

Dr. Adam Riess

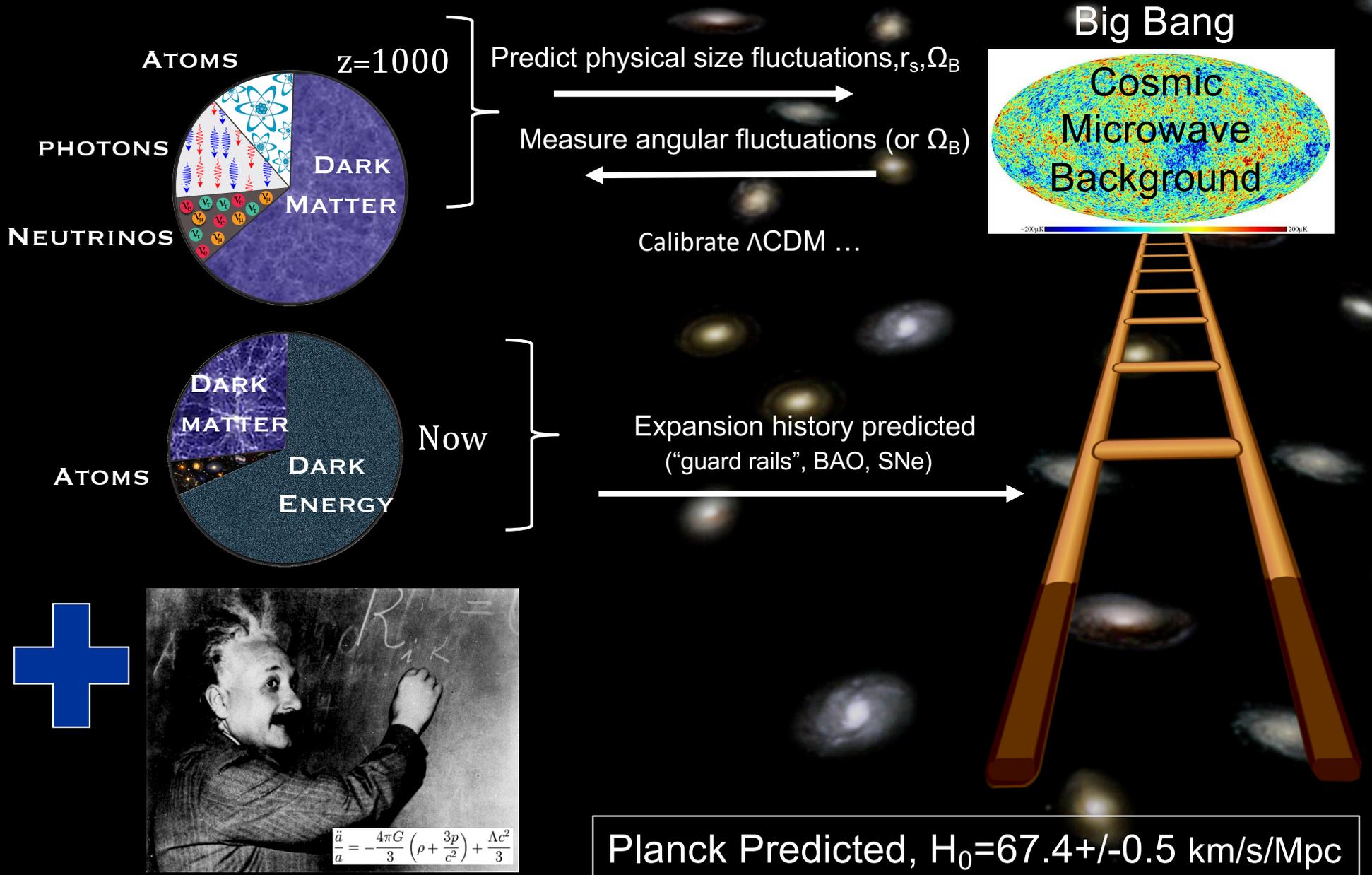
Johns Hopkins University  
Space Telescope Science Institute

**NEW MEASUREMENTS OF THE  
EXPANSION RATE OF THE UNIVERSE,  
FASTER THAN WE THOUGHT**

SH<sub>0</sub>ES Team

# Ultimate "End-to-end" test for $\Lambda$ CDM, Predict and Measure $H_0$

Standard Model: (Vanilla)  $\Lambda$ CDM, 6 parameters + ansatz ( $w$ ,  $N_{\text{eff}}$ ,  $\Omega_K$ , etc)



# A Direct, Local Measurement of $H_0$ to percent precision

## The $SH_0ES$ Project (2005)

(Supernovae,  $H_0$  for the dark energy Equation of State)

A. Riess, L. Macri, D. Scolnic, S. Casertano, A. Filippenko, W. Yuan, S. Hoffman, +

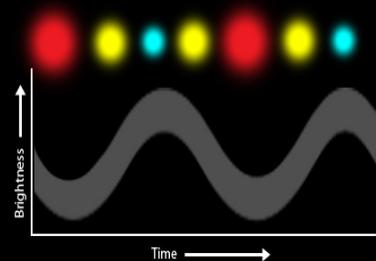
Measure  $H_0$  to percent precision empirically by:

- A strong, simple ladder: **Geometry  $\rightarrow$  Cepheids  $\rightarrow$  SNe Ia**

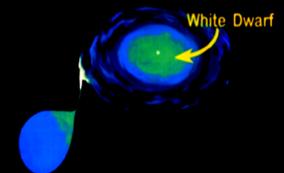
Multiple ways



Pulsating Stars,  
 $10^5 L_{\odot}$ , P-L relation



Exploding Stars,  
 $10^9 L_{\odot}$ ,  $\sigma \sim 5\%$



An explosion resulting from the thermonuclear detonation of a White Dwarf Star.

- Reduce systematics w/ consistent data along ladder and NIR
- Thorough propagation of statistical and systematic
- HST Cycle 11-28, 17 competed GO proposals,  $\sim 1000$  orbits

# A Direct, Local Measurement of $H_0$ to percent precision

## The $SH_0ES$ Project (2005)

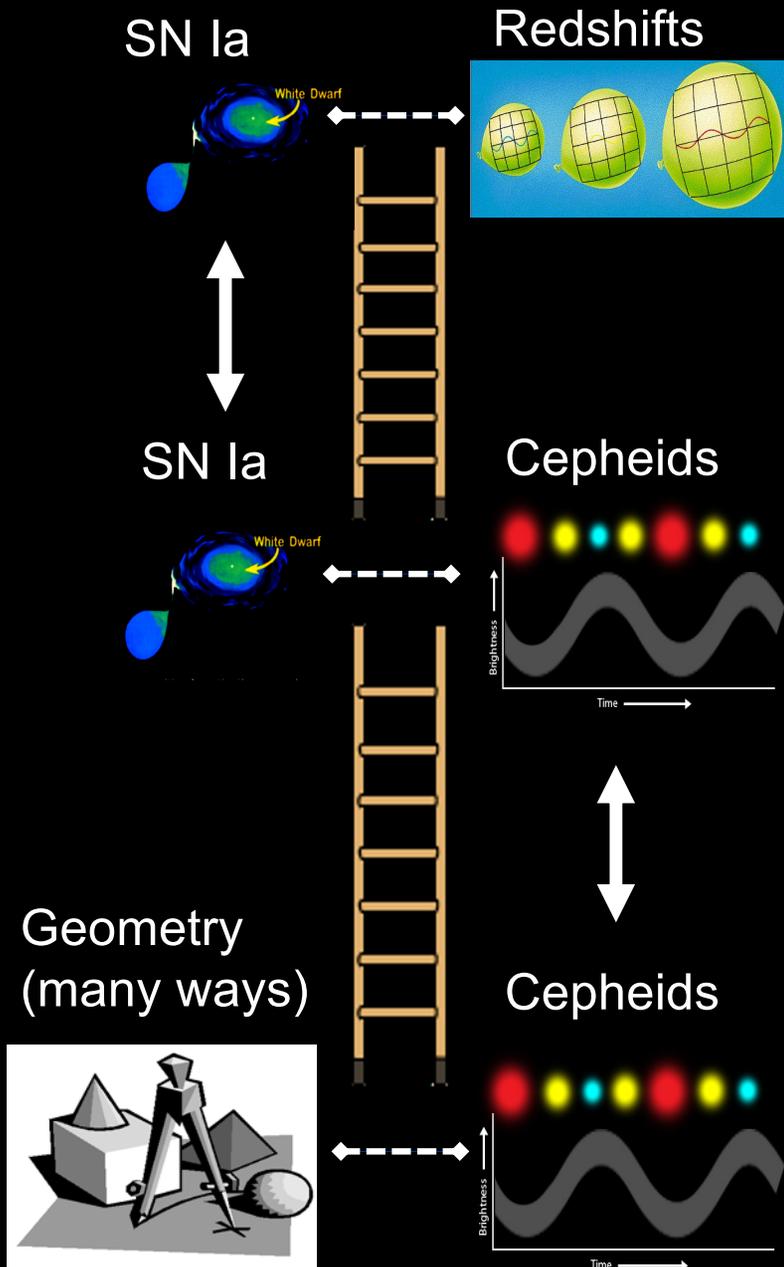
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Measure  $H_0$  to percent precision empirically by:

- A strong, simple ladder: **Geometry → Cepheids → SNe Ia**
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# Distance Ladders: Simple & Empirical, Must be Consistent



**Hubble Flow:**  
 $D \sim \text{Gpc}, z \sim 0.1$

**Cross-calibrate:**  
 $D \sim 10\text{-}40 \text{ Mpc}$

**Anchors:**  
 $D \sim \text{Kpc or Mpc}$

## Nutrition Facts

Serving size 1 potato (148g/5.2oz)

