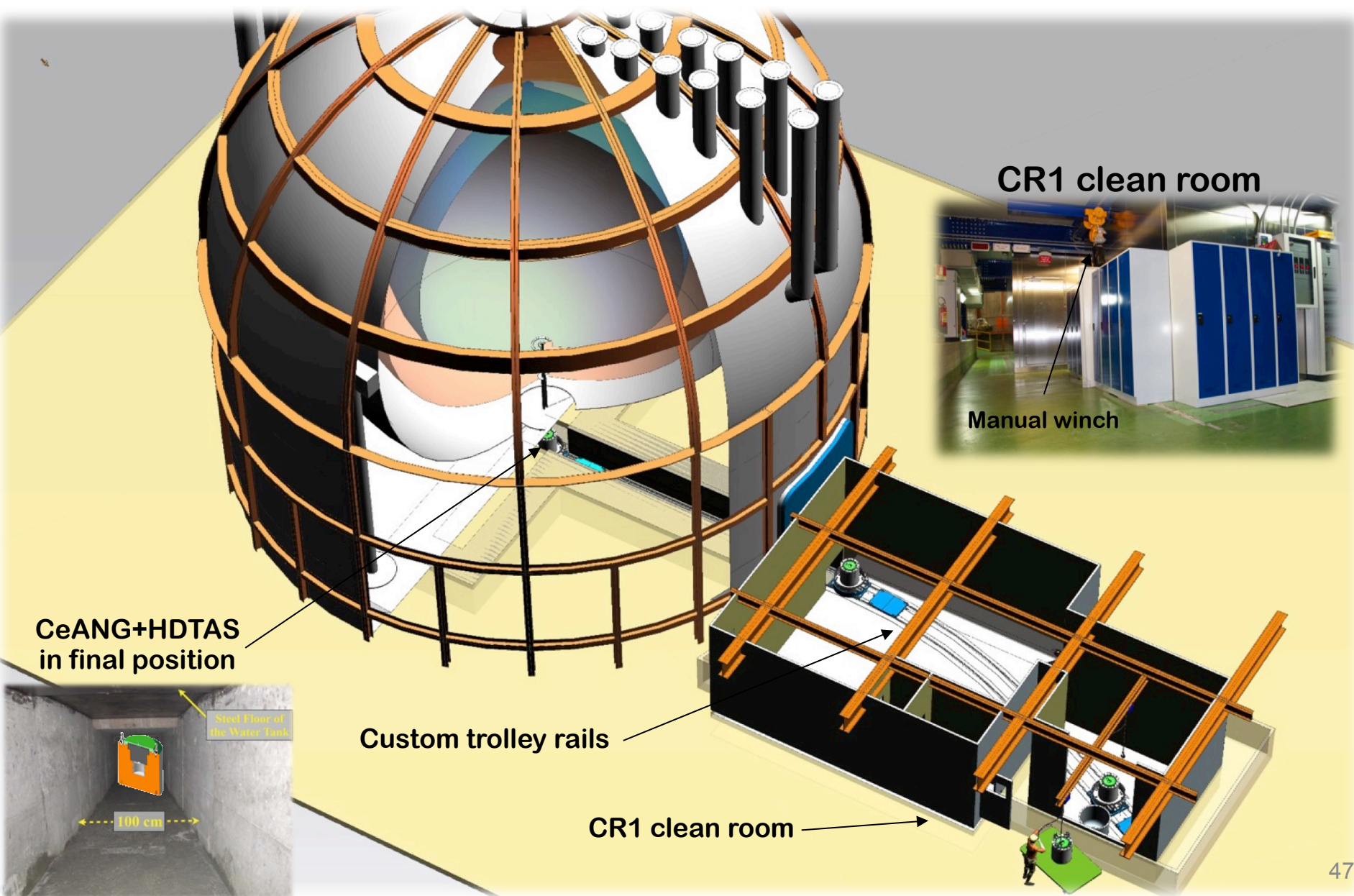


**Gran Sasso National Laboratory**

**Hall C (Opera / Borexino)**

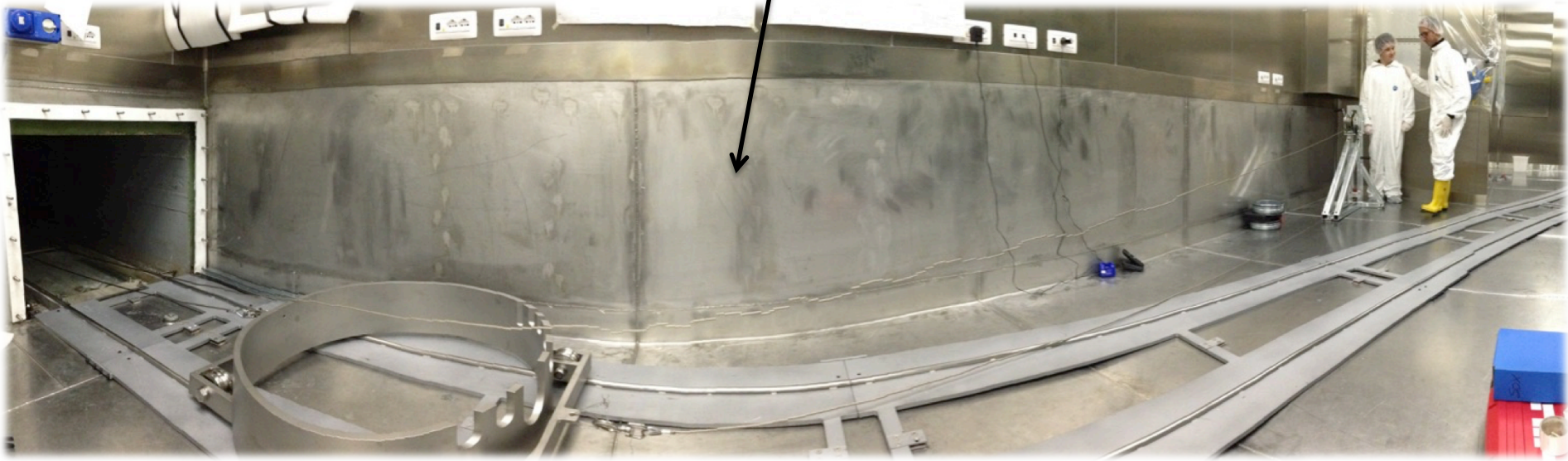
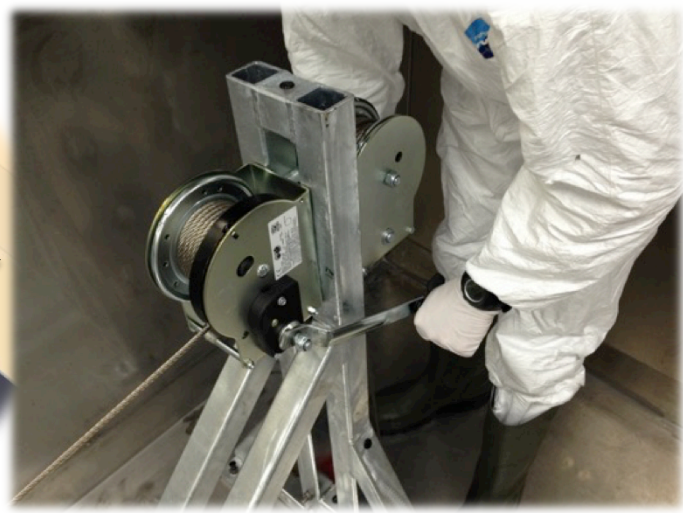
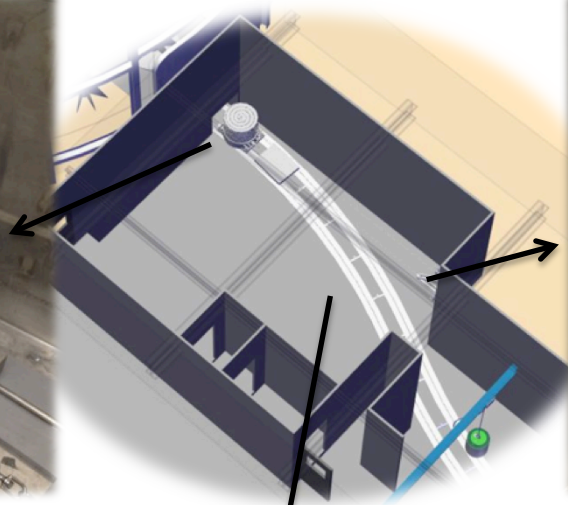


# Place CeANG next to BX





# Installation Status at LNGS

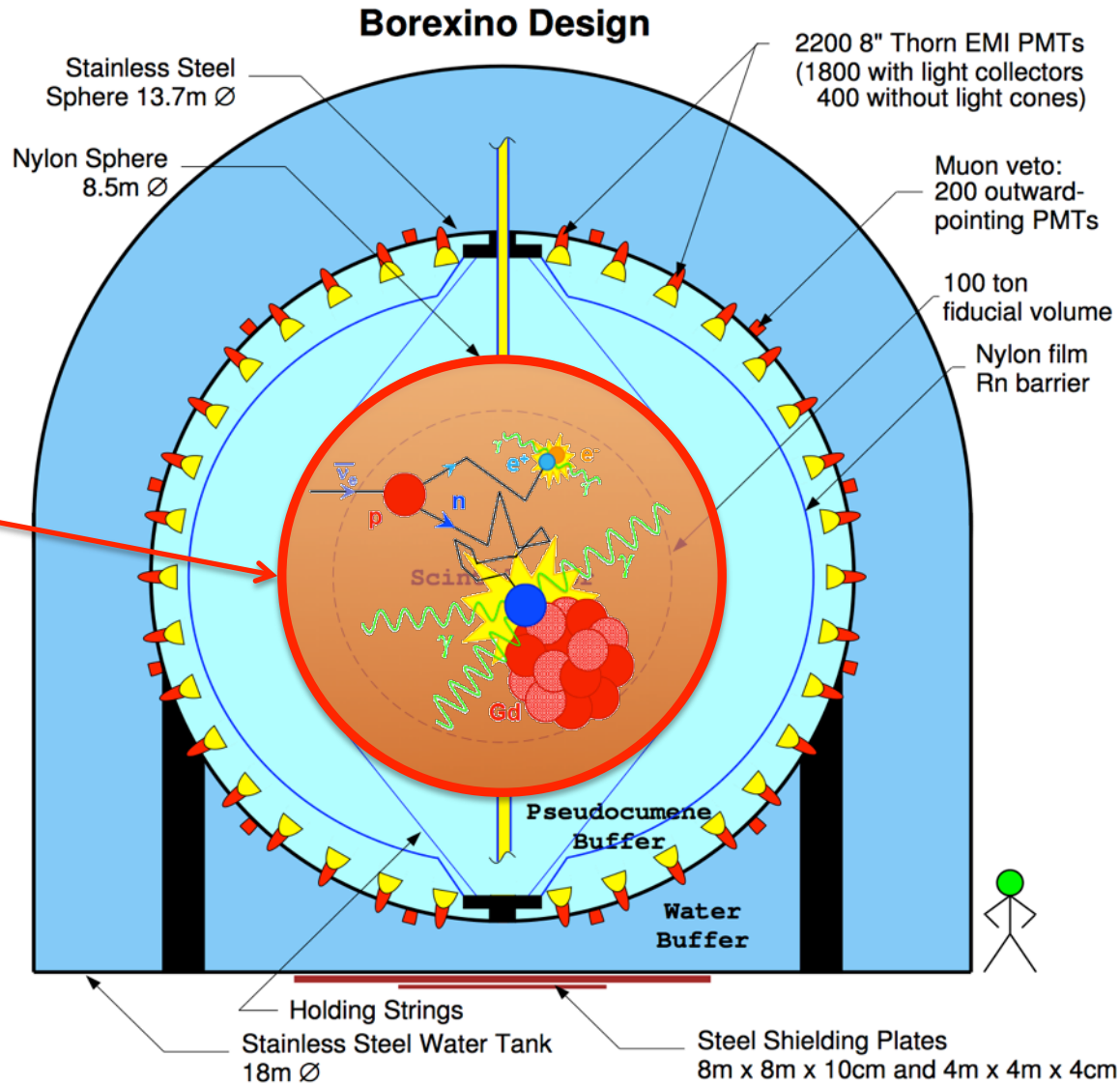




# Minimal Configuration

## CeSOX target

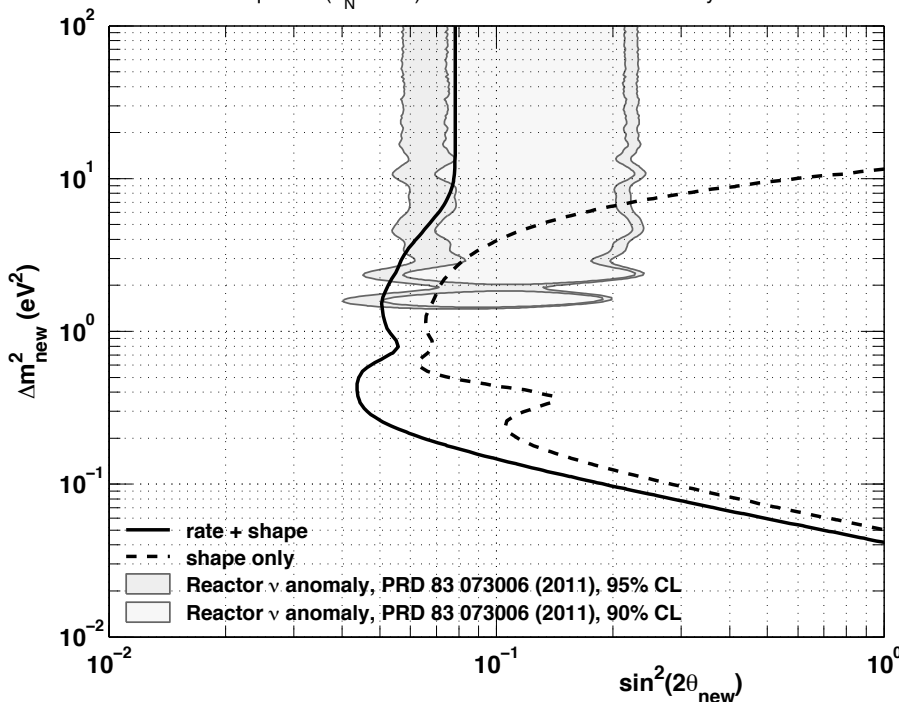
- $R < 4.25$  m
- 280 tons
- $C_6H_{12}$
- $\#H: 1.7 \cdot 10^{31}$



# Expected Sensitivity

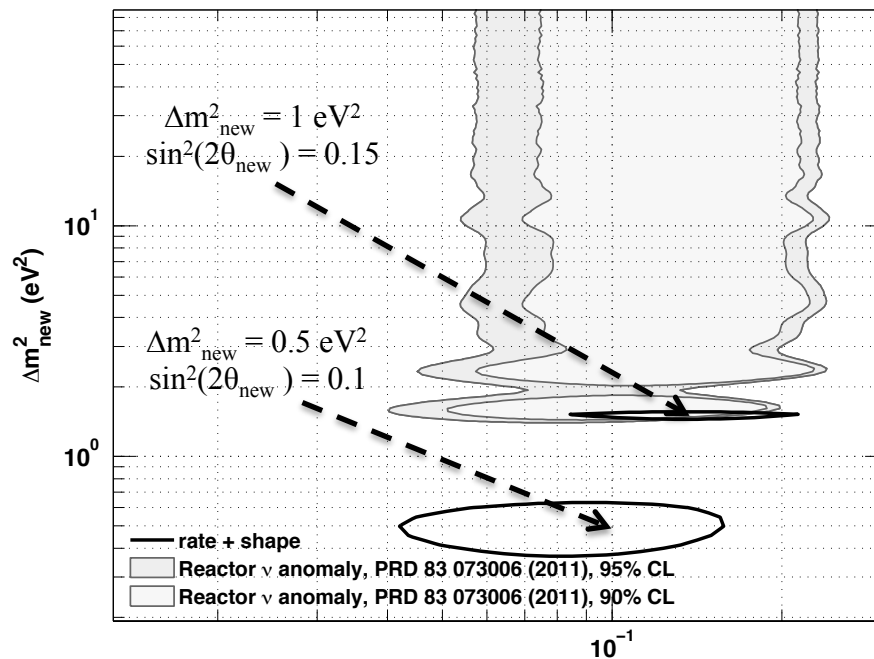
3.7 PBq (100 kCi) - 1.5 year of data taking  
 Activity measurement uncertainty: 1.5%  
 Shape only analysis (---) & Rate + Shape analysis (—)

3.7 PBq  $^{144}\text{Ce}$  ( $\sigma_N=1.5\%$ ) @ 8.2 m from Bx center - 1.5 y - 90.000 % CL



Exclusion contour (90% CL)

3.7 PBq  $^{144}\text{Ce}$  ( $\sigma_N=1.5\%$ ) @ 8.2 m from Bx center - 1.5 y - 99.000 % CL



Discovery potential (99% CL)

# Search for $\nu_s$ with ${}^3\text{H}$ $\beta$ -decay

erc

- Source:  ${}^3_1\text{H} \rightarrow {}^3_2\text{He} + e^- + \bar{\nu}_e$
- A new branch in the  $\beta$ -spectrum :

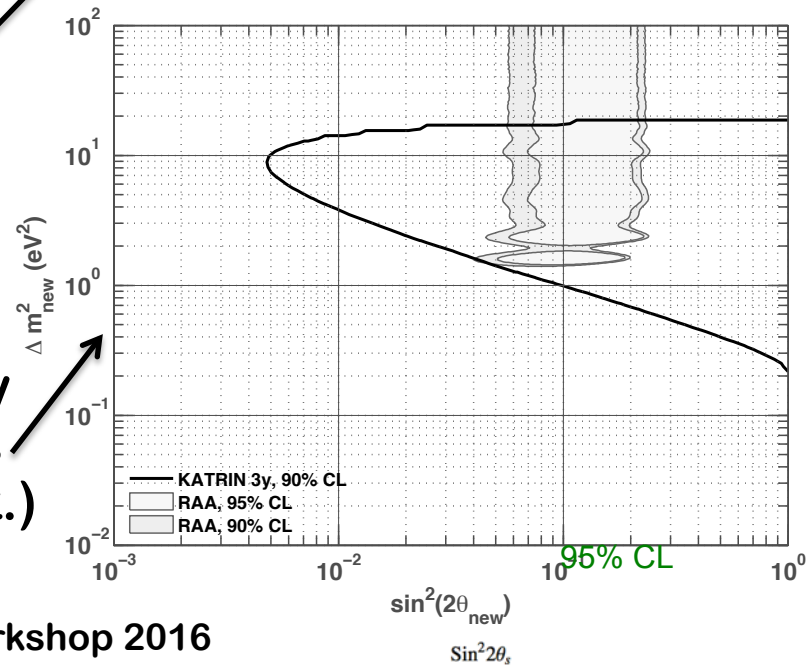
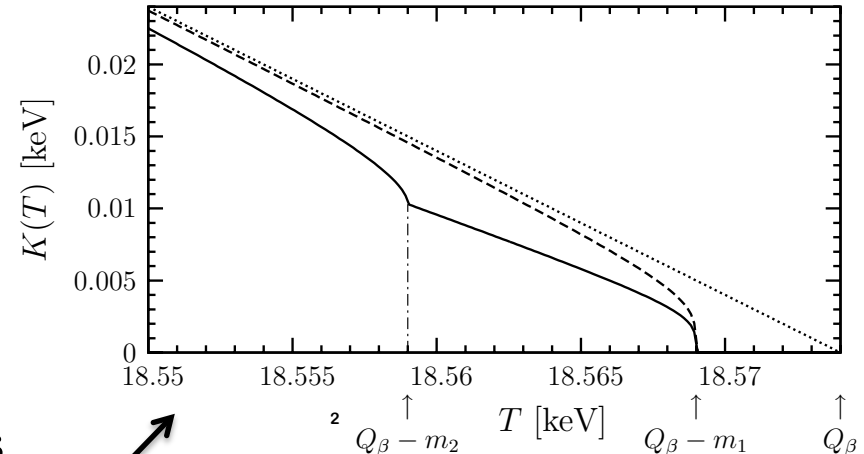
$$\langle m_\beta \rangle_4 = |U_{e4}| \sqrt{\Delta m_{41}^2}$$

- Modification of the effective mass

$$\langle m_\beta \rangle = \sqrt{\sum_{1,2,3,\dots} |U_{ei}|^2 m_i^2}$$

→ Search for a –kink–  
a few eV below end point

- KATRIN –as designed– can partially test the  $\nu_e$  disappearance anomalies (sensitivity to be assessed with syst.)



## @ Neutrino Beam



**Test of LSND/MinibooNE/reactor/gallium anomalies**  
**If positive signal, detailed study of sterile- $\nu$  phenomenology**

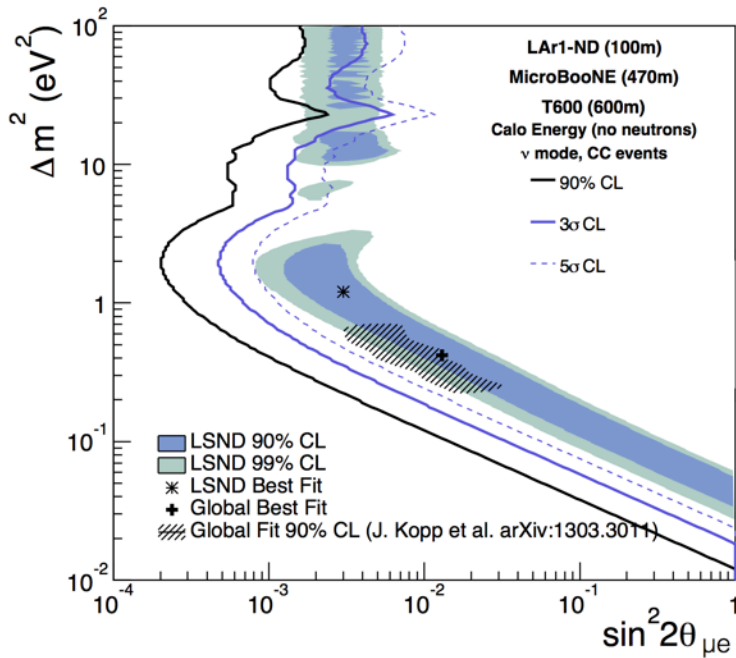
# $\nu$ Beam Proposals

Type	Source	App. /Dis.	Oscillation Channels	Projects
<b>Isotope Decay at Rest</b>	$p + {}^9\text{Be} \rightarrow {}^8\text{Li} + 2p$ $n + {}^7\text{Li} \rightarrow {}^8\text{Li}$ ${}^8\text{Li} \rightarrow {}^9\text{Be} + e^- + \bar{\nu}_e$	Dis.	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	IsoDAR
<b>Pion (Kaon) Decay at Rest</b>	$\pi^+ \rightarrow \mu^+ \nu_\mu$ $\quad \quad \quad \searrow$ $\quad \quad \quad e^+ \bar{\nu}_\mu \nu_e$	App. & Dis.	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	OscSNS, KDAR, JPARC-MLF
<b>Pion Decay in Flight</b>	$\pi^+ \rightarrow \mu^+ \nu_\mu$ $\quad \quad \quad \searrow$ $\quad \quad \quad e^+ \bar{\nu}_\mu \nu_e$	App. & Dis.	$\nu_\mu \rightarrow \nu_e$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_\mu \rightarrow \nu_\mu$ $\nu_e \rightarrow \nu_e$	MINOS+, MicroBooNE, LAr1kton Icarus/Nessie
<b>Low-E Neutrino Factory</b>	$\mu^+ \rightarrow e^+ \bar{\nu}_\mu \nu_e$ $\mu^- \rightarrow e^- \nu_\mu \bar{\nu}_e$	App. & Dis.	$\nu_e \rightarrow \nu_\mu$ $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$ $\nu_\mu \rightarrow \nu_\mu$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$	$\nu$ STORM



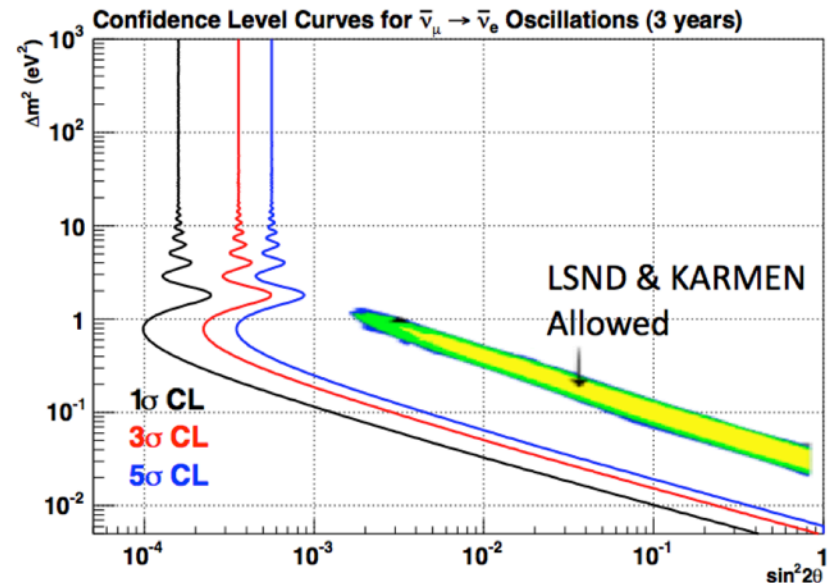
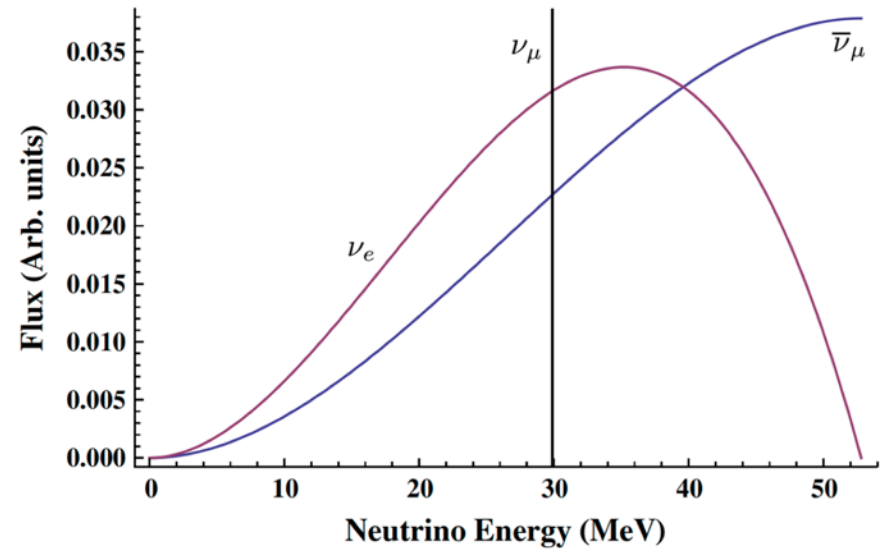
# Pion Decay in Flight $\nu$ -sources

- ICARUS (770t LAr, 600 m)
  - rebuilt @CERN (WA104)
  - Data taking at FNAL
- New near detector (82t, 100m)
- LAr1-ND  $\rightarrow$  T-1053 R&D
- MicroBooNE (470 m)



# Pion Decay at Rest $\nu$ -sources

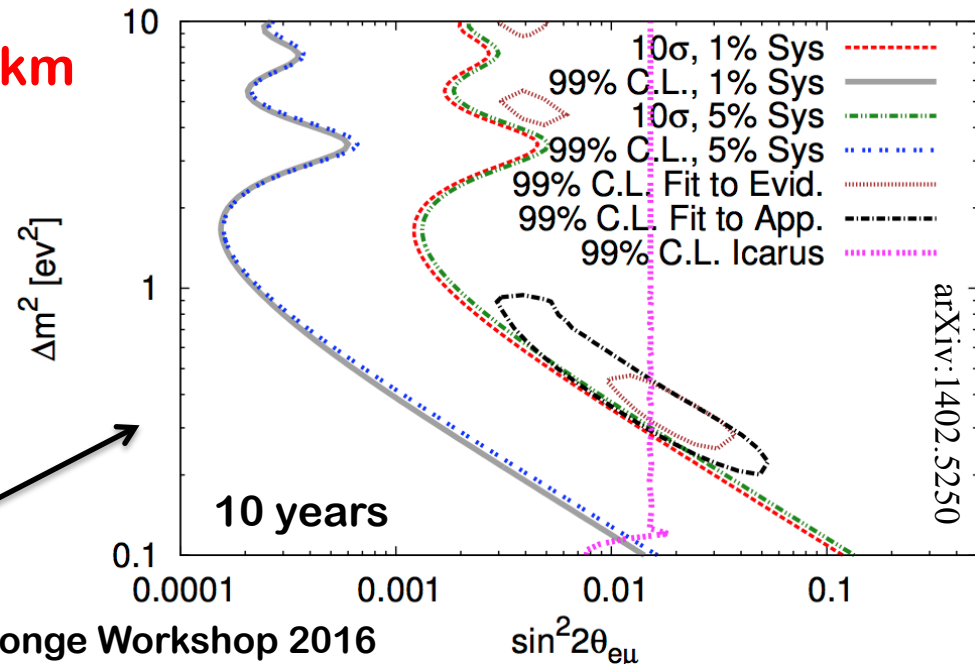
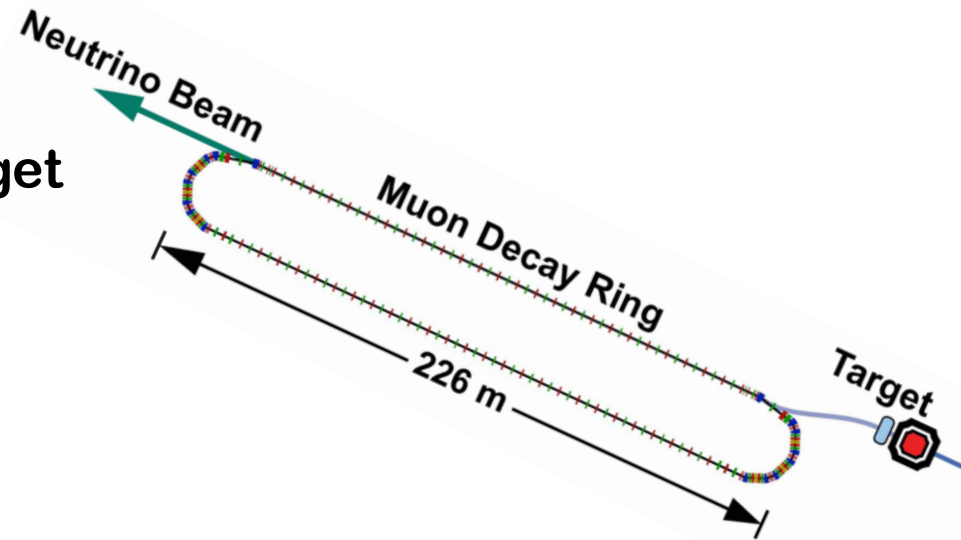
- **High Energy Proton source**
  - Each  $\pi^+$  decay
    - $\nu_\mu, \nu_e, \bar{\nu}_\mu$
    - well known E spectrum
- **Detection channels**
  - $\nu_e \rightarrow \nu_e$  Disappearance
  - $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  Appearance
- **Direct Test of LSND**
- **OscSNS (ORNL, 1.4 MW)**
  - 800 t LS-det @ 60 m
- **JPARC-MLF**
  - 2x25ton Gd-LS-det @17 m





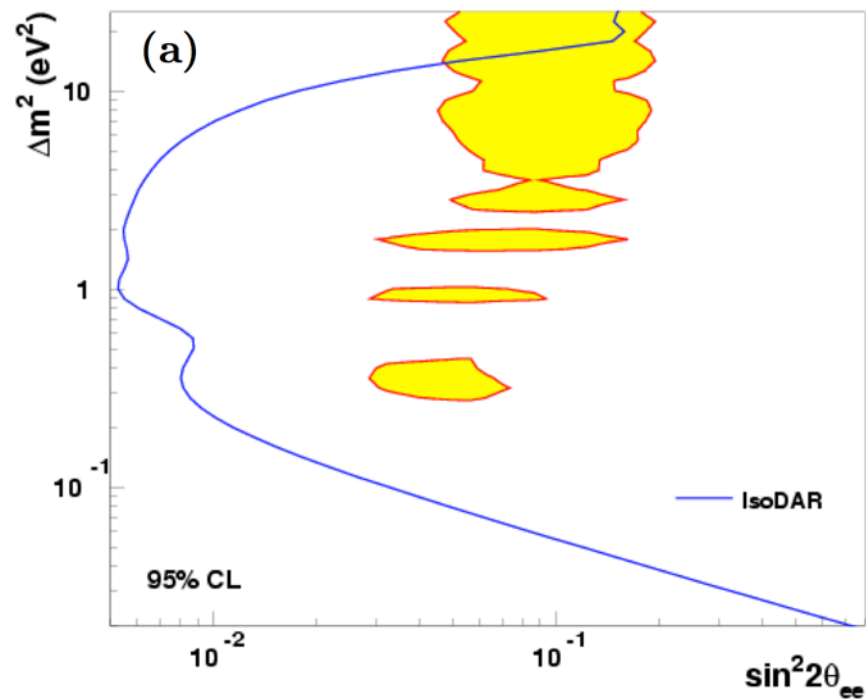
# Muon Decay Rings: $\nu$ -STORM

- **Neutrino Factory Concept**
  - 60 GeV protons on solid target
  - Horn capture and  $\pi$  transfer
  - Low-E muon Decay ring
- **APP and DIS channels with:**
  - $(\bar{\nu})_{\mu}, (\bar{\nu})_e$
- **kT-scale Fe+PS detector @ 2km**
  - Magnetized to tag wrong charged muons
- **Golden Mode**
  - $(\bar{\nu})_{\mu}$  APP in a  $(\bar{\nu})_e$  beam
- **Ultimate sterile  $\nu$  search**

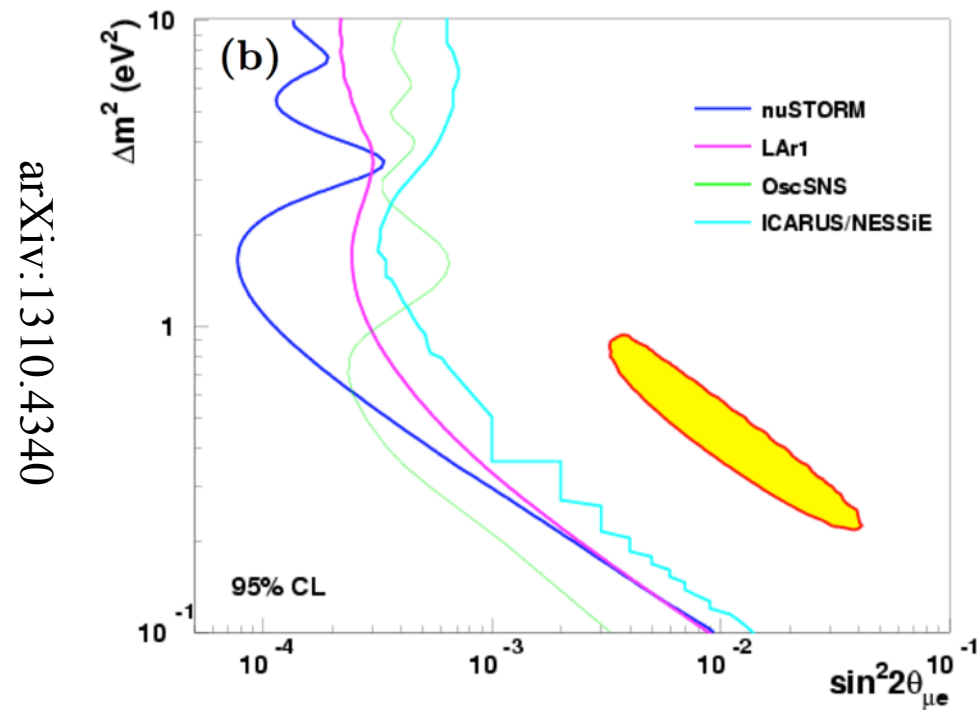


# Beam Experiment Sensitivities

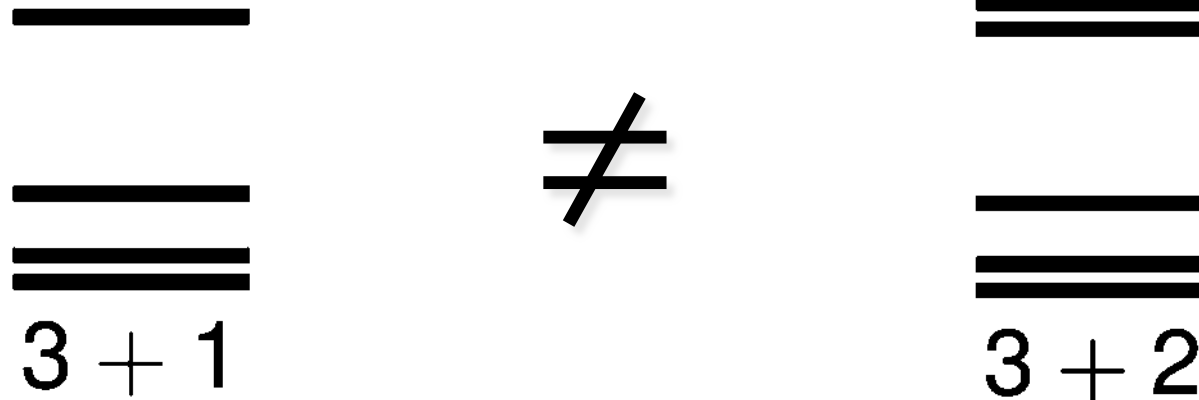
## Disappearance



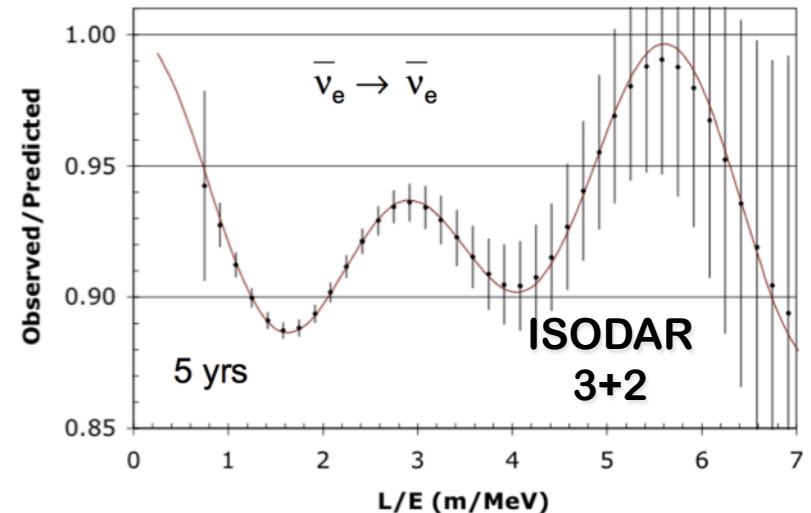
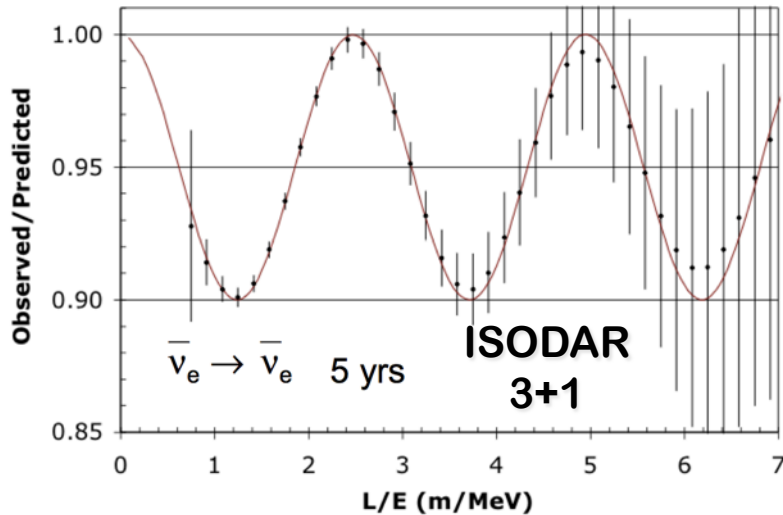
## Appearance



# Isotope Decay at Rest $\nu$ -sources



$\bar{\nu}_e$  disappearance: L/E Waves with  $1e6 \nu / 5y @ 16m$



- **2.7 – 3.8  $\sigma$  anomalies (each) calling for clarification**
  - LSND & MiniBooNE?
  - Gallium Anomaly
  - Reactor Anomaly
- $\Delta m^2 \approx eV^2$  Sterile Neutrino? Or Experimental Artifacts?
- **But also negative indications:**
  - No deficit in  $\Delta m^2 \approx eV^2$  muon disappearance exp.
  - Tensions in global fits (APP vs DIS)
- **A definitive probe of this parameter space is necessary → need several experiments**

- Many proposals with capabilities to unambiguously test  $L/E \approx 1$  m/MeV oscillatory behavior with low systematics
- **Reactor Neutrinos**
  - Results within 5 years, Cost :1-10 M\$
  - Challenge: Background mitigation
- **Neutrino Generator**
  - Results within 5 years, cost  $\approx 5$  M\$
  - Challenge: the isotope production and transportation
- **Neutrino 'Beam'**
  - Middle or long terms, cost 10-300 M\$
  - Would allow studying sterile neutrino phenomenology
  - Relevant for X-section study and R&D for next generation projects
- **Other tests through  $\beta$ -decay and  $(\beta\beta)0\nu$ -decay, Cosmo**