

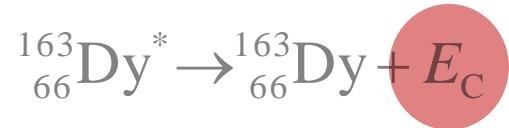
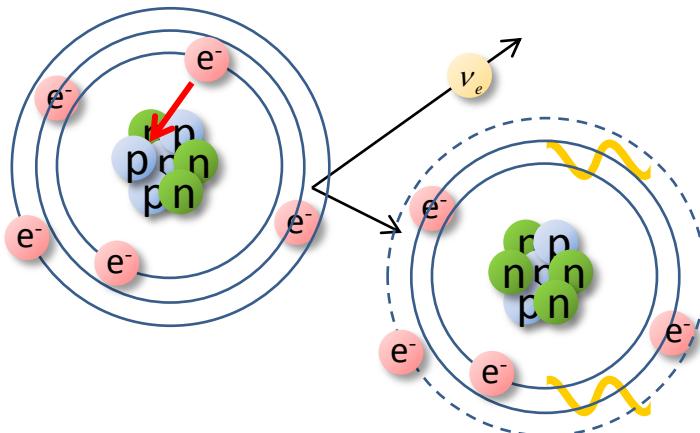


The Status of the ECHo experiment for investigation of keV sterile neutrinos

CLEMENS HASSEL
for the ECHo-Collaboration



^{163}Ho and neutrino mass



- $\tau_{1/2} \cong 4570$ years (2* 10^{11} atoms for 1 Bq)
- $Q_{EC} = (2.555 \pm 0.016)$ keV *

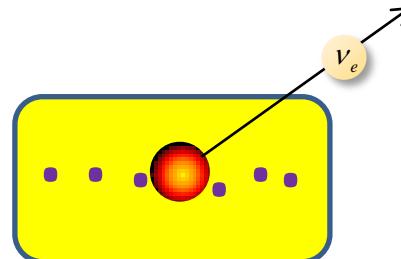
*M. Wang, G. Audi et al., *Chinese Phys. C* **36**, 1603, (2012)

A non- zero neutrino mass affects the **de-excitation energy spectrum**

Atomic de-excitation:

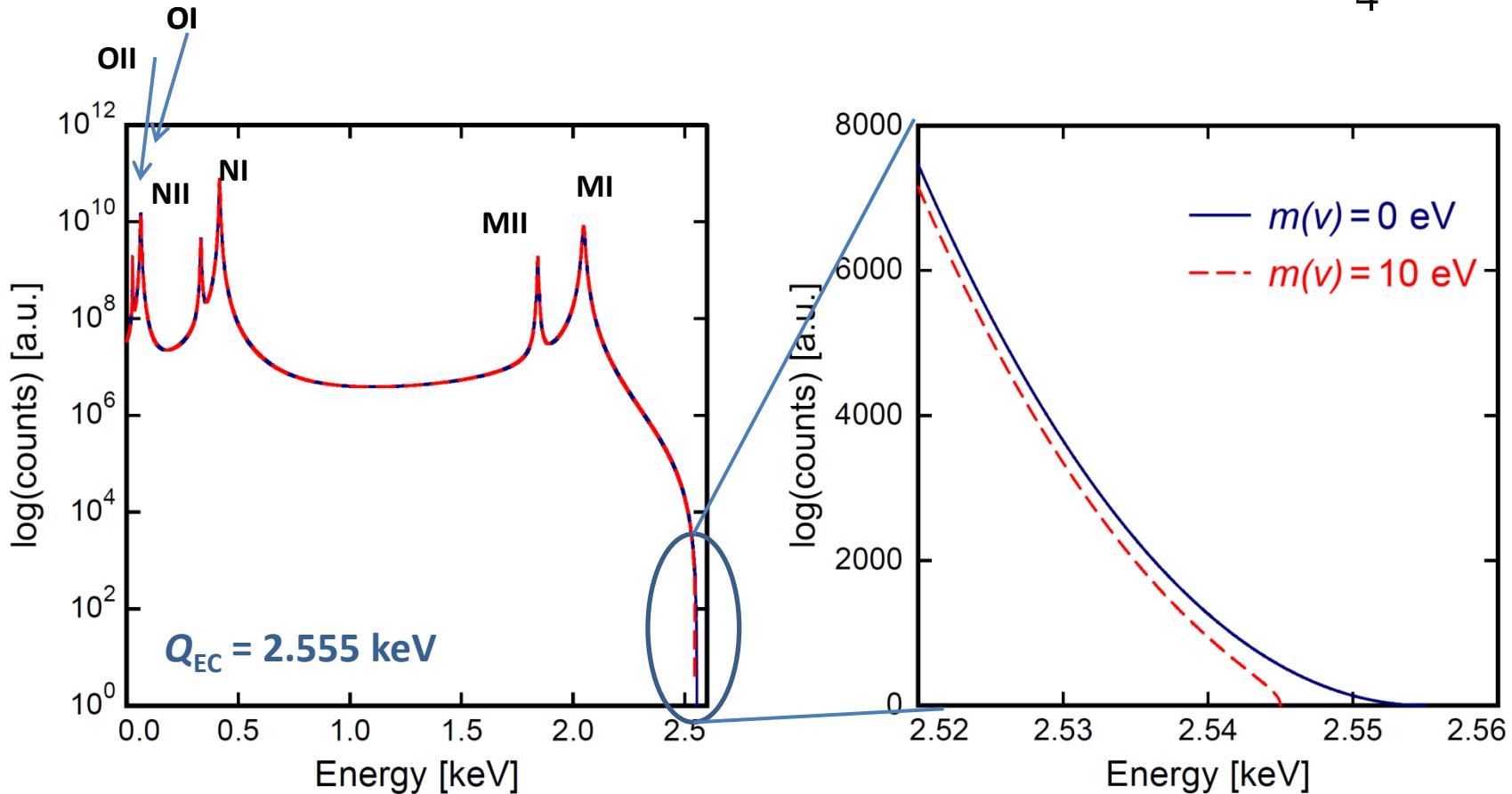
- X-ray emission
- Auger electrons
- Coster-Kronig transitions

