The "Bosma Effect" Revisited

Correlations between the ISM and DM in Galaxies

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A Brief History of DM in Galaxies

- "Dark matter" needed to explain local stellar kinematics (Kapteyn 1922, Oort 1932)
- Flat rotation curve of M31 (Babcock 1939, Mayall 1951)
- MW globular cluster kinematics (Kurth 1950)
- Local Group kinematics (Kahn & Woltjer 1959)
- Problem with stability of massive discs (Toomre 1964)
- CDM halo provides stability (Ostriker & Peebles 1973)
- Ubiquity of "flat" rotation curves

(Rogstad & Shostak 1972, Bosma 1978, Rubin, Ford & Thonnard 1980)

- Stellar disc-halo "conspiracy" (van Albada & Sancisi 1986)
- "Galaxies are irrelevant": CDM needed for LSS, ΛCDM cosmology (1990's-present)
- Bullet cluster: DM not in baryonic intracluster medium stripped from galaxies
- CDM halos can't be cuspy, so add toy gastrophysics until it fits



Classical Lines of Evidence for **Cold** DM

WDM

- Galaxy Dynamics
 - Stellar dynamics in the solar neighborhood
 - Spiral galaxy rotation curves
 - Stability of galaxy disks, spiral density waves
 - Projected kinematics of elliptical galaxies
 - Local Group kinematics
 - X-ray gas in elliptical galaxies
 - Strong gravitational lensing
- Galaxy Clusters
 - X-rays
 - Strong gravitational lensing
 - Weak gravitational lensing (e.g. "bullet cluster")
 - Sunyaev-Zeldovich Effect
- Cosmic Background Radiation (e.g. WMAP)
- Large-scale structure formation
 - Baryonic acoustic oscillations
 - Galaxy correlation functions
 - Number and distribution of galaxy masses

Baryons matter (not just the stars)!

- Tully-Fisher relation (Tully & Fisher 1977)
- V_{DM}² ∝ V_{gas}² (Bosma 1978, 1981)
- Stellar disc halo conspiracy (URC) (Bahcall & Casertano 1985; van Albada & Sancisi 1986)
- Maximum discs
 - (van Albada & Sancisi 1986)
- MOdified Newtonian Dynamics (Milgrom 1983)
- Baryonic Tully-Fisher relation (McGaugh et al. 2000, Pfenniger & Revaz 2005)
- Mass discrepancy acceleration relation (McGaugh 2004)
- Galaxies are a 1-parameter family (Disney et al. 2008)
- Constant mean DM & baryonic mean surface densities (Donato et al. 2009; Gentile, Famaey & Zhao 2009)

MOdifed Newtonian Dynamics (MOND)

- Milgrom (1983)
- Fundamental universal acceleration scale a₀ due to
 - modified inertia $F = [m^*\mu(a/a_0)]^* a$
 - modified gravity $a = a_{Newton} / \mu(a/a_0)$ where $\mu(x>>1)=1$, $\mu(x<<1)=x$
- For modified gravity, Poisson equation is

$$\nabla \cdot \left[\mu(|\nabla \phi|/a_0) \nabla \phi \right] = 4\pi G \rho$$

- Fits to rotation curves yield $a_0 = 1.2 \ge 10^{-10} \ge h_{75}^2 \sim 0.1 \ge 0.1$
- Many successful predictions for properties of galaxies
- Theoretical basis could be TensorVectorScalar gravity
- Can also explain "bullet cluster", WMAP angular power, gravitational lensing, ...
- Functions so well, that if not an alternative to Einstein gravity then MOND says DM physics produces really bizzare correlations with baryons.

MOND Successes



The Problem with MOND



- Basically pure (though spectacularly successful) phenomenology
 - MOND-ish theories (TeVeS, conformal gravity) are inelegant
- No laboratory / Solar system tests



MOND is telling us that baryons are more important than we thought.

Baryonic Tully-Fisher Relation



McGaugh et al. (2000)

Mass Discrepancy - Acceleration Relation



McGaugh (2004)

Galaxies are a 1-Parameter Family



Disney et al. 2008

The "Bosma Effect"





"... the ratio [of dynamic to gas surface densities] ... is more or less constant beyond about onethird of the optical radius, with HI being the dominant contributor ... in the outer parts"

Dependence on Galaxy Parameters



Testing the Bosma Effect



Conclusions of HvA&S

- "The model curve [of the poorer fits] does not agree with the observed rotation curve in the inner region."
- There are "... large wiggles that are not present in the observed rotation curve."
- "The model rotation curve drops below the observed rotation curve at large radii."
- "... scaling of HI to represent the dark component only works in combination with maximal discs."
- "... our sample is biased against galaxies with R_{out}/h_{HI} substantially larger than 3."
- "... for about two-thirds of the galaxies we obtain good fits to the data."
- "... the good fits are somewhat coincidental."

Simply the Effects of CDM?

1978 : DM can be in the disc 2001 : CDM <u>must</u> be in the halo



The Bosma Effect in Nearby Galaxies



Spitzer Infrared Nearby Galaxy Survey



Rotation Curve Models

Normally :

$$V_{\text{tot}}^{2} = \Upsilon_{\text{disk}} V_{\text{disk}}^{2} + \Upsilon_{\text{bulge}} V_{\text{bulge}}^{2} + V_{\text{HI+He}}^{2} + V_{\text{DM}}^{2} + V_{\text{mol-H+He}}^{2}$$
$$V_{\text{DM}}^{2} = f(\rho_{0}, r_{c}), f(V_{200}, c), f(V_{200}, c(V_{200})), \dots$$

"Simple" Bosma effect = "HI-scaling" :

$$V_{\text{tot}}^2 = \Upsilon_{\text{d,IR}} V_{\text{disk}}^2 + \Upsilon_{\text{b,IR}} V_{\text{bulge}}^2 + (1 + f_{\text{HI}}) V_{\text{HI+He}}^2$$

"Classic" Bosma effect :

 $V_{tot}^{2} = (1 + f_{disc}) \Upsilon_{d,IR} V_{disk}^{2} + \Upsilon_{b,IR} V_{bulge}^{2} + (1 + f_{HI}) V_{HI+He}^{2}$

The "simple" Bosma Effect: Pure HI-scaling





HI Distributions of Galaxies



Rhee & van Albada (1996)

The "classic" Bosma Effect



Bosma effect vs. CDM











Results



- Self-consistent NFW model ruled out
- "Simple" Bosma effect = "HI scaling" only works outside of the stellar disk
- "Classical" Bosma effect with stellar proxy nearly as good as URC/Burkert

Implied Surface Densities



Paper II: Including More of the ISM



Paper III: What does the Bosma effect mean?

• CDM? : disk potential fundamentally different from that of a spherical distribution





CCM has no means of teaching small amounts of gas in a disk to behave as if it was distributed exactly as a spherical CDM halo - is galactic DM then baryonic??

What does the literal Bosma effect mean?

- Only about 10-70% of the baryons are visible
- The Utility of "maximal disks" is explained
- The relative mean surface density constancy is explained

$$<\Sigma>_{\rm DM}/<\Sigma>_{\rm baryons} \approx _{\rm DM}/_{\rm baryons} \approx 5$$



• The extended baryonic Tully-Fisher relation (Pfenniger & Revaz 2005) $log (M_*+c M_{HI+He}) = a+b*V_{rot} - c \sim 3$

Discs are More Efficient Sources of V²

DM sphere with flat rotation curve: $\rho(r) = (M_{vir}/4\pi r_{vir})(r_{vir}/r)^2$ $V(r)^2 = G M_{vir}/r_{vir} = const$

Mestel disc: $\Sigma(R) = (M_{disc} / (2\pi R_{disc}^2)) (R_{disc} / R) a\cos(R_{disc} / R)$ $V(r)^2 = \pi G M_{vir} / 2r_{vir} = const$

 $M_{disc}/M_{vir} = (2R_{disc}/\pi r_{vir}) \approx 10 \text{ kpc} / 100 \text{ kpc} = 0.1$

Are there other Signs of a Hidden ISM?

- Cold H₂ "clumpescules" (Pfenniger & Combes 1994)
- "Extreme Scattering Events", 1~AU (Walker & Wardle 1998)
- MSX, PLANCK "cold cores", l~pc (Egen et al. 1998, Ade et al. 2011a)
- EGRET "dark gas"

(Grenier et al. 2005)

- Dwarf galaxies from collisional debris (Bournaud et al. 2007)
- PLANCK "dark gas" phase (Ade et al. 2011)
- HERSCHEL dwarf galaxy survey (Madden et al. 2011)

Egen et al. 1998



Local Stellar Dynamics Revisited

Milky Way (NASA)





The Local Mass-Density Revisited



The Bosma Effect & MOND

Define : $g_{tot} = g_* + g_{gas} + g_{dDM}$ $= g_{vis} + g_{dDM} = g_* + g_{ISM} = g_* + (1 + f_B) g_{gas}$ Thus (Dunkel 2004) : $\varepsilon = g_{tot}/g_{dDM} - 1 = g_{vis}/g_{dDM}$ $\epsilon/(\epsilon+1) - g_{vis}/g_{tot} = \mu(\epsilon)$ $g_{vis} = \mu(\varepsilon) g_{tot}$ MOND : $g_{vis} = \mu(x) g_{tot}$ $x/(x+1) = \mu(x)$ $x = g_{tot} / a_0,$ If $x = \varepsilon$: g_{tot}/g_{dDM} - 1 = g_{tot}/a_0 $1/a_0 = 1/g_{dDM} - 1/g_{tot}$ $g_{tot} / a_0 = g_{vis} / g_{dDM} = g_* / g_{dDM} + 1 / f_B$ $\approx g_* / g_{dDM}$



The local gravitational field (and approximately the total local surface density) determines how much mass is in stars vs. in dDM

The Mass Discrepancy-Acceleration Relation



Halo Mass of the Milky Way Revisited

- Estimates for total baryonic mass $M_b+M_d+M_g+M_{dDM} \approx (0.1+0.7+0.1+0.3) \ 10^{11} M_{Sun} \approx 1.210^{11} M_{sun}$
- Concordance assumptions & result (Watkins, Evans & An 2010)
 - Most of mass in NFW halo with scales >> $\rm r_{vis},$ all satellite galaxies observed are bound
 - Result: $M_{halo}(r < 300 \text{ kpc}) \sim 13 \ 10^{11} M_{Sun} \sim 15 \text{x}$ visible disc
- Non-standard assumptions & result
 - Kinematics of satellite galaxies with r > 40 kpc
 - Isotropy parameter $\beta \sim 0$
 - Leo I & Hercules are not bound (2 most extreme outliers from 28)
 - Result: $M_{halo}(r < 300 \text{ kpc}) \sim 4 \ 10^{11} M_{Sun} \sim 3x$ total disc



The results of satellite kinematics depends upon poor statistics & what one assumes, but one needs a modest (W?)DM halo at scales of the Local Group.

The Bosma Effect & Warm DM



- Disk DM cannot explain kinematics at large distances (e.g. Milky Way & M31 satellites, massive ellipticals)
- Disk DM cannot explain galaxy clusters
- WDM naturally fills in the gap at large radii.

Disk Stability?

- Increase $\Sigma(R)$ by a factor of ~3, Q = $\sigma \kappa / \pi G \Sigma < 1$
- Real discs are not uniform, axisymmetric, thin
- Real ISM chemistry complicated
- Read ISM is fractal
- Stability of turbulent media complicated (Romero, Burkert, Agertz 2010)
- Spiral structure is non-stationary (Sellwood 2010)
- m=1 structure seen in 56% of non-interacting galaxies (Van Eymeren et al. 2011)
- m=1 structure seen in inner galaxies (Rix & Zaritsky 1995)
- Dark component in discs are stabler than one expects (Revaz, Pfenniger, Combes & Bournaud 2009)



Q > 1 keeps galaxies from looking like galaxies

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 - Sunyaev-Ze
- Cosmic Backg
- Primordial nu
 - Baryonic a fig 400 Galaxy cor: @ 200
 - Number an





WDM

Summary

- The "Bosma effect" the correlation between the centripetal contribution of the dynamically unimportant visible gas and DM is clearly seen in the THINGS+SINGS data.
- It is physically implausible for DM in a spherical halo to force the ISM in a disk to show exactly the same centripetal signature, despite different geometries.
- The Bosma effect appears to be telling us that there is more baryonic matter in the discs of spiral galaxies and no need for a halo of COLD DM.
- The Bosma effect explains lots (but not all) of the baryon-DM correlations
- The theory and implications of disc DM need to be reconsidered.
- A non-cold DM component is still needed for the Local Group, massive galaxies, clusters, and LSS
- W/HDM and baryonic disc DM seem to be a perfect match.