Amount per serving  
**Calories** **73**

% Daily Value\*

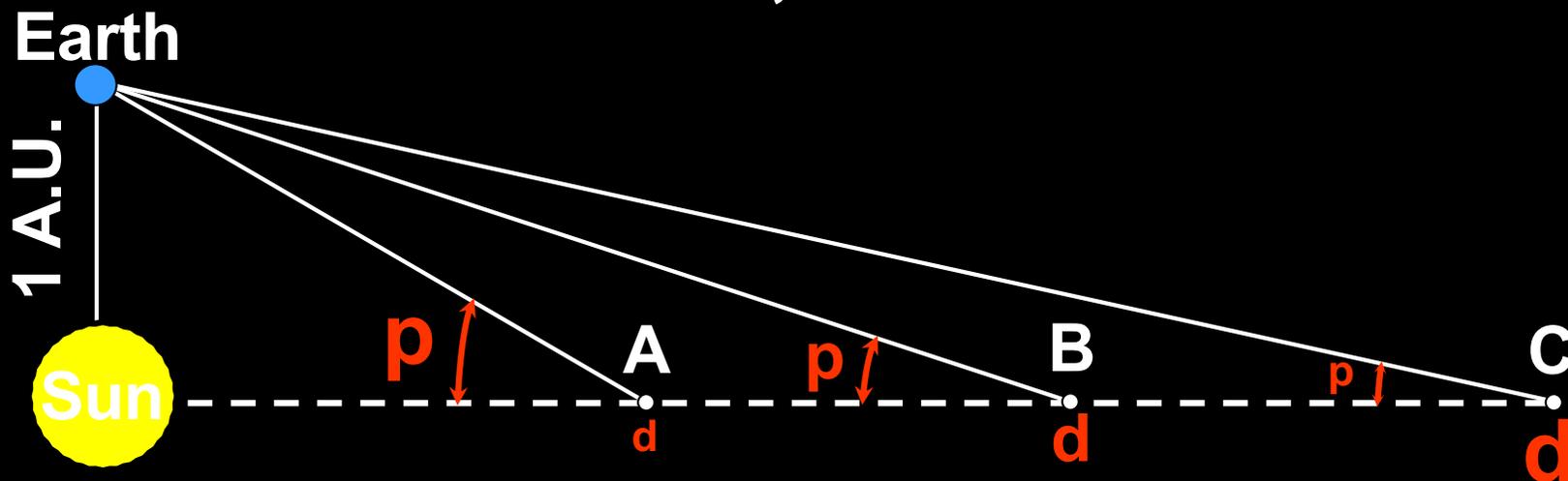
|                        |     |
|------------------------|-----|
| Astrophysical modeling | 0%  |
| General Relativity     | <1% |
| LCDM                   | <1% |

## ⚠ WARNING

Same object types on different rungs must be standardized and measured consistently!

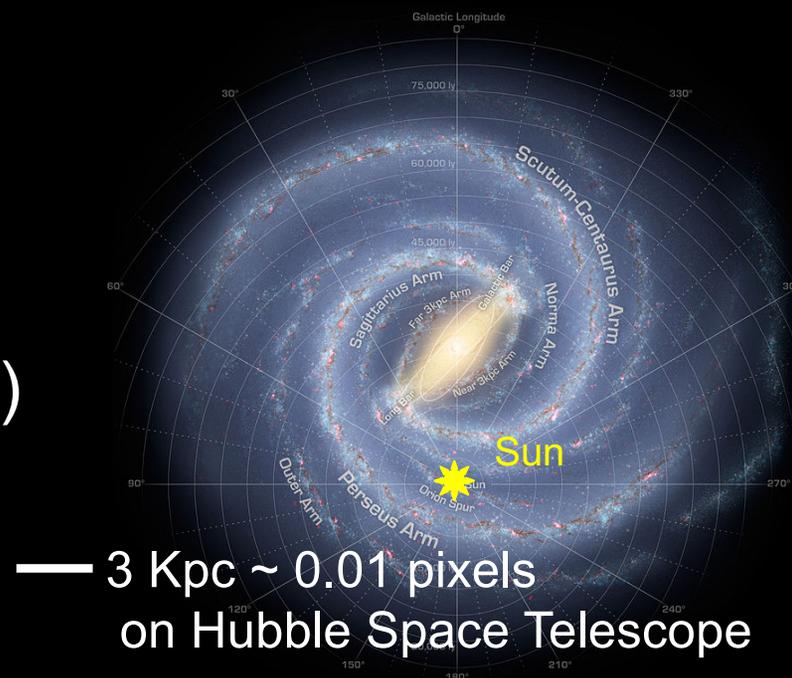
# Parallax in the Milky Way at Kiloparsec Distances

**Stars are far, Parallax is small !**

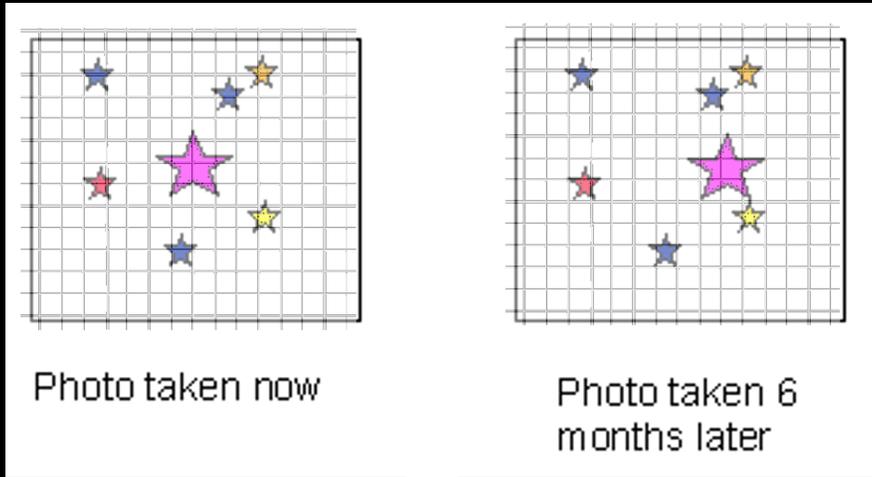


$$d \text{ (kpc)} = \frac{1}{p \text{ (milliarcsec)}}$$

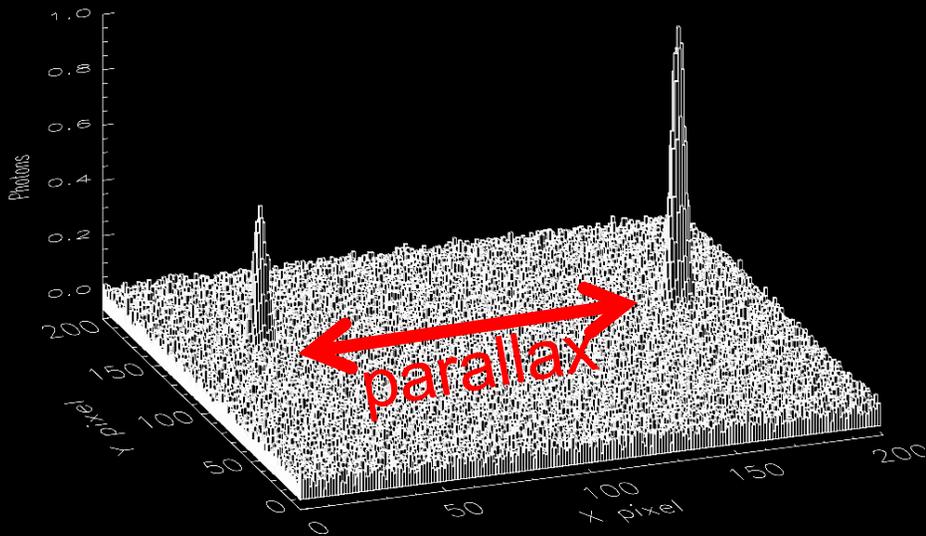
Nearly all long-period ( $P > 10$  days)  
MW Cepheids  $D > \text{kpc}$



# Extending Parallax with WFC3 Spatial Scanning

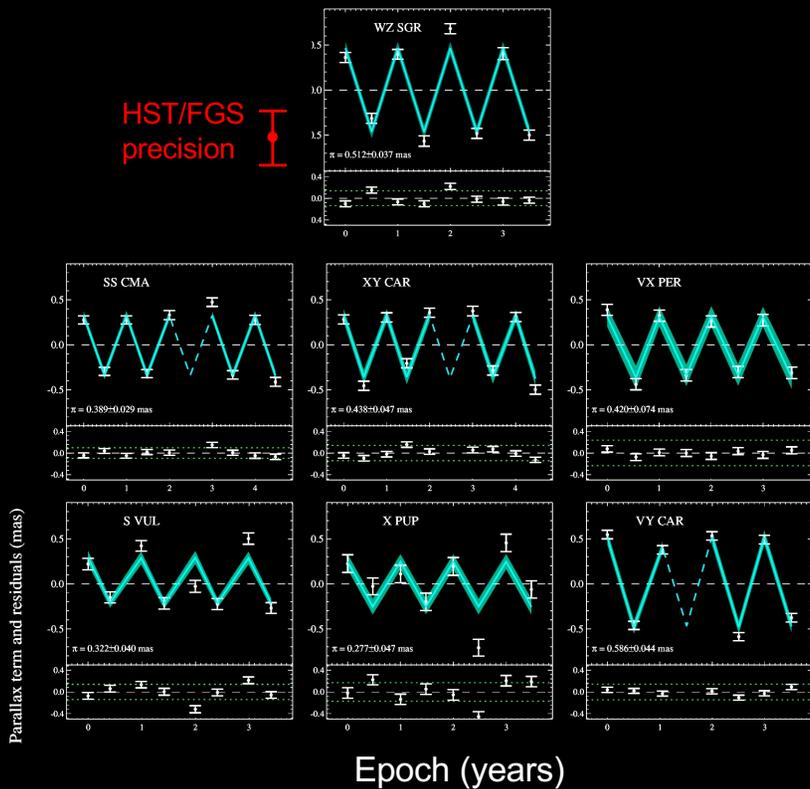


Imaging, precision=0.01 pix  $\sigma$   
WFC3:  $\sim 1\sigma$  @ 3 kpc



# New Tool: WFC3 Spatial scanning for long range parallaxes, photometry

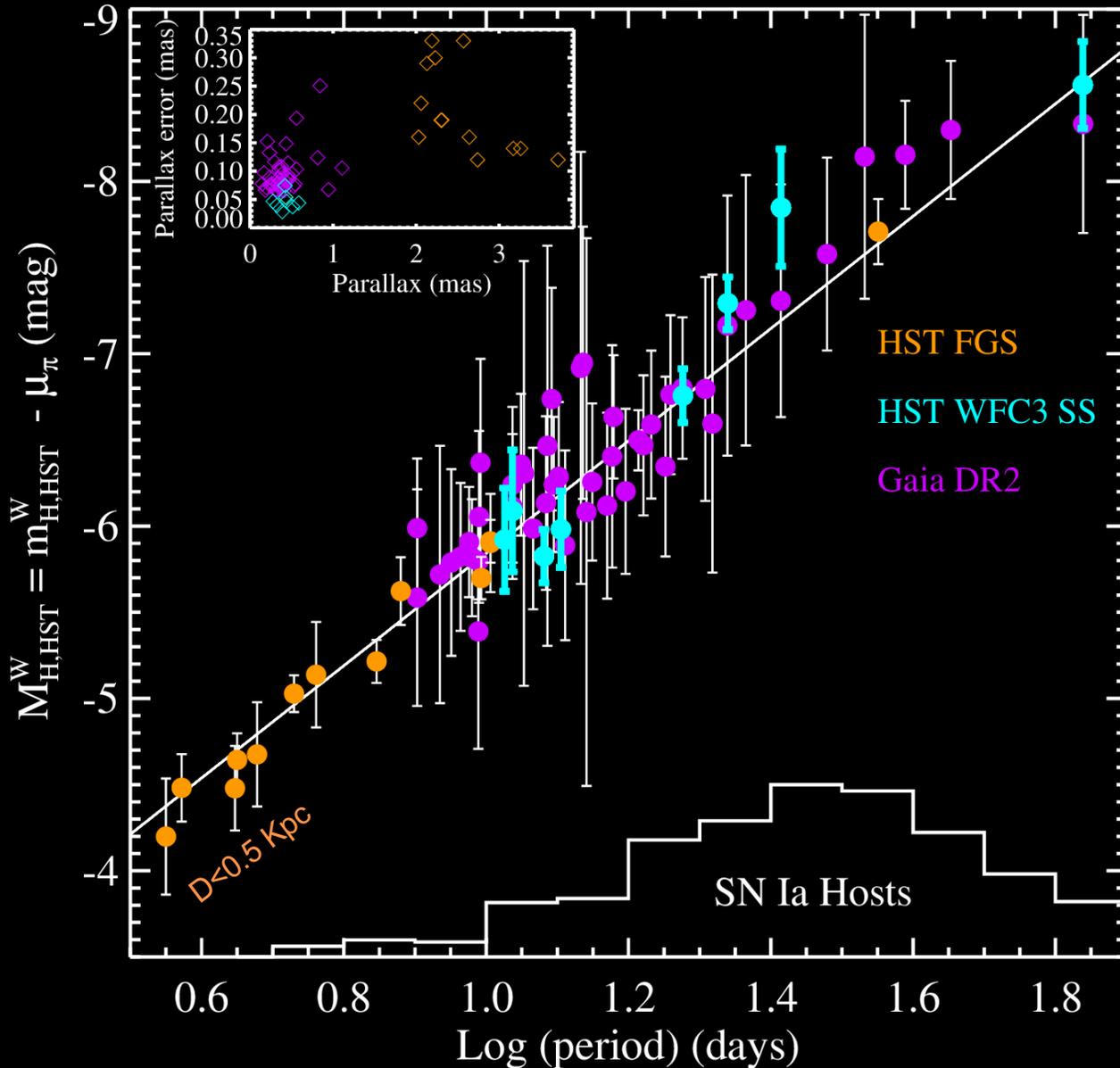
WFC3 Spatial Scanning  $\rightarrow$  20-40  $\mu\text{as}$   
4 Years Later: Proper Motion subtracted,  
8 MW long-P Cepheid Parallaxes  
1.7 < D < 3.6 Kpc, error in mean = 3.3%



Riess et al. (2018a), ApJ, 855, 136

# Milky Way Cepheid P-L Relation, Now w/ HST photometry, Long Periods

## Milky Way PL Relation



Final Gaia Parallaxes  
+ HST Photometry  $\rightarrow$   
 $H_0 \sim 0.4\%$

}  
Periods > 10 days  
matching  
Cepheids HST sees  
in SN Ia hosts

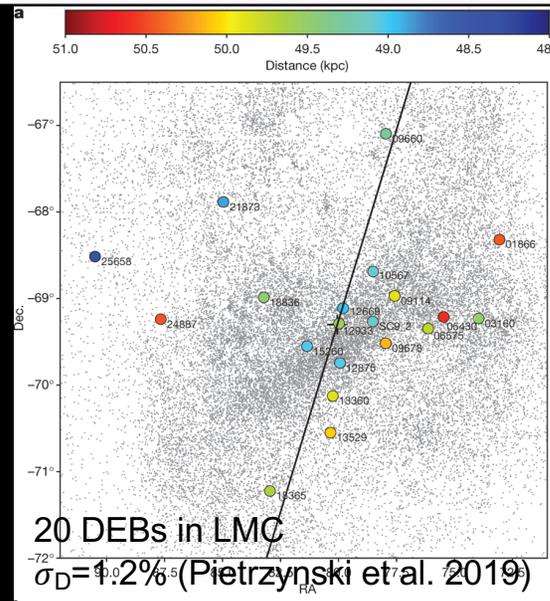
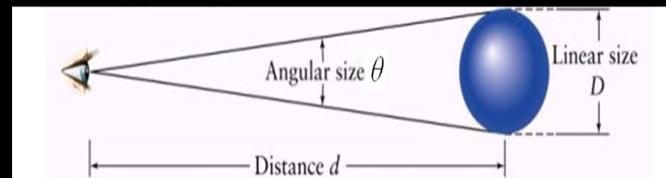
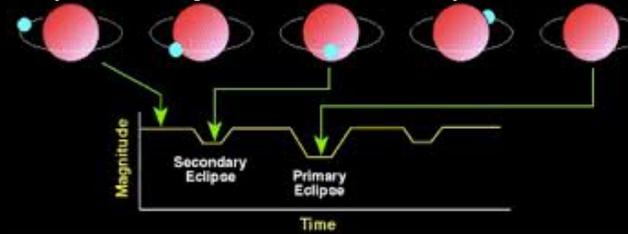
Gaia EDR3! In prep

# Three Sources of Geometric Distances to Calibrate Cepheids

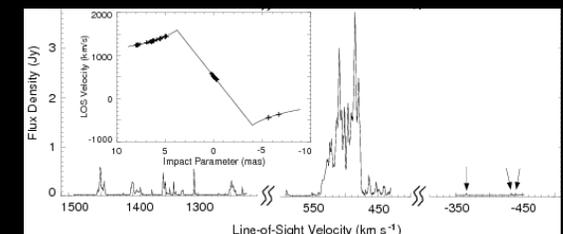
Parallax in Milky Way (WFC3 SS, HST FGS, Gaia)



Detached Eclipsing Binaries in LMC (Pietrzynski+2019)



Masers in NGC 4258, Keplerian Motion (Reid+2019)



# Step 2: Cepheids to Type Ia Supernovae

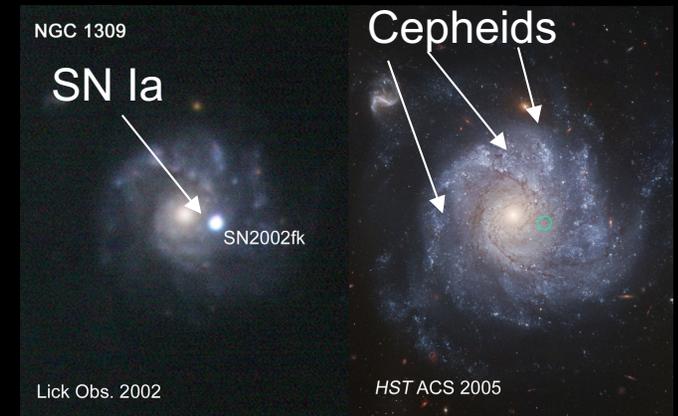
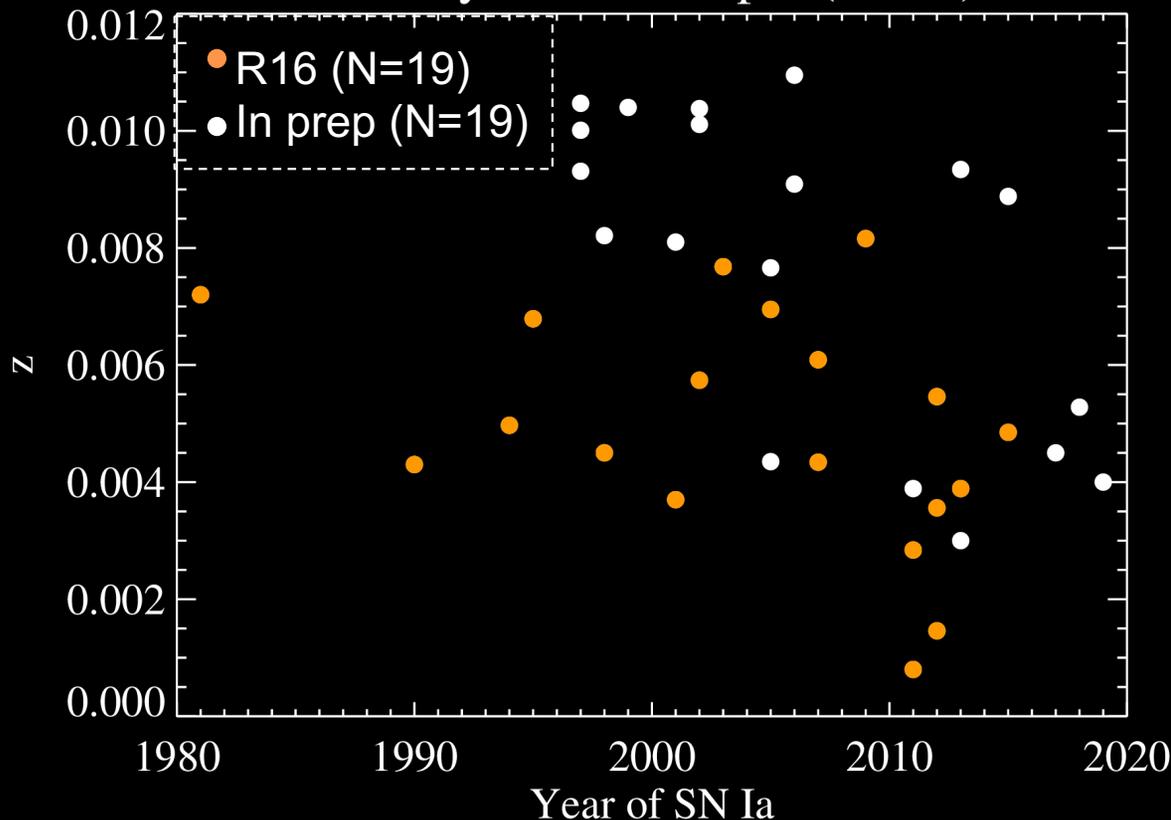
Number nearby SN Ia limits  $H_0$  precision,  $\sigma = \frac{6\%}{\sqrt{N}}$

SN Ia Requirements:  $A_V < 0.5$ , normal, pre-max, digital

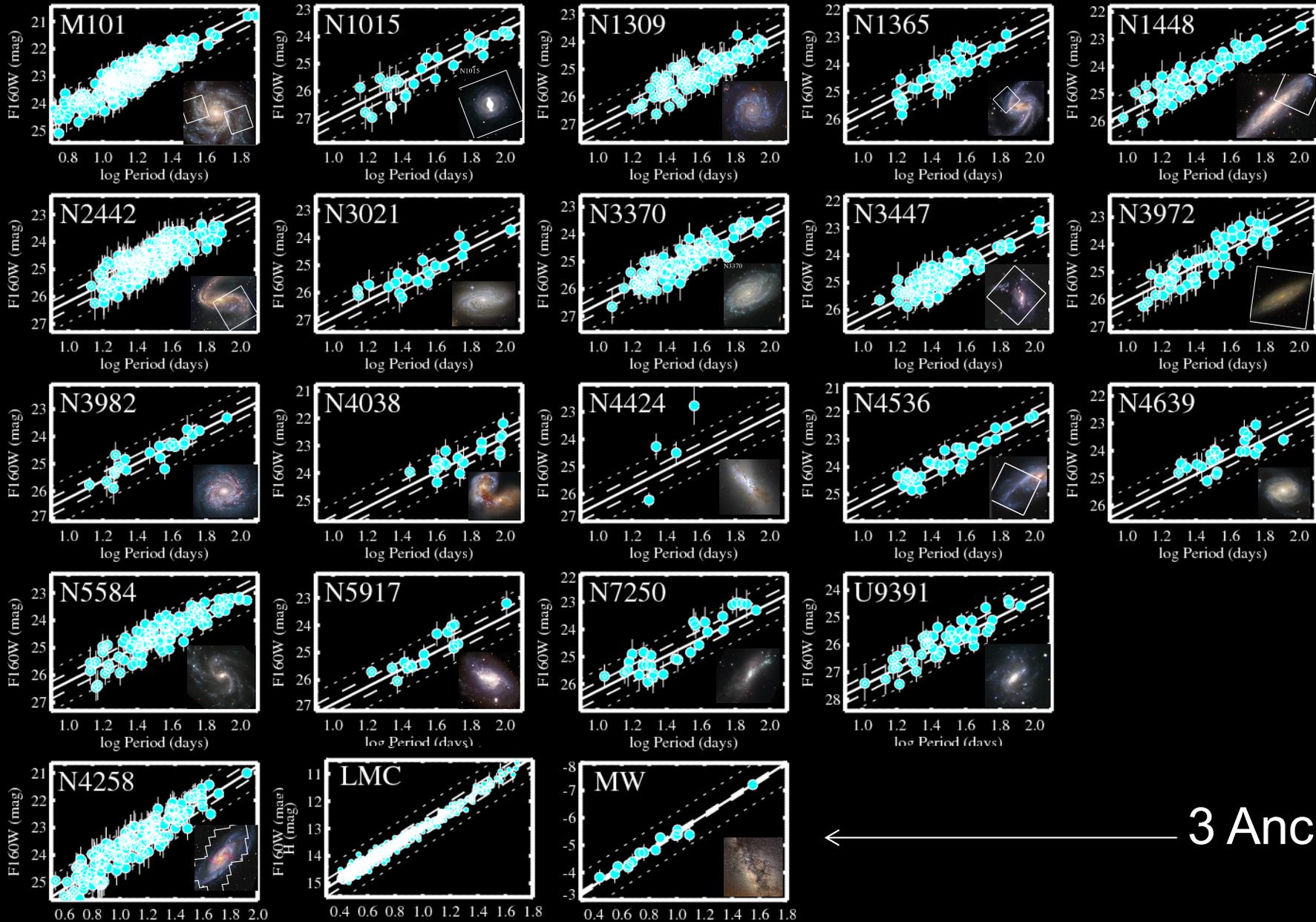
Host Requirements: Late-type,  $z \leq 0.01$ , not-edge on

2020 Complete sample (new ones @ 1.5/yr)

Nearby SN Ia Sample (N=38)



# Cepheid V,I,H band Period-Luminosity Relationships: 19 hosts, 3 anchors



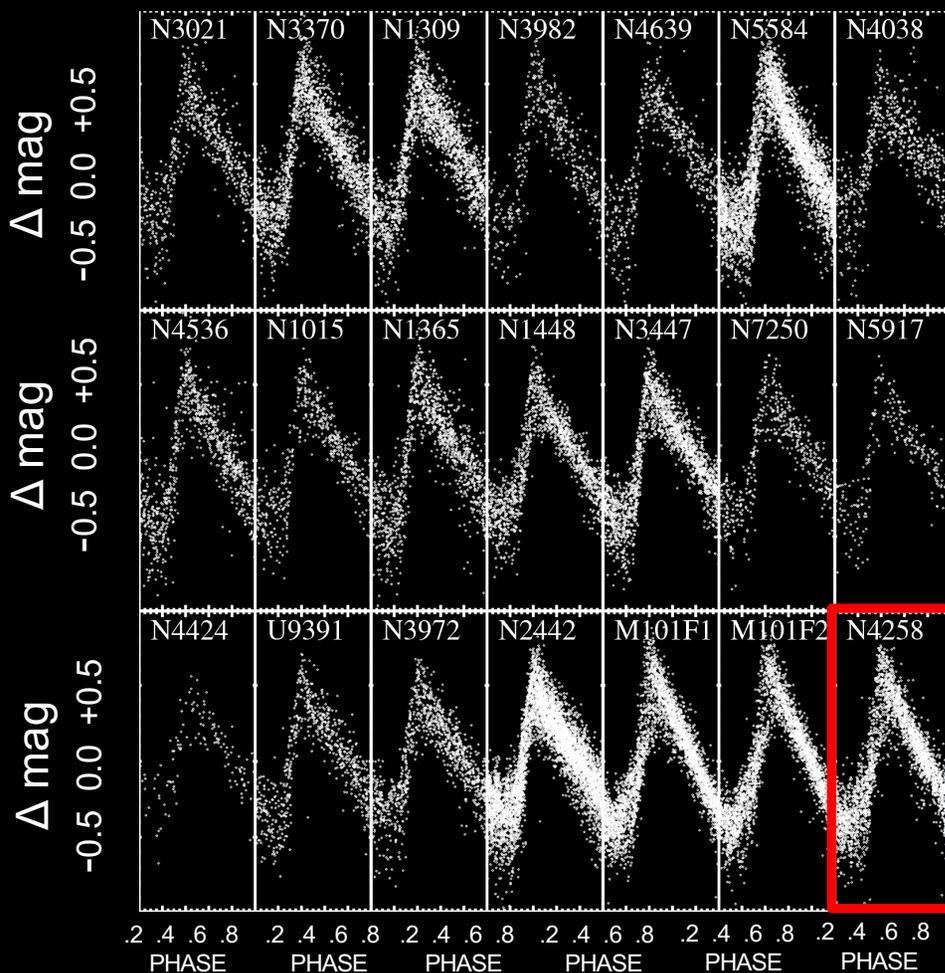
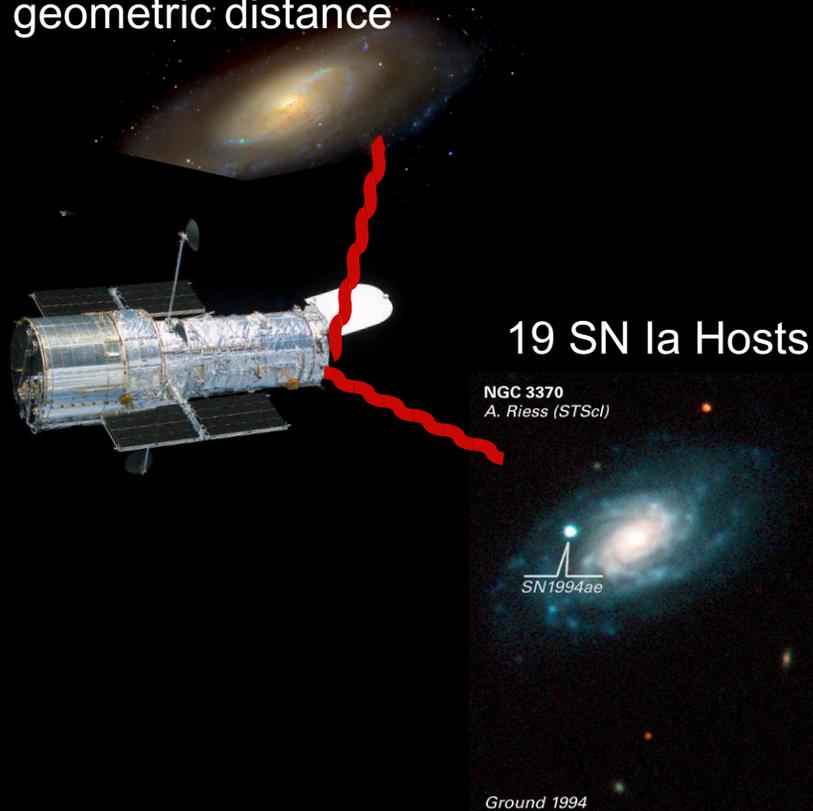
← 3 Anchors

# Lower Systematics from *Differential* Flux Measurements

To reduce systematic errors: measure all Cepheids with same instrument, filters, similar metallicity, period range

ANCHORS: NGC 4258, MW, & LMC  
geometric distance

Cepheid composite LC's, >2400



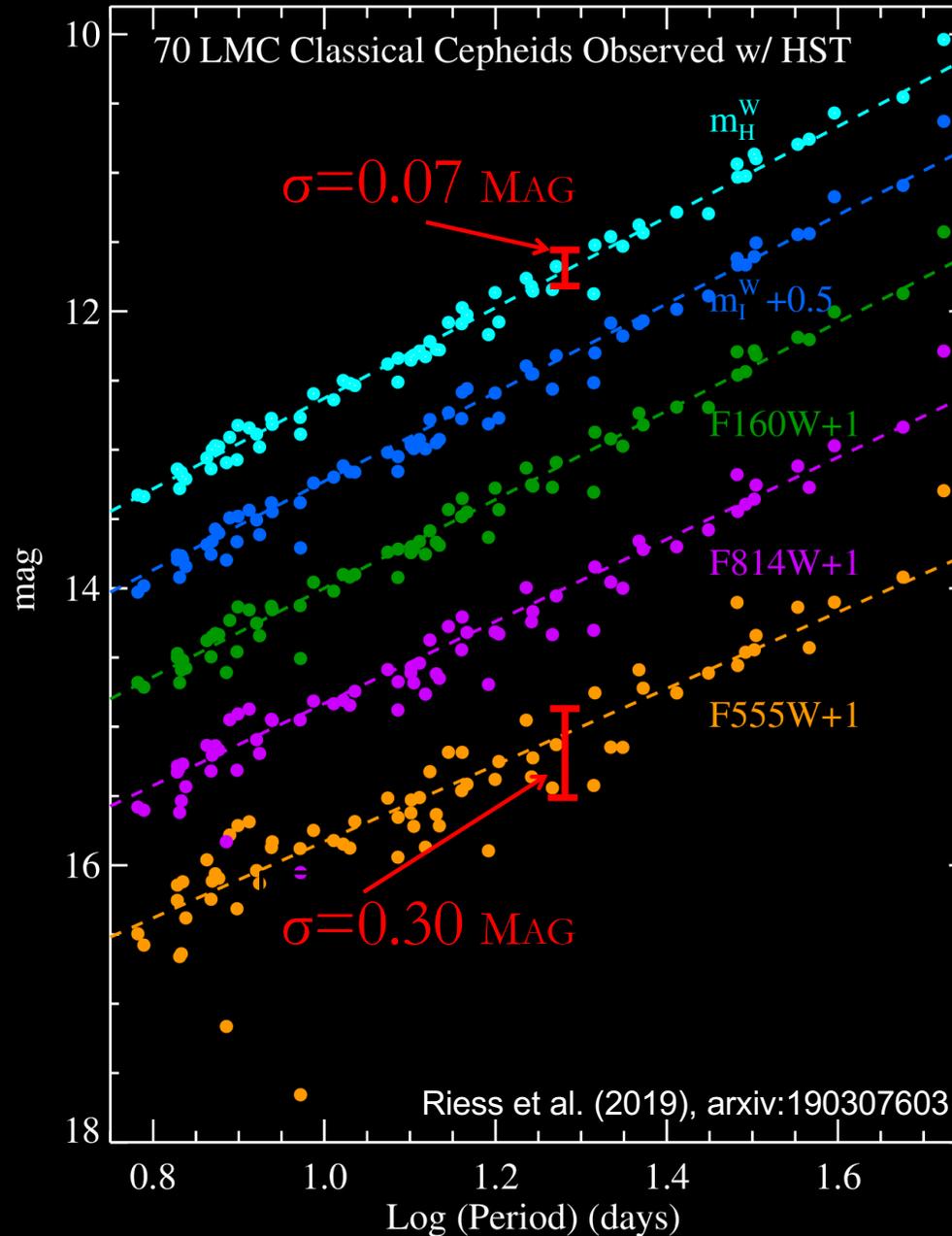
# Lowering Systematics: Near-IR Cepheid Observations + HST, Now in LMC!

-Negligible sensitivity to metallicity in NIR (F160W)

-Dependence on reddening laws 6x smaller than optical

We use F160W-band as primary +F555W,F814W

Key Project used F555W and F814W



Dereddened:  
F160W-0.386(F555W-F814W)

1.6 $\mu$ m

0.8 $\mu$ m

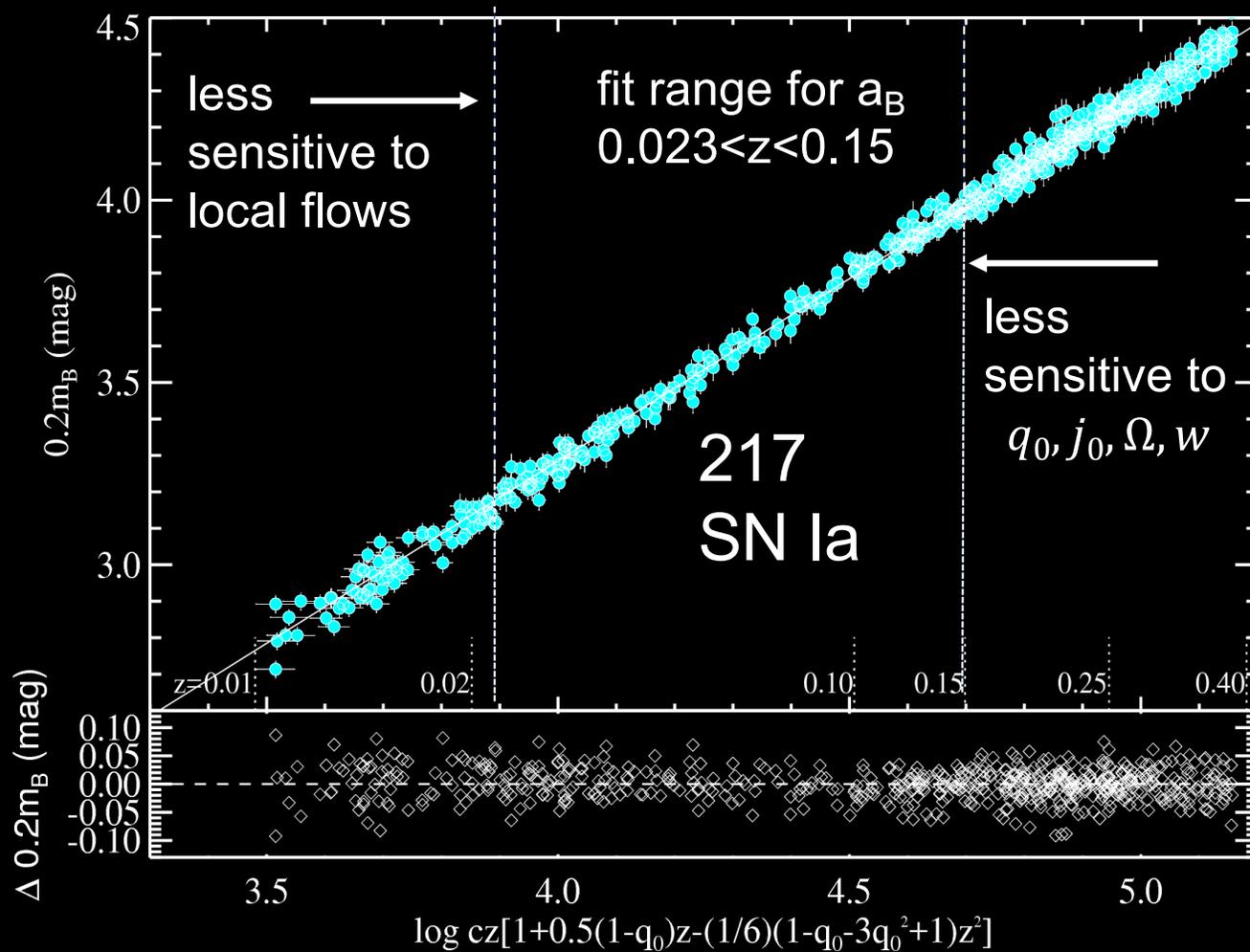
0.55 $\mu$ m



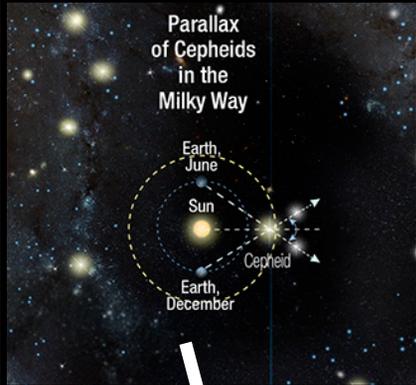
Leavitt

# Step 3: Intercept of SN Ia Hubble Diagram: Distance vs Redshift

$$a_B = \log cz \left\{ 1 + \frac{1}{2} [1 - q_0] z - \frac{1}{6} [1 - q_0 - 3q_0^2 + j_0] z^2 + O(z^3) \right\} - 0.2m_B^0 \leftarrow \text{Kinematic Intercept equation}$$

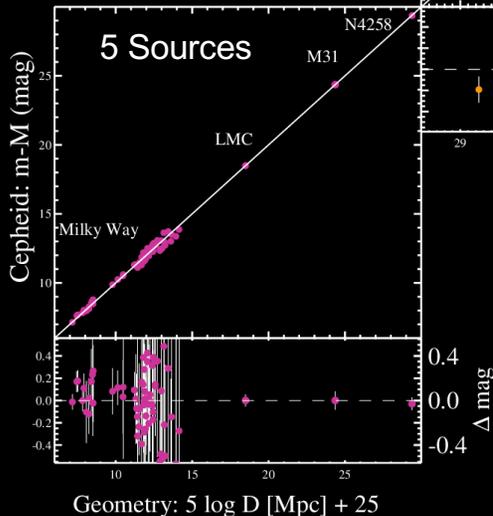


# The Hubble Constant in 3 Steps: Present Data

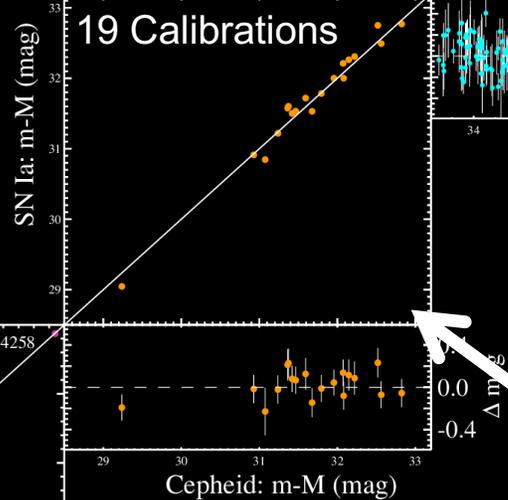


1

Geometry → Cepheids



Cepheids → Type Ia Supernovae

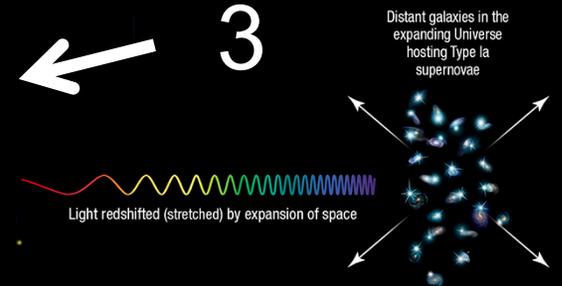
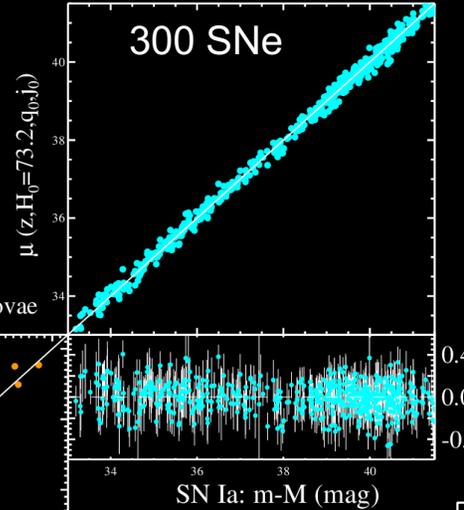


2

Galaxies hosting Cepheids and Type Ia supernovae



Type Ia Supernovae → redshift(z)



$H_0 = 73.5 \pm 1.4$ ,  
 $\text{Km s}^{-1} \text{Mpc}^{-1}$   
(Riess et al. 2019,  
Reid, Pesce, Riess 2019)

1.9% total uncertainty

4.2 $\sigma$  from CMB +  $\Lambda$ CDM !

\*Simultaneous Fit: Retain interdependence of data and parameters

# Robust? Seven Sources of Cepheid Geometric Calibration

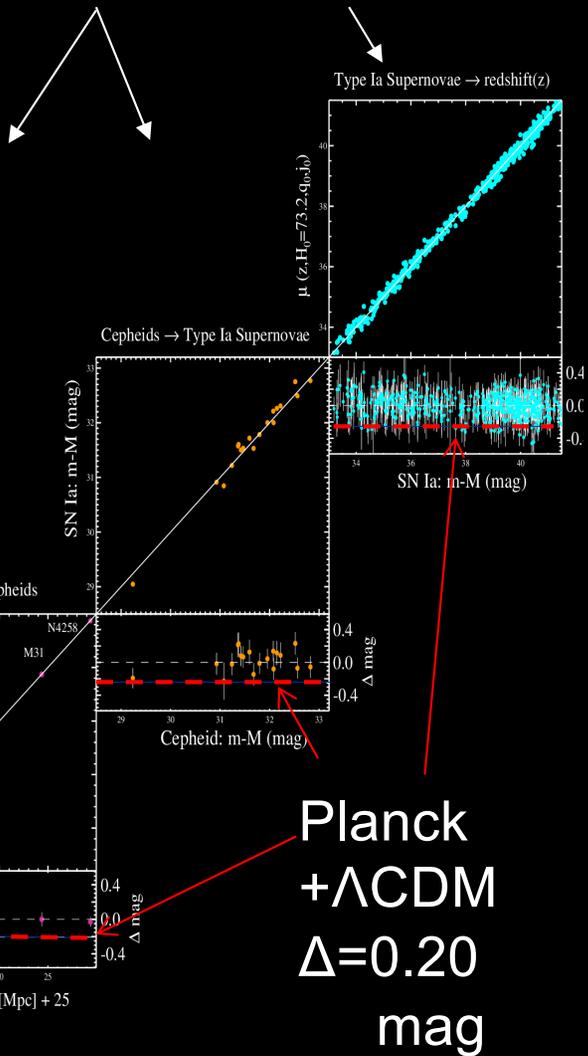
| Independent Geometric Source  | $\sigma_D$ | $H_0$ |
|---|------------|-------|
| NGC 4258 H <sub>2</sub> O Masers: Reid, Pesce, Riess 2019   | 1.5%       | 72.0  |
| LMC 20 Detached Eclipsing Binaries: Pietrzynski+ 2019 + 70 HST<br>LMC Cepheids: Riess+(2019) <b>AGREES WITH GAIA EDR3</b> | 1.3%       | 74.2  |
| Milky Way 10 HST FGS Short P Parallaxes: Benedict+2007 --also<br>Hipparcos (Van Leeuwen et al 2007)                       | 2.2%       | 76.2  |
| Milky Way 8 HST WFC3 SS Long P Parallaxes: Riess+ 2018  | 3.3%       | 75.7  |
| Milky Way 50 Gaia+HST, Long P Parallaxes: Riess+ 2018   | 3.3%       | 73.7  |
| Milky Way Short P Cepheid Binary Gaia Companion Parallax:<br>Breuval+20   | 3.8%       | 72.7  |
| Milky Way Short P Cepheid Cluster Gaia Parallax: Breuval+20   | 3.2%       | 73.6  |

Consistent Results ( $\leq 2\sigma$ ), *Independent Systematics*

# Systematics? 23 Analysis Variants—we propagate variation to error

Best Fit:

$$5 \log H_0 = M_B^0 + 5a_B + 25$$



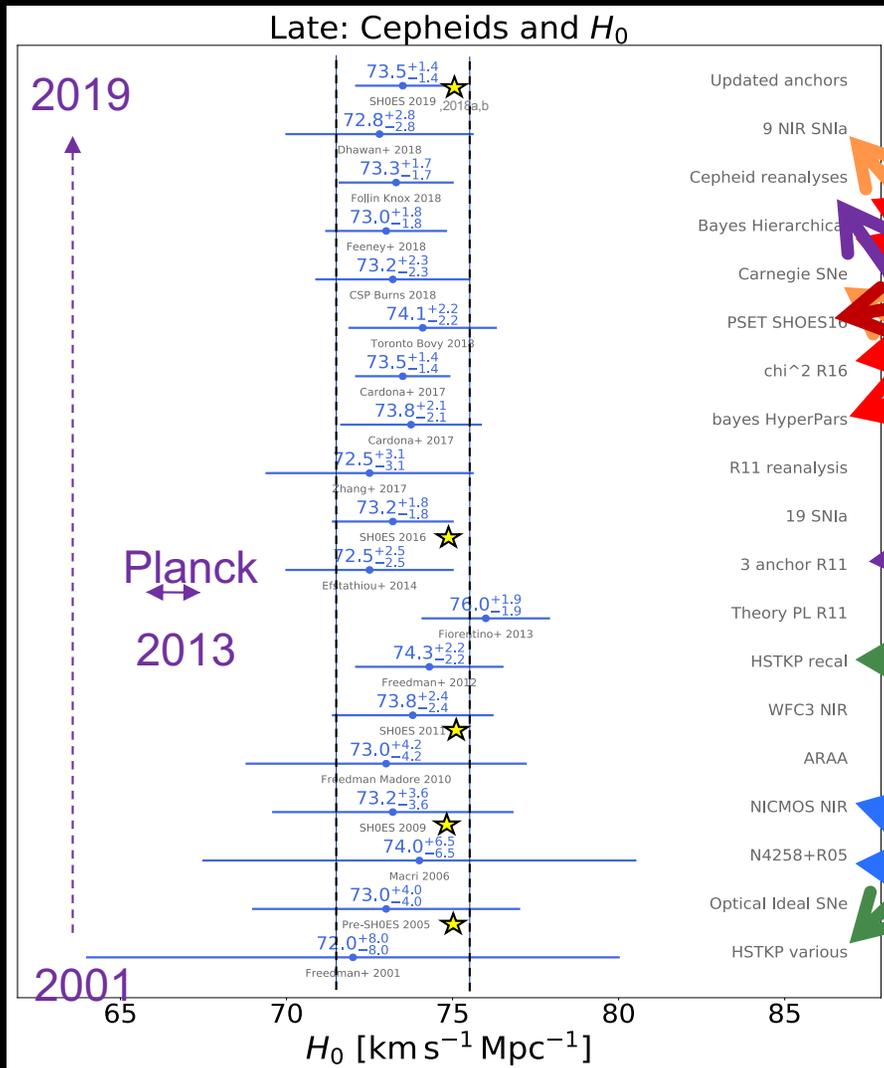
Planck  
+  $\Lambda$ CDM  
 $\Delta=0.20$   
mag

| Analysis Variants                              | $H_0$ |
|--|-------|
| Best Fit (2019)                                | 73.5  |
| Reddening Law: LMC-like ( $R_V=2.5$ , not 3.3) | 73.4  |
| Reddening Law: Bulge-like (N15)                | 73.9  |
| No Cepheid Outlier Rejection (normally 2%)     | 73.8  |
| No Correction for Cepheid Extinction           | 75.2  |
| No Truncation for Incomplete Period Range      | 74.6  |
| Metallicity Gradient: None (normally fit)      | 74.0  |
| Period-Luminosity: Single Slope                | 73.8  |
| Period-Luminosity: Restrict to $P > 10$ days   | 73.7  |
| Period-Luminosity: Restrict to $P < 60$ days   | 74.1  |
| Supernovae $z > 0.01$ (normally $z > 0.023$ )  | 73.7  |
| Supernova Fitter: MLCS (normally SALT)         | 75.4  |
| Supernova Hosts: Spiral (usually all types)    | 73.6  |
| Supernova Hosts: Locally Star Forming          | 73.8  |
| Optical Cepheid Data only (no NIR)             | 72.0  |

- Could we live in a giant void (9% in  $H_0$ )?  
No, LSS Theory and SN Ia mag-z limit  $\sigma \sim 0.6\%$  in  $H_0$   
[Odderskov et al. \(2016\)](#) , [Wu & Huterer \(2017\)](#), [Kenworthy, Scolnic, Riess 2019](#)
- Is HST WFC3-IR flux scale linear to 1%?  
Yes, calibrated to  $\sigma = 0.3\%$  in  $H_0$  across 15 mag  
[Riess, Narayan, Calamida 2019](#)
- Does Cepheid crowding compromise accuracy?  
No, amplitude data confirms locality of crowding  
[Riess, Yuan, Casertano, Macri, Scolnic 2020](#)
- Is there a difference in SN Ia at ends of distance ladder?  
No, correlations of Hubble residuals  $< \sigma = 0.3\%$  in  $H_0$   
[Jones et al 2018](#)

# Cepheids+SN Ia Ladder, Most Widely Replicated: 2001-2019

Why Cepheids? Advantages: 1) longest-range 2) most calibrations 3) consistent photometry along ladder 4) most tested...



SH<sub>0</sub>ES results (★) *cumulative* but compared to present... consistent

grad student problem set! (Toronto) Different analyses

Different SNe, wavelength

“Planck People”

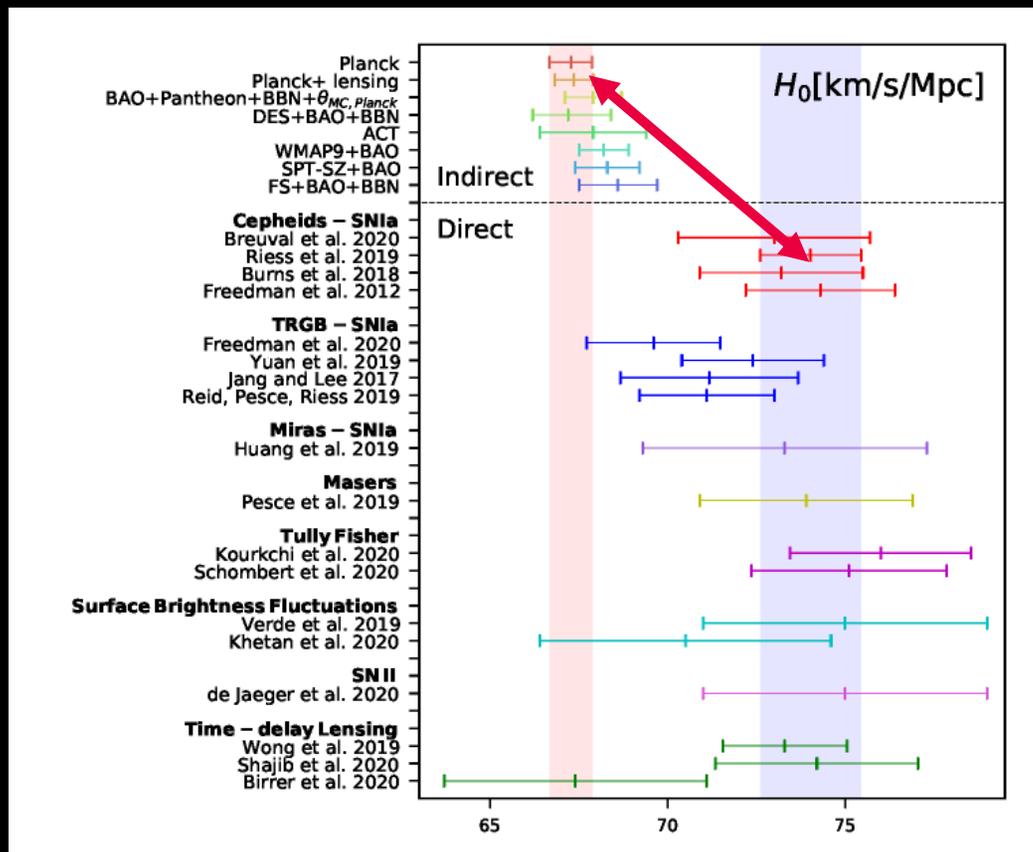
Different Team (KP), photometry, Cepheids, wavelengths

Different HST Instruments

# The Hubble Constant Tension, Discrepancy, Problem, Crisis

## Status late 2020

KITP 2019 (Verde, Treu, Riess 2019)



*“does not appear to depend on the use of any one method, team or source”*

No Cepheids:  $4.5-5.3\sigma$

No TRGB:  $5.7-6.3\sigma$

No lens:  $5.0\sigma$

No SN Ia:  $4.9\sigma$

No Cepheids or TRGB:  $5.3\sigma$

No Planck:  $4.4-4.9\sigma$

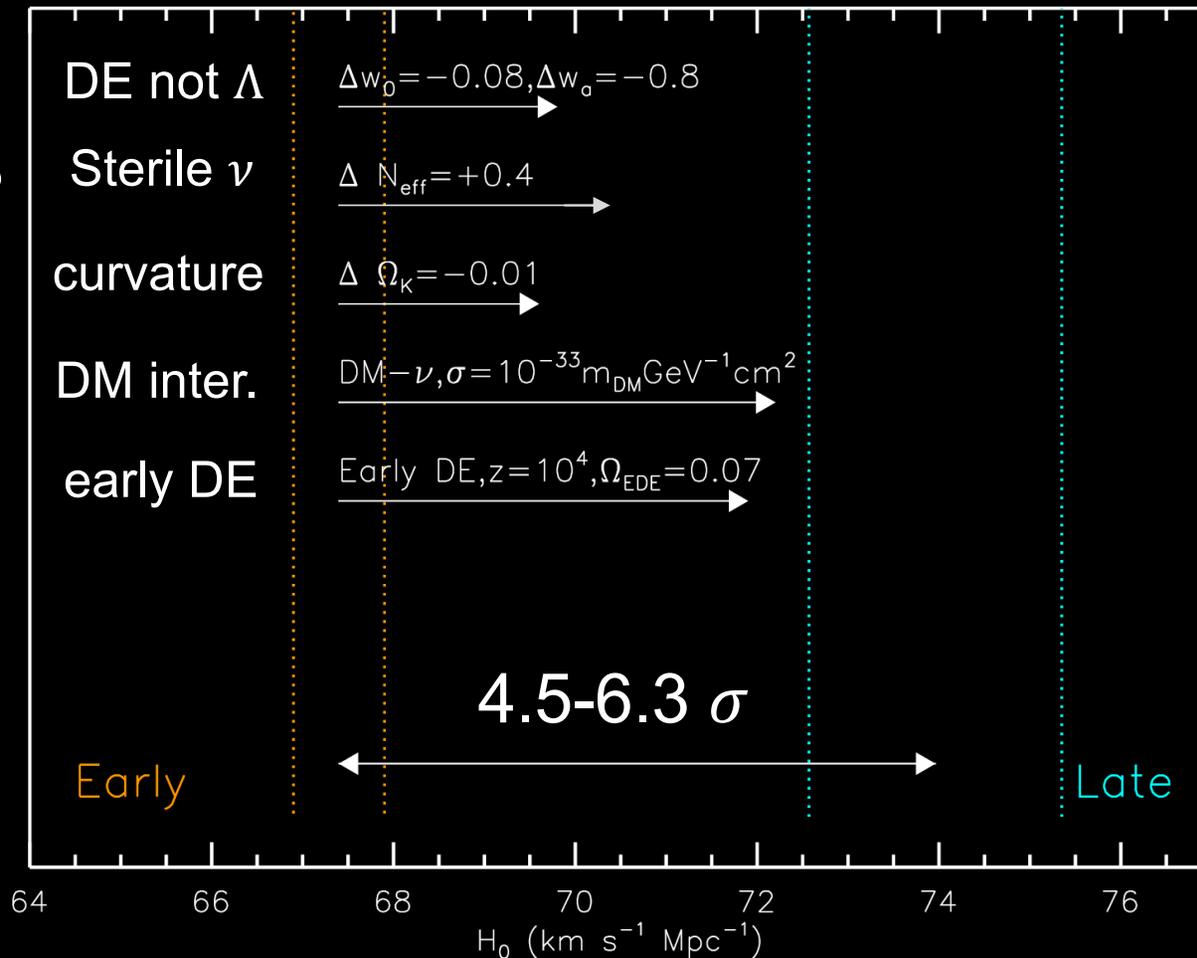
No CMB:  $4.0-4.5\sigma$

(Riess 2019, Nature Reviews)

Compilation from Di Valentino(2020)

# Cause Early vs Late Difference? Newton: “Feign No Hypothesis”

NEW  
PHYSICS  
?



“The Hubble Hunter’s Guide”, Knox and Millea, 2019: “Most Likely”: Increase Expansion Rate Pre-recombination  $\rightarrow$  reduce sound horizon by 5-8%

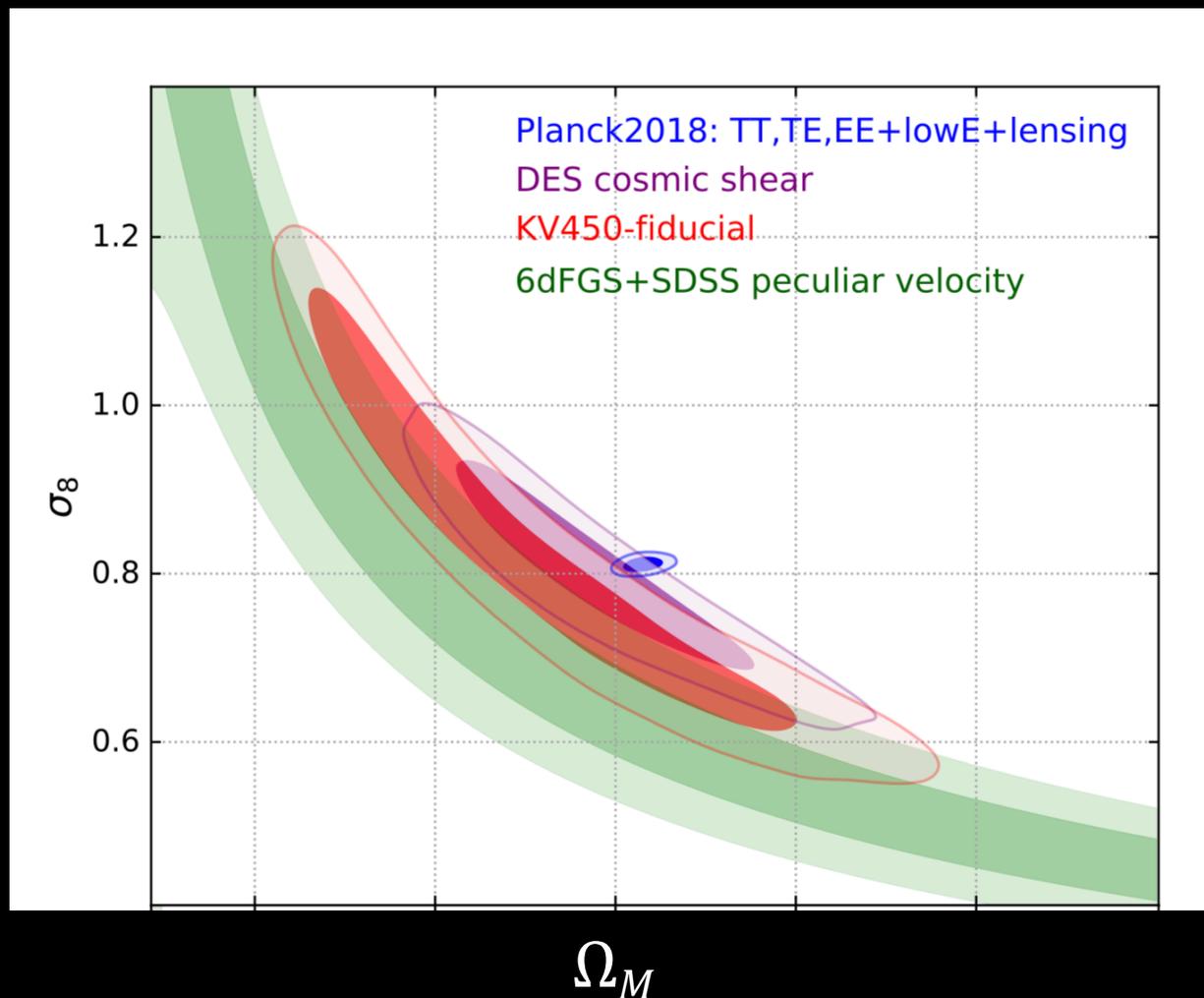
Mechanisms: Early DE or sterile (self-interacting) neutrinos

Claims: better fit to CMB, new CMB features, cosmic birefringence as evidence of CMB coupling to EDE/ALPs or pNG Boson (Capparelli+20, Fujita+20)?

# Another Early vs Late Tension? Matter clumpiness, $\sigma_8$

RMS matter fluctuation,  $\sigma_8$ , ( $r=8 h^{-1}$  Mpc), 0.8 Early vs late divide

$\sim 3\sigma$  from lensing and peculiar velocities, independently



6dFGS+SDSS

Said, K et al 2020,  
MNRAS, 497, 1275

“...deviates by more than  $3\sigma$  from the latest Planck CMB measurement. Our results favour ... a Hubble constant  $H_0 > 70 \text{ km s}^{-1} \text{ Mpc}^{-1}$  or a fluctuation amplitude  $\sigma_8 < 0.8$  or some combination of these.”

# Can We Believe Measurements without Explanation?

Don't sweep "problems" under the rug



"Problems" are often clues!

~~Precession of Mercury~~

Solved!

~~Solar Neutrino Problem~~

Solved!

~~Missing Baryon Problem~~

Solved!

Lithium Problem

CMB Cold Spot

Flat rotation curves/  
what/where is dark matter?

Accelerating Universe/  
why  $\Lambda$  so small?

# Can We Believe Explanation without hypothesis (*how*)?

Present data provides formidable challenge!

“Its New Physics”—constrained precise  $H(z)$  data, CMB high- $l$

“Its Systematics”—many measures, many independent rungs, duplicate measurements, Copernican principle

I don't think so.

Reasons for optimism:

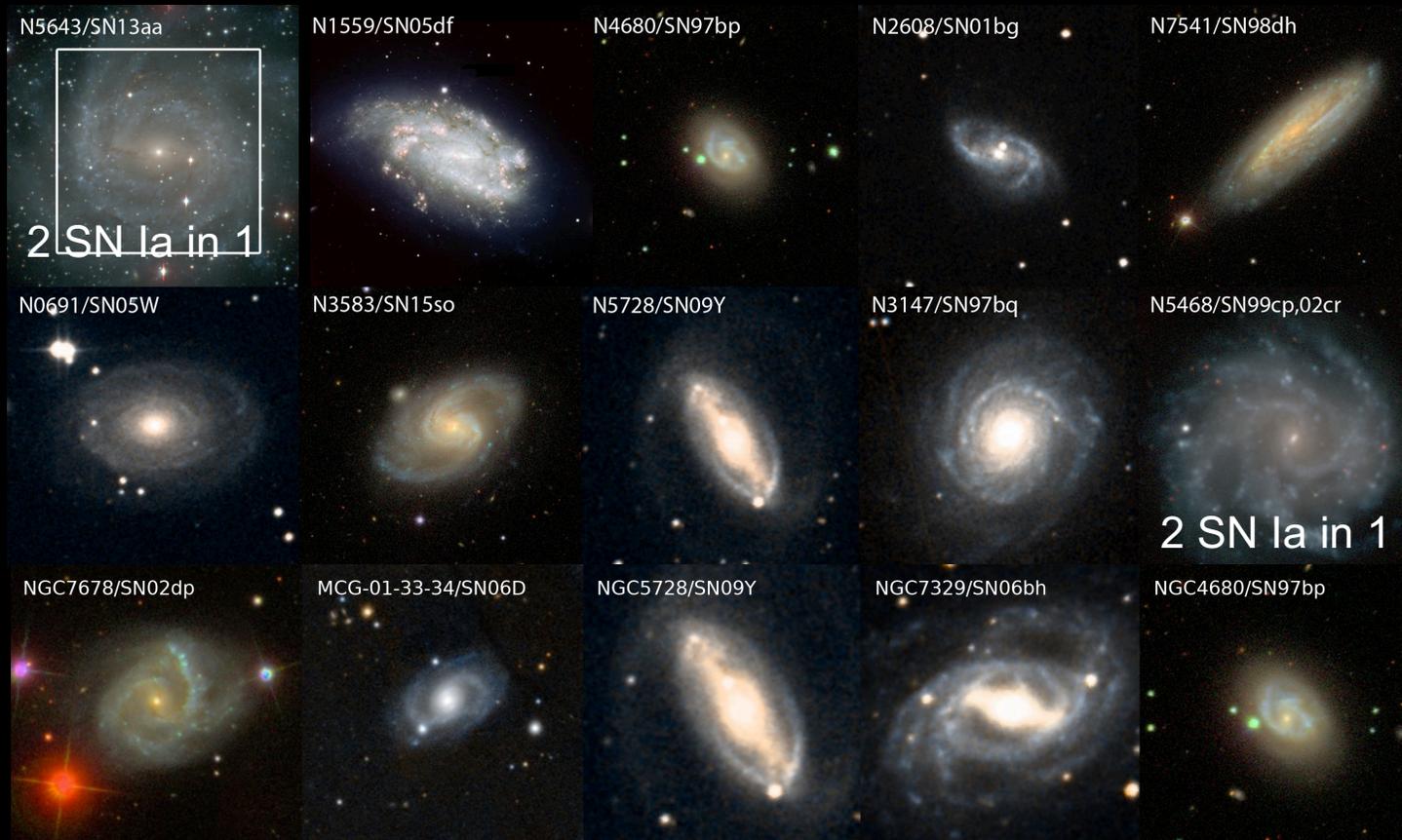
New data: LIGO, DESI, Roman, Rubin, Euclid, JWST, Simons, S4

New clues: Early vs late  $\sigma_8$ , Cosmic Birefringence?, BBN ?

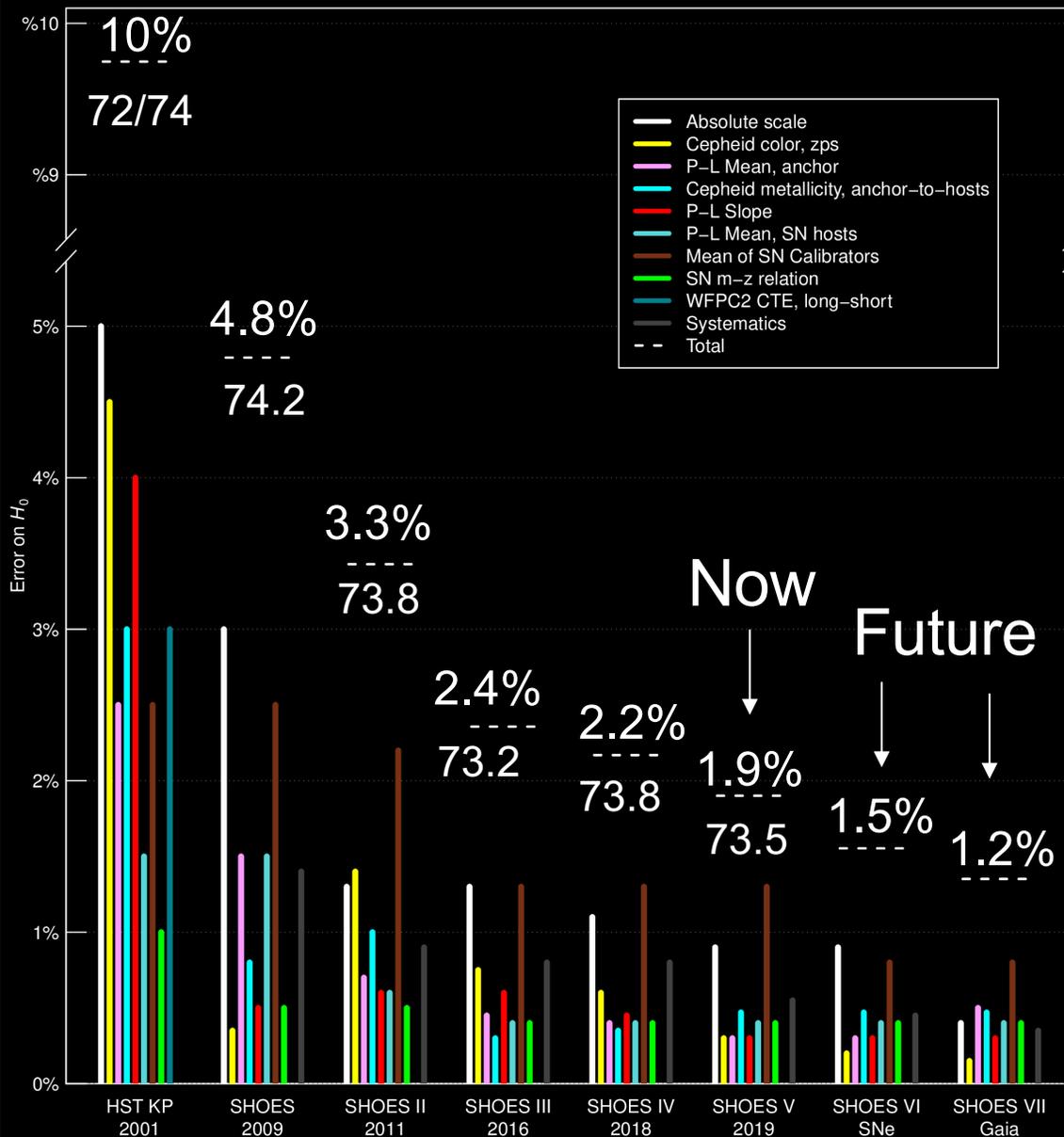
Big Playground: Lambda CDM is 95% dark, quantum gravity

# Next Steps: Increasing Number of SN-Cepheid Calibrations

**\*NEW\*** SHOES Large HST Programs, Cycles 25,26,28  
24 more Cepheid-SN Ia Calibrators underway,  
to reach total=43, + Cepheids to Coma!



# Future Prospects...



- **New low-z SN samples**
- **Doubling SN Calibrator sample, 19→40**
- **Gaia EDR3 !!!**
- **LIGO  $H_0$  (Late Universe)**
- **DESI, LSST, WFIRST, Euclid → better  $w(z)$**
- **Next generation CMB: signatures (e.g., EDE)**
- **Stay tuned...**

# Final Thoughts

- Discrepancy is  $\sim 5\sigma$  (4-6)  $\sigma$  (depending on combination)  
No Late Universe measurements lower than any Early
- Appears robust, requires multiple catastrophic failures to avoid
- Very interesting! (unless your Bayesian prior on  $\Lambda$ CDM  $> 5 \sigma$ )
- Feign No Hypothesis, let's follow evidence, find the *how*
- Universe may be more clever than we are *now*