} Calorimetric measurement



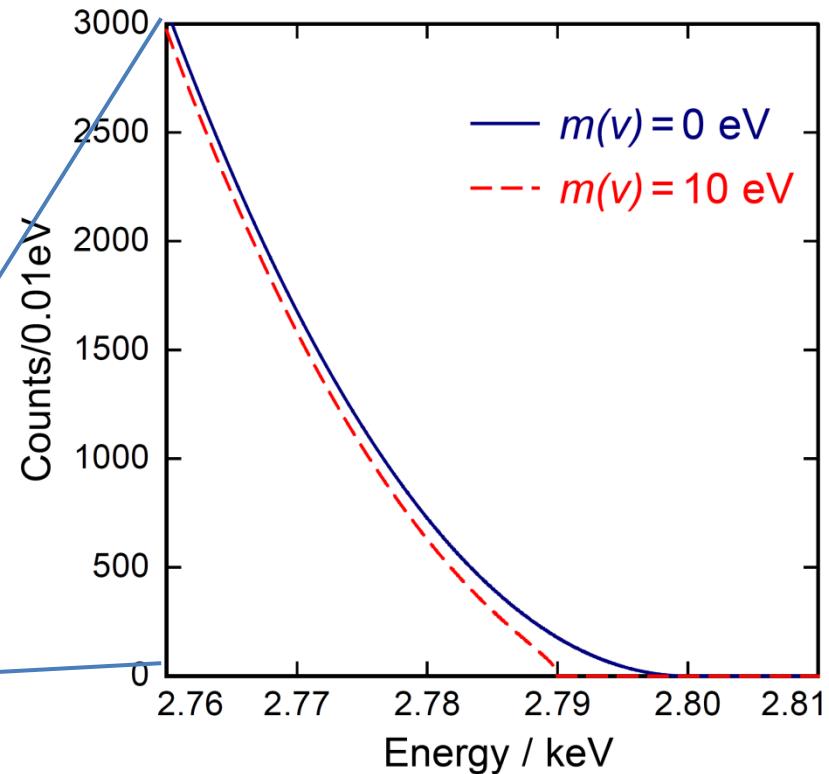
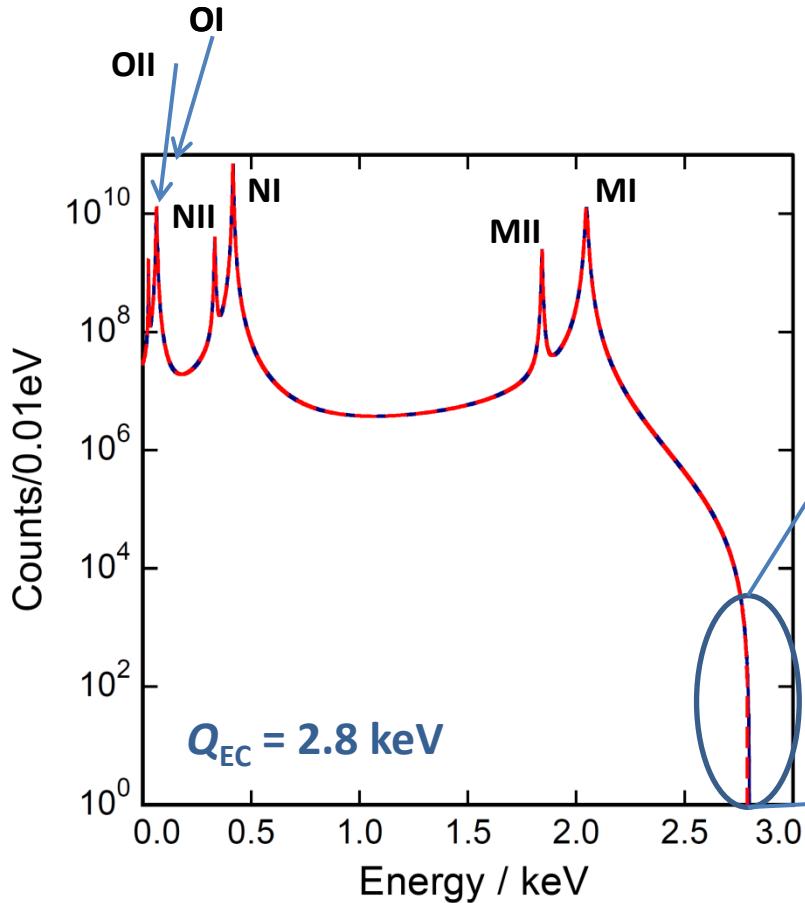
^{163}Ho and neutrino mass

$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



^{163}Ho and neutrino mass

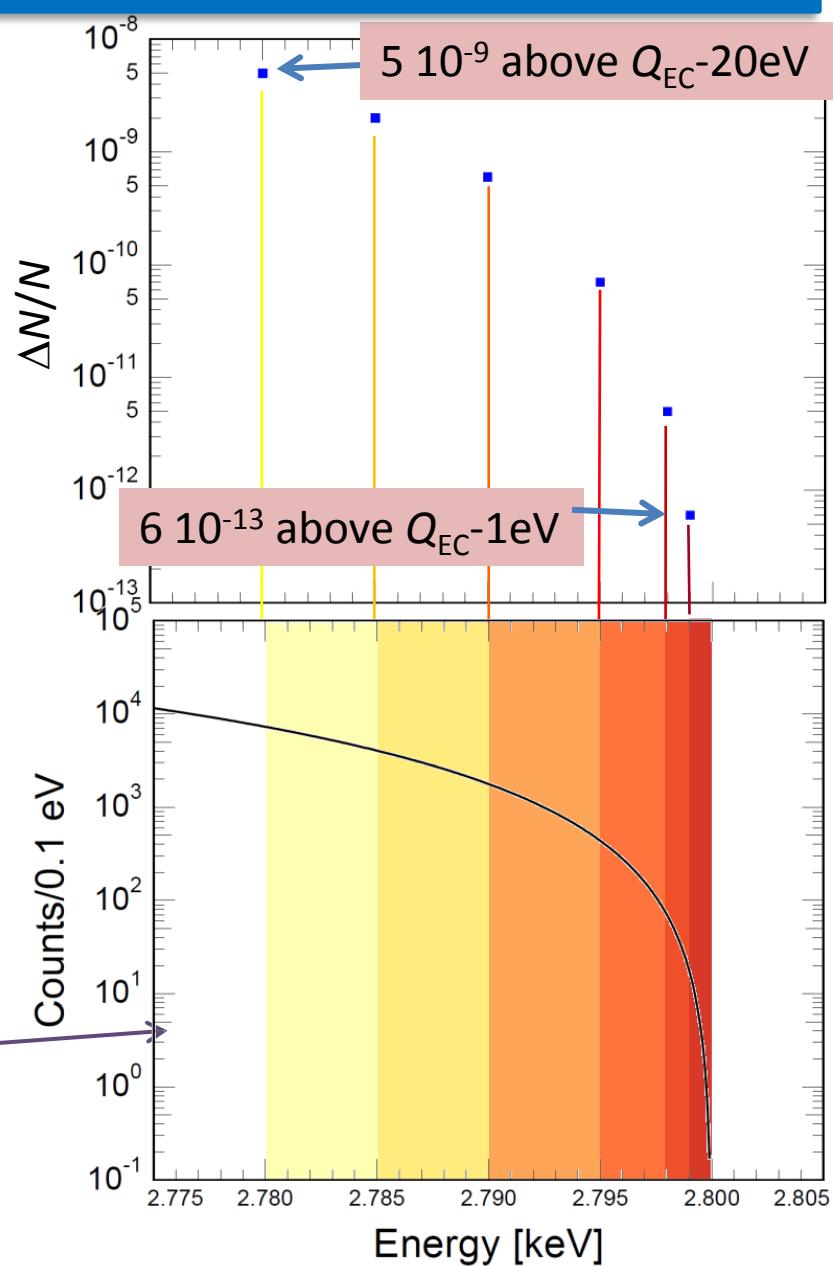
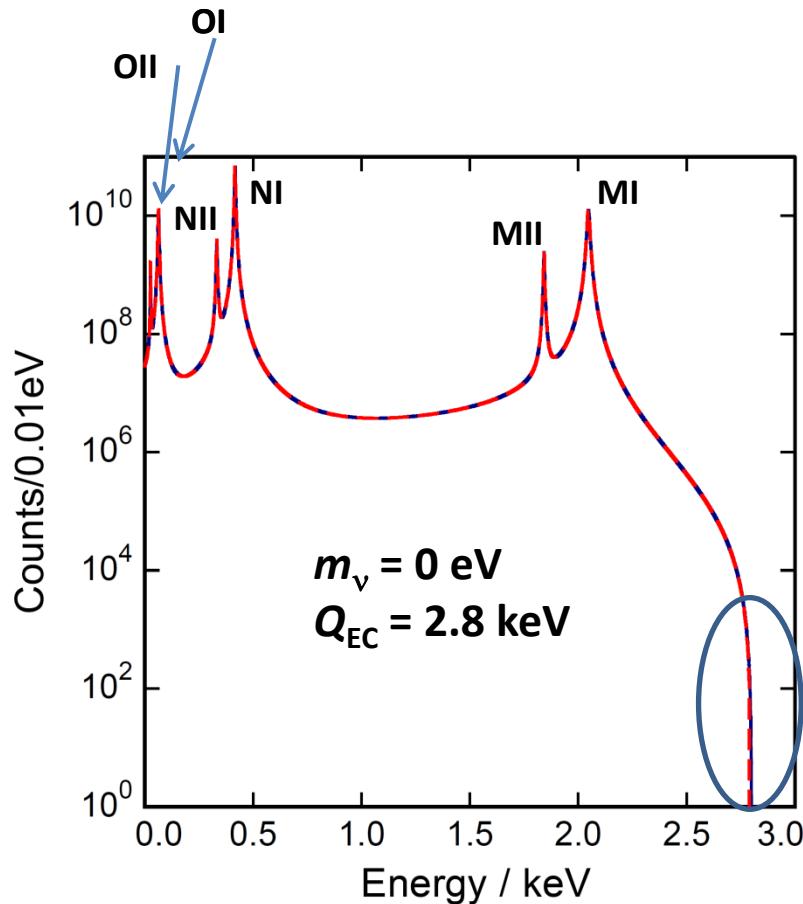
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^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$



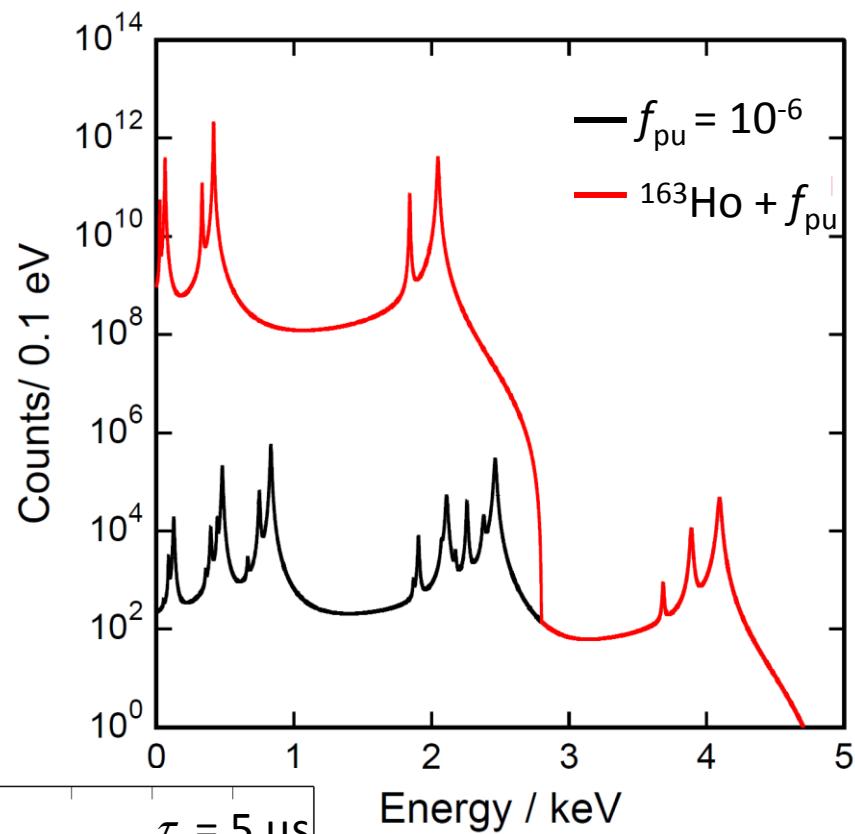
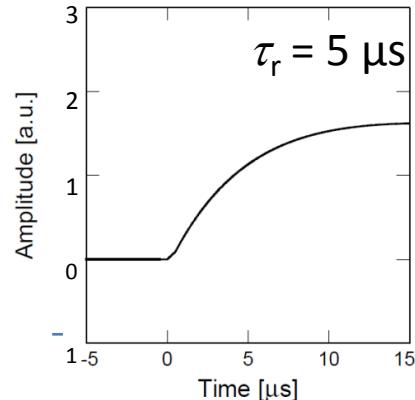
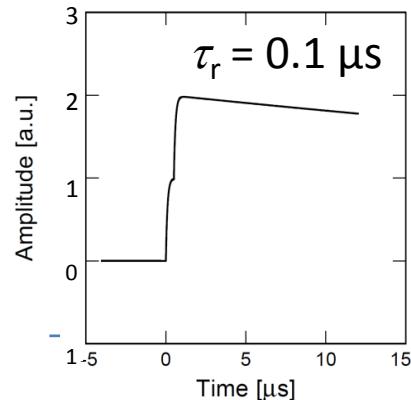
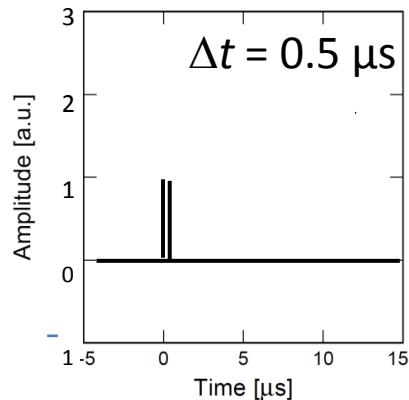
^{163}Ho and neutrino mass: sub-eV sensitivity

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Unresolved pile-up ($f_{\text{pu}} \approx a \cdot \tau_r$)

- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a \approx 10 \text{ Bq}$



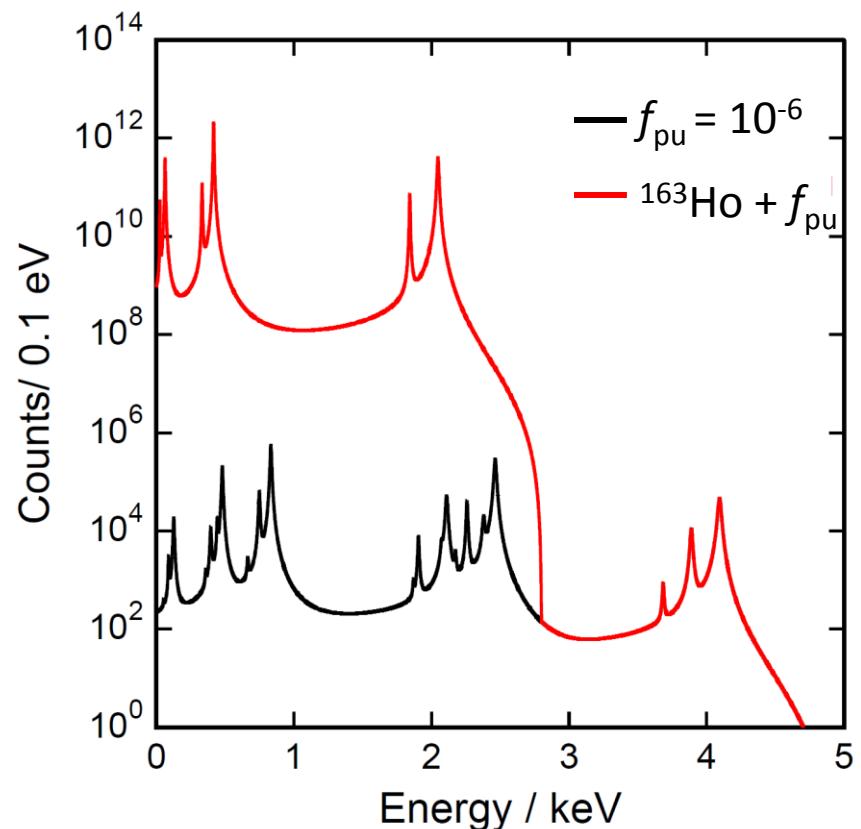
^{163}Ho and neutrino mass: sub-eV sensitivity

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- 10^5 pixels



^{163}Ho and neutrino mass: sub-eV sensitivity

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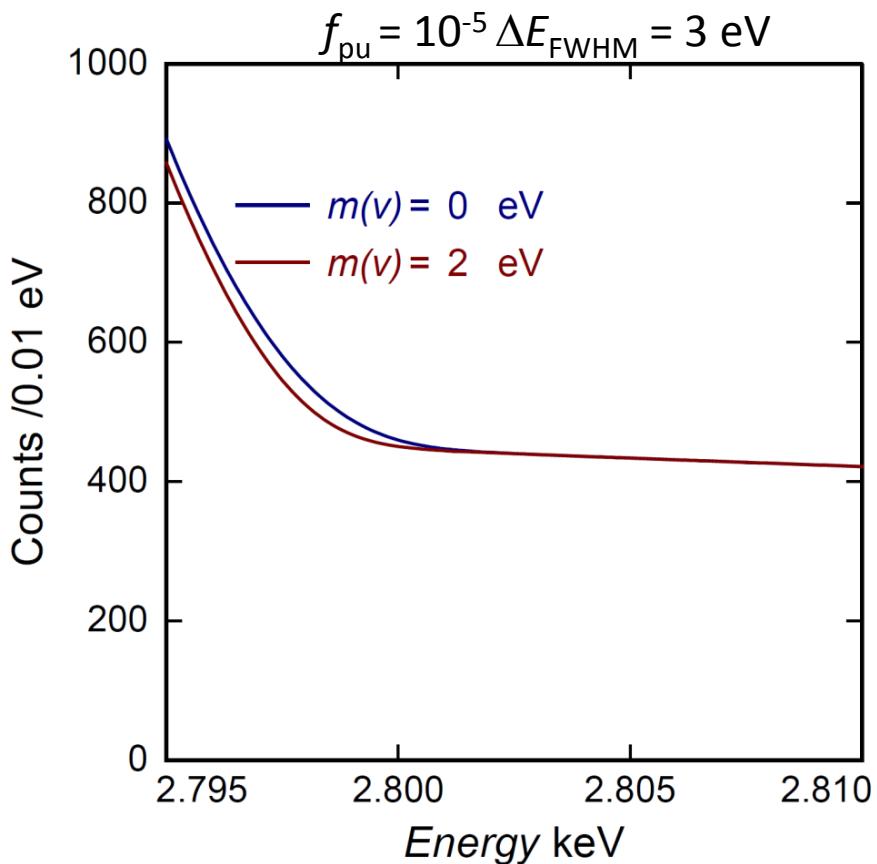
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- $\tau_r < 1 \mu\text{s} \rightarrow a \approx 10 \text{ Bq}$
- **10^5 pixels**

Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 2 \text{ eV}$



^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

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- $\tau_r < 1 \mu\text{s} \rightarrow a \approx 10 \text{ Bq}$
- 10^5 pixels

Metallic Magnetic Calorimeters
and
Microwave multiplexing
fullfil all the requirements

Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 2 \text{ eV}$

^{163}Ho and neutrino mass: sub-eV sensitivity

Statistics in the end point region

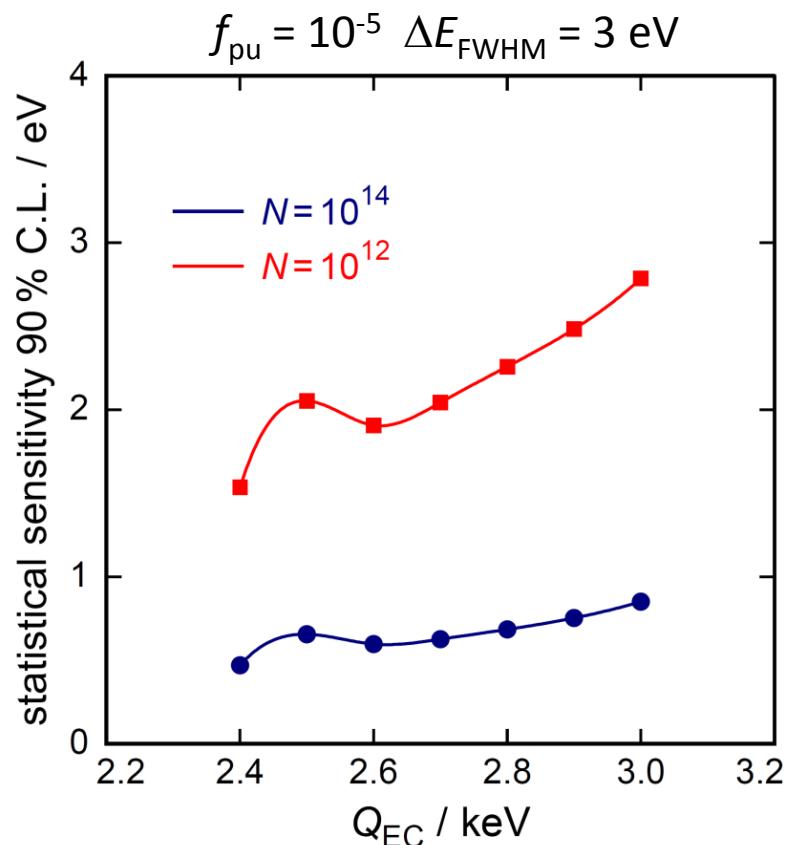
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- 10^5 pixels

Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 2 \text{ eV}$



Q_{EC} determination of ^{163}Ho

Penning Trap mass spectrometry

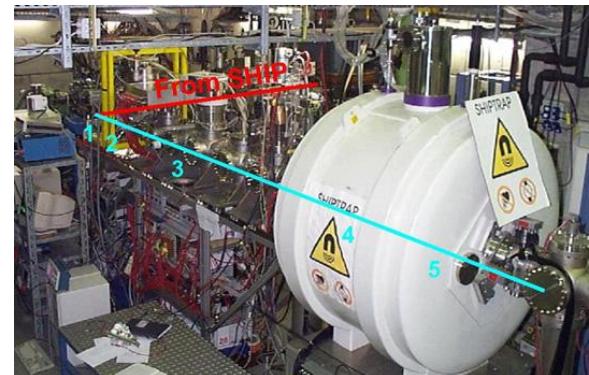
- First experiments at [TRIGA-TRAP \(Uni-Mainz\)](#) in 2014 *
 - Development of efficient Ho ion source using laser ablation
 - Uncertainties on ^{163}Dy and ^{163}Ho mass reduced by a factor of 2
 - Know-how to be applied in SHIPTRAP

*Preparatory studies for a high-precision Penning trap measurement of the ^{163}Ho electron capture Q-value
F. Schneider et al., submitted to EPJ

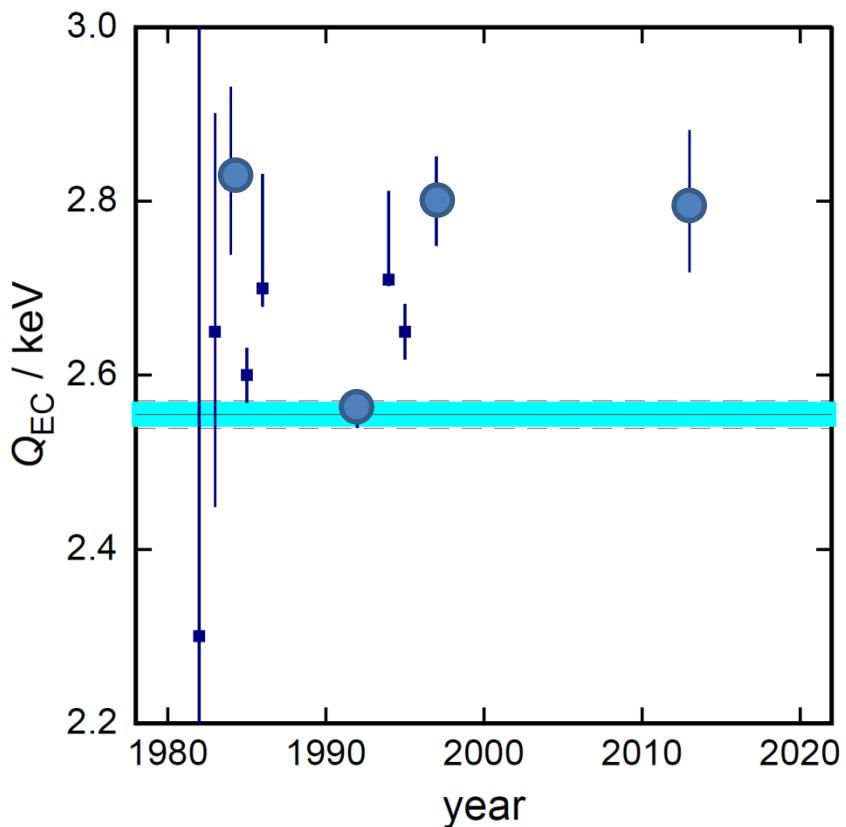


- Presently: [SHIPTRAP \(GSI\)](#) **
 - Q_{EC} determination with smaller uncertainties
 - Define scale of the experiment

** Direct measurement of the mass difference of ^{163}Ho and ^{163}Dy as prerequisite to a determination of the electron neutrino mass
S. Eliseev et al., submitted to PRL



Q_{EC} determination of ^{163}Ho

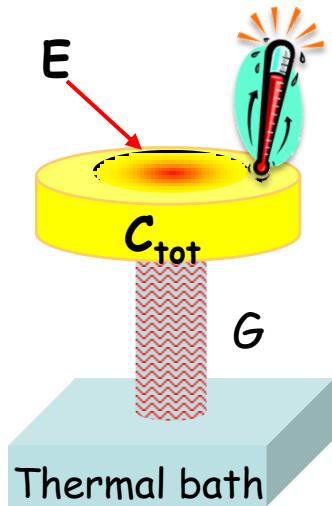


- Calorimetric measurements
- Other methods

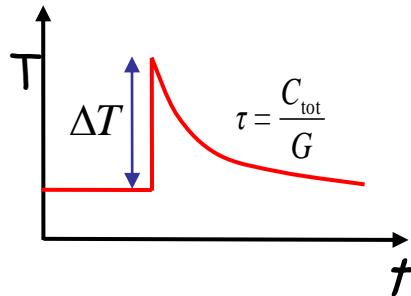
To reduce uncertainties in the analysis:
 Q_{EC} determination within **1 eV**
→ **PENTATRAP (MPIK HD)**



Low temperature micro-calorimeters



$$\Delta T \approx \frac{E}{C_{\text{tot}}}$$



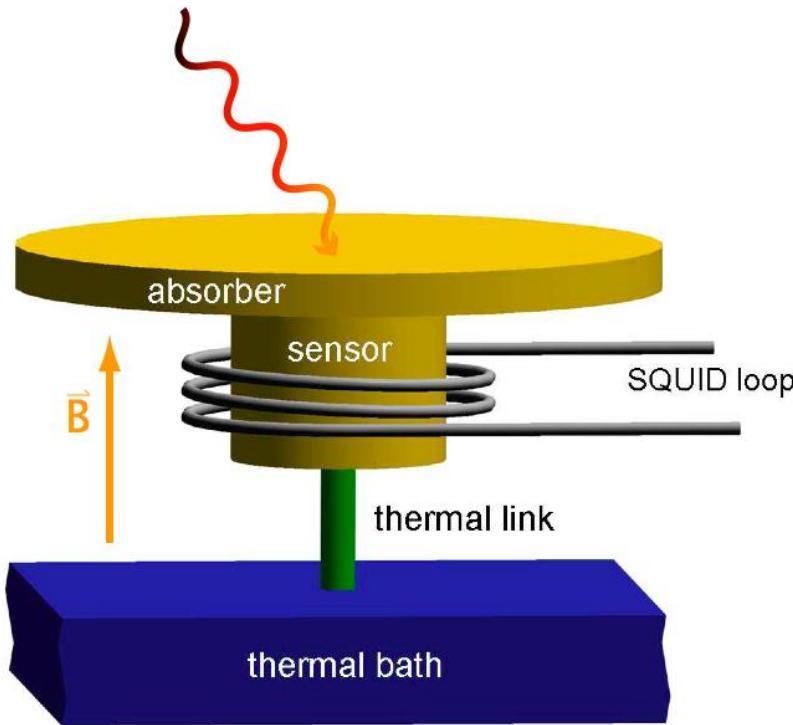
$$\left. \begin{array}{l} E = 10 \text{ keV} \\ C_{\text{tot}} = 1 \text{ pJ/K} \end{array} \right\} \rightarrow \sim 1 \text{ mK}$$

- Very small volume
- Working temperature below 100 mK
 - small specific heat
 - small thermal noise
- Very sensitive temperature sensor

Metallic Magnetic Calorimeters - MMC

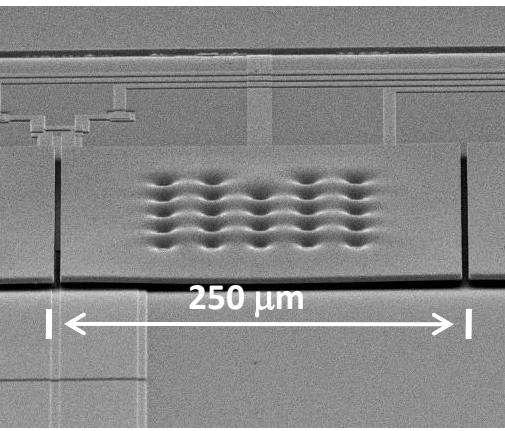
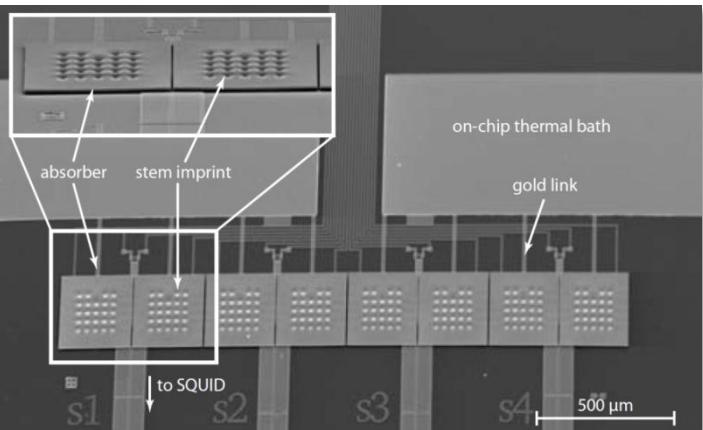
A. Fleischmann et al.,
AIP Conf. Proc. **1185**, 571, (2009)

- Paramagnetic sensor: Au:Er

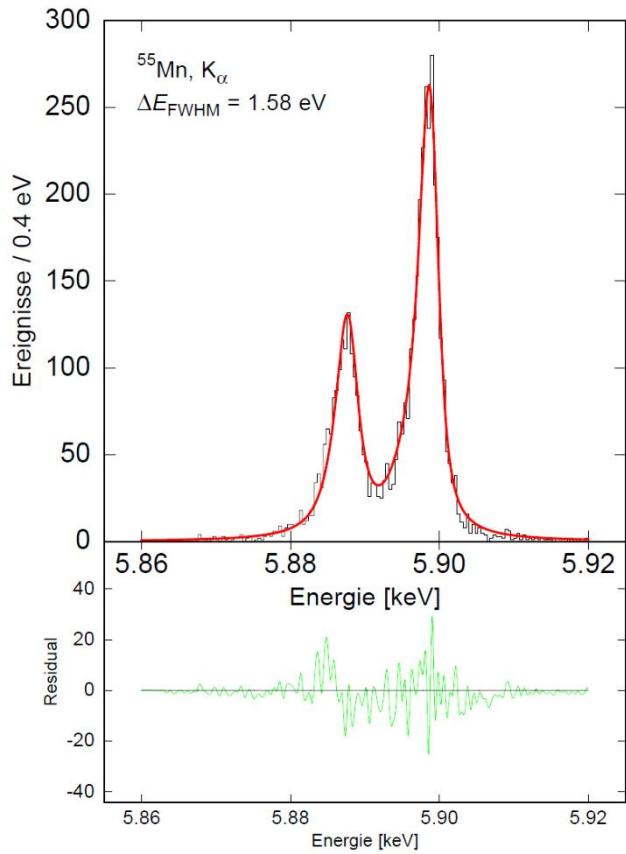


$$\Delta\Phi_s \propto \frac{\partial M}{\partial T} \Delta T \rightarrow \Delta\Phi_s \propto \frac{\partial M}{\partial T} \frac{E}{C_{\text{sens}} + C_{\text{abs}}}$$

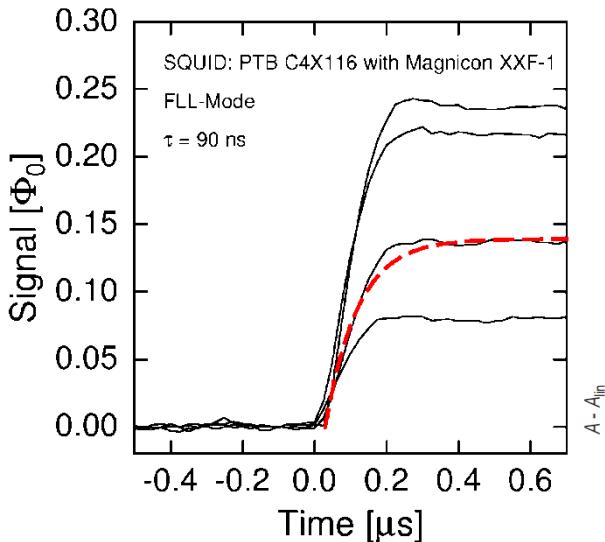
MMCs: 1d-array for soft x-rays ($T=20$ mK)



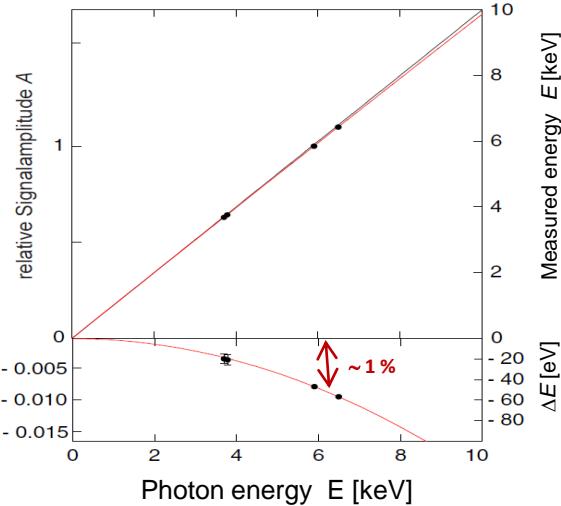
$\Delta E_{\text{FWHM}} = 1.6 \text{ eV} @ 6 \text{ keV}$



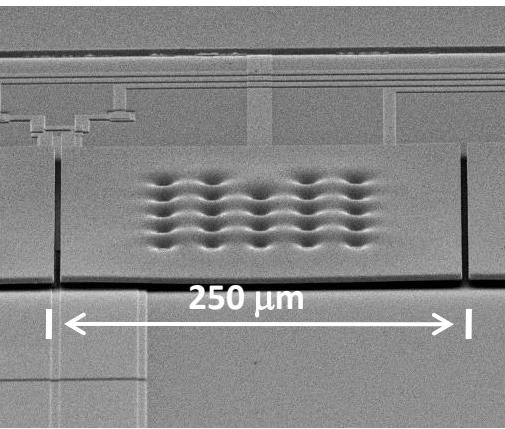
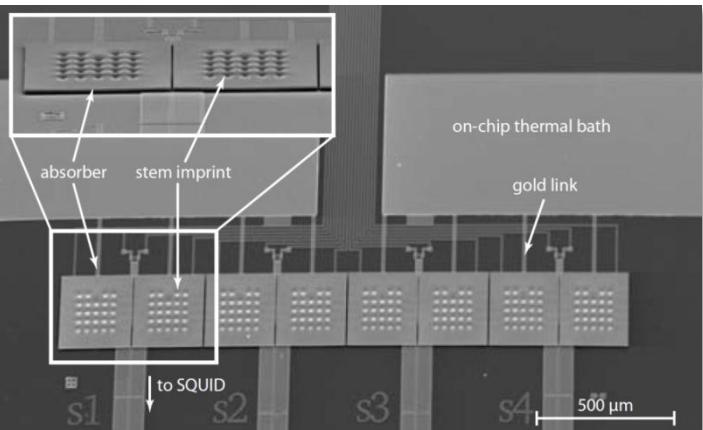
Rise Time: 90 ns



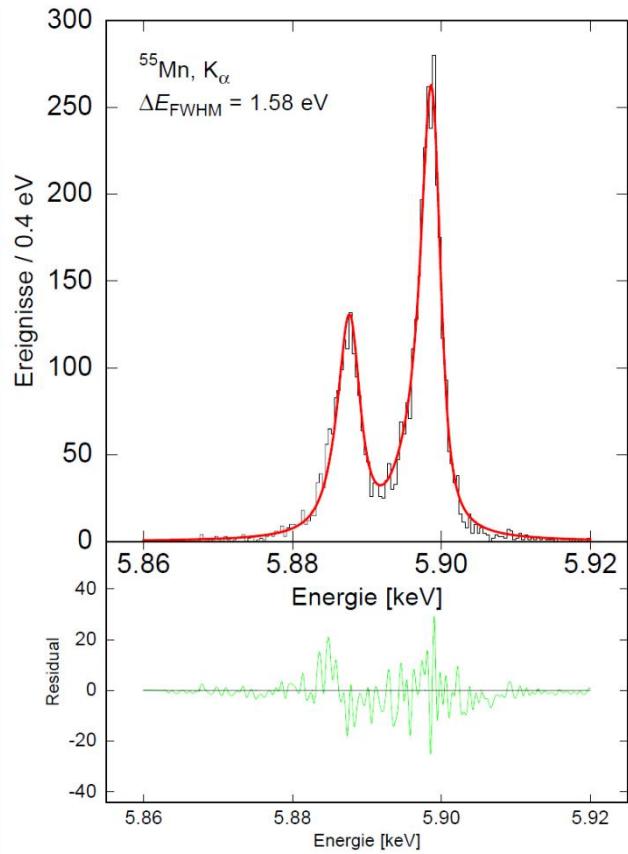
Non-Linearity < 1% @6keV



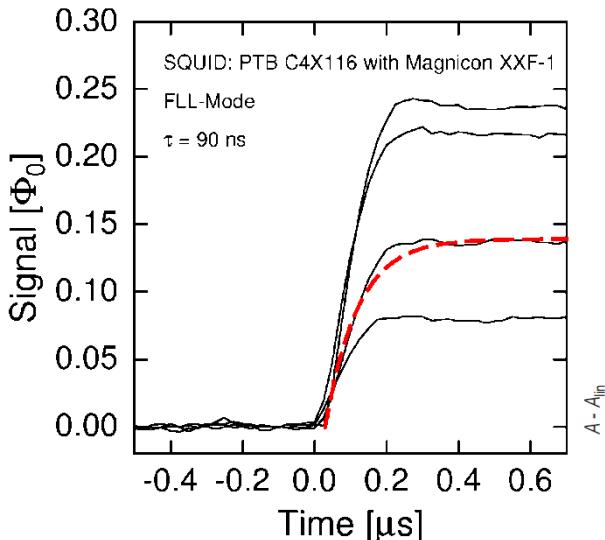
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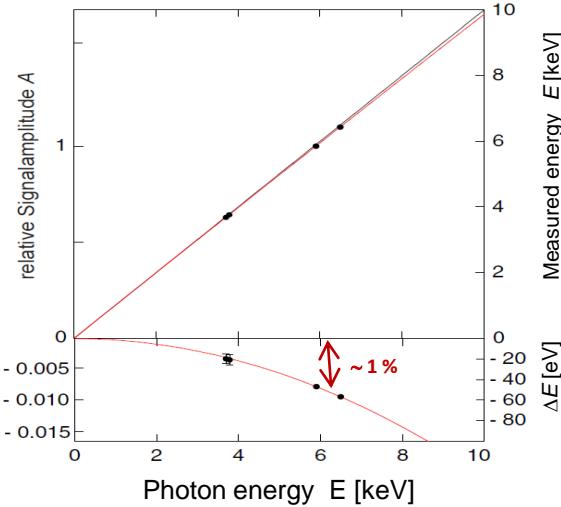
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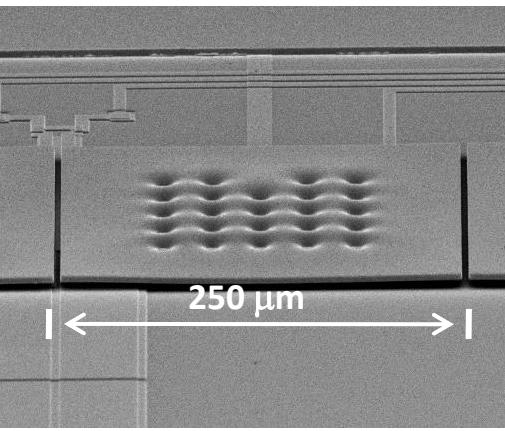
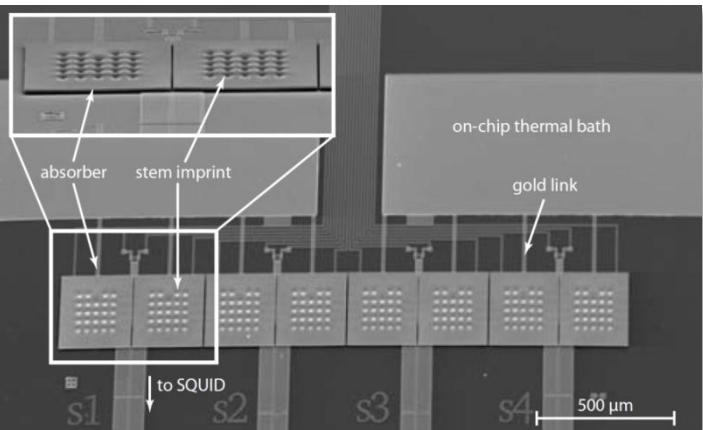


Non-Linearity < 1% @6keV

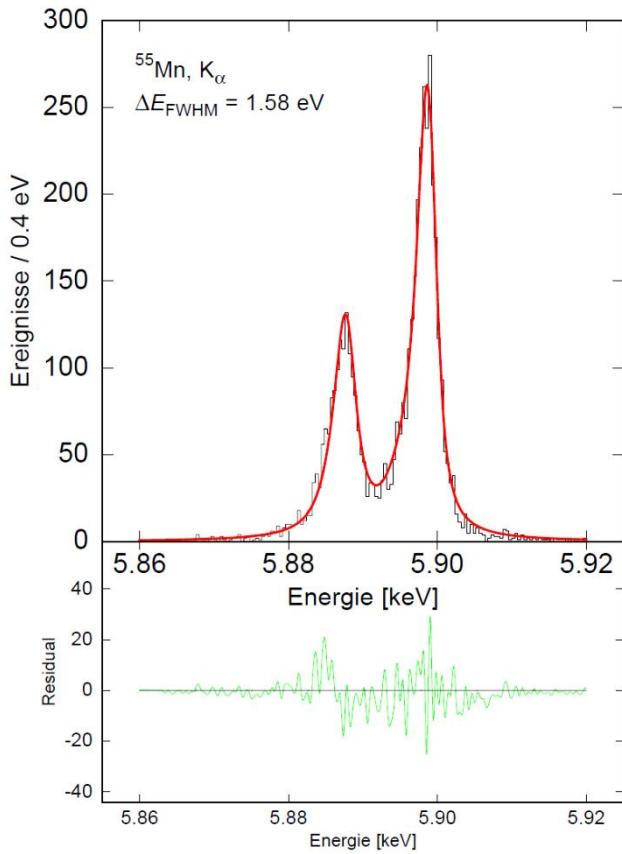


Reduction
un-resolved pile-up

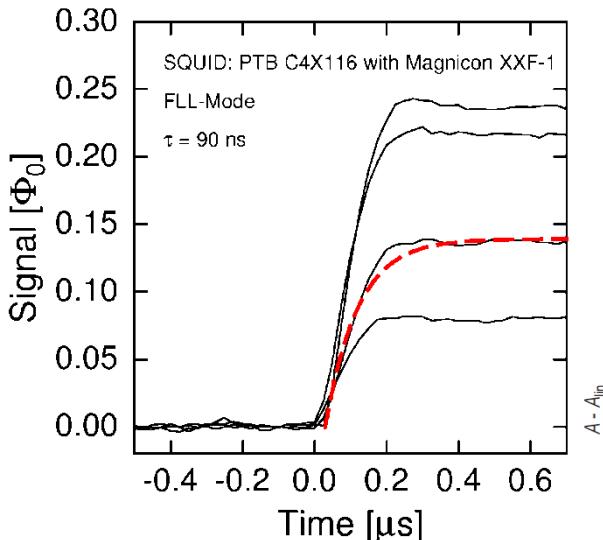
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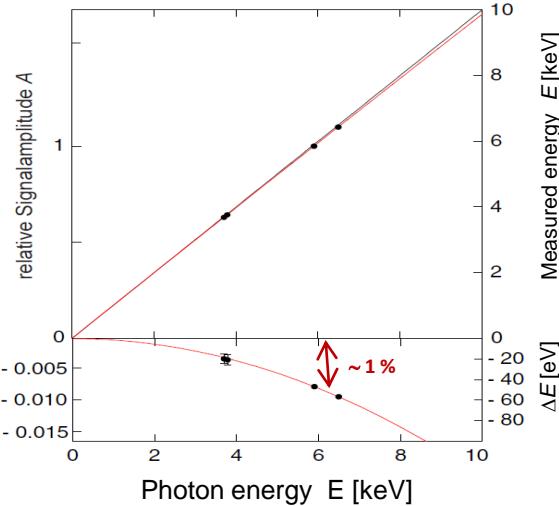
$$\Delta E_{\text{FWHM}} = 1.6 \text{ eV} @ 6 \text{ keV}$$



Rise Time: 90 ns



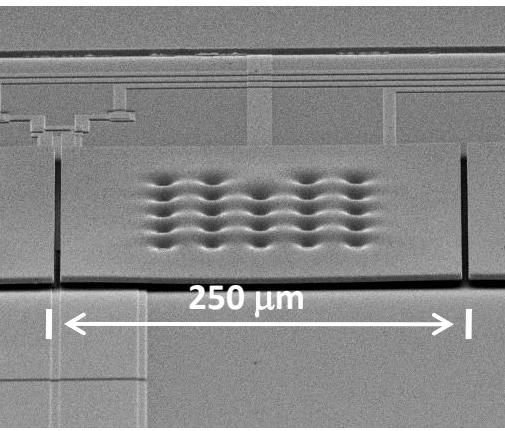
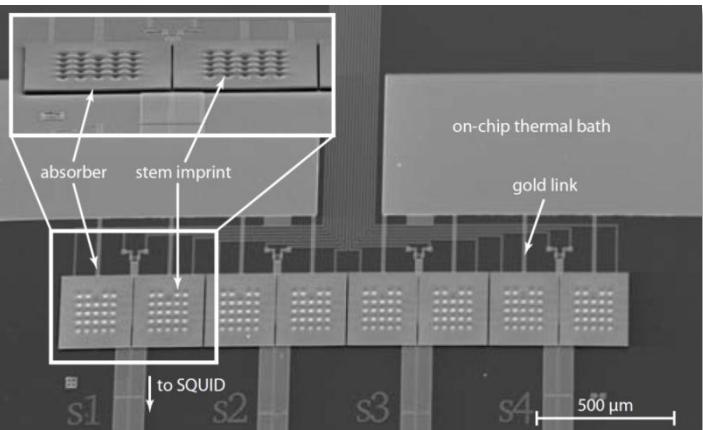
Non-Linearity < 1% @6keV



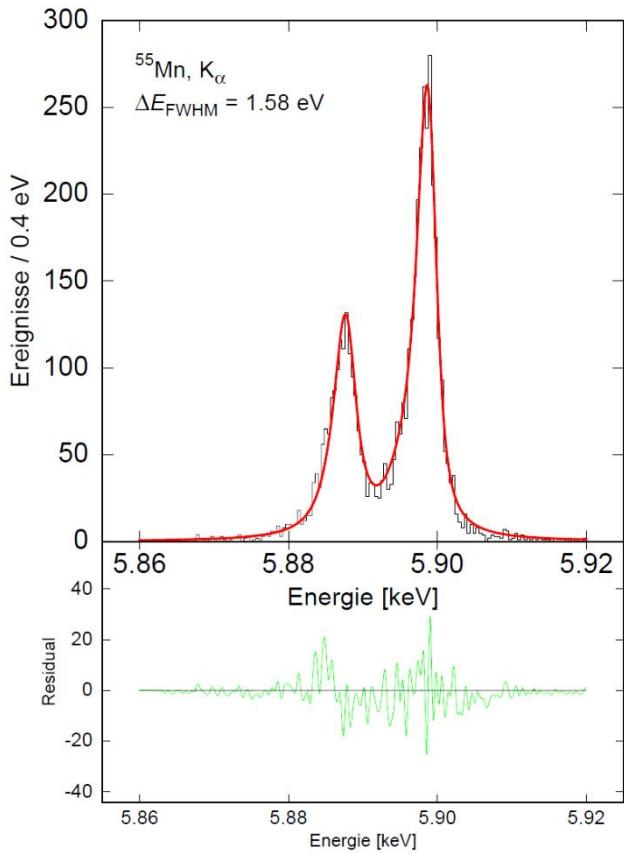
Reduction
un-resolved pile-up

Definition
of the energy scale

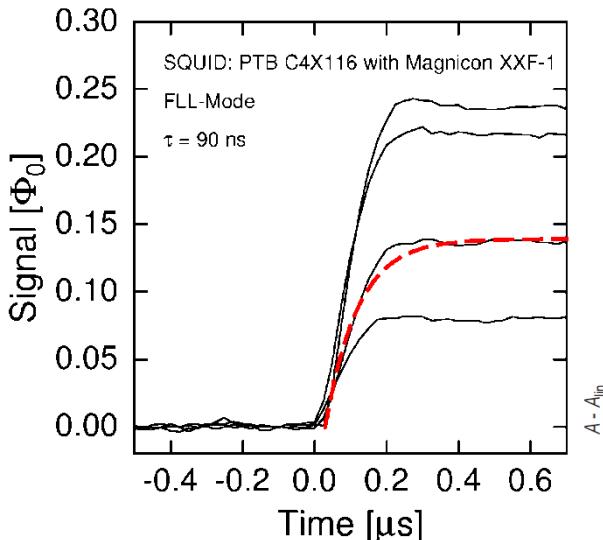
MMCs: 1d-array for soft x-rays ($T=20$ mK)



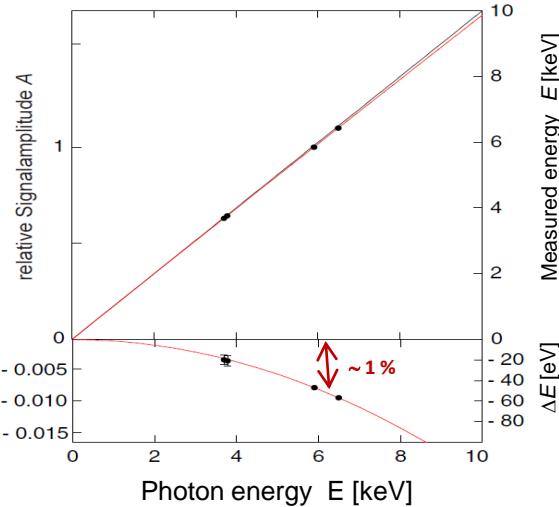
$$\Delta E_{\text{FWHM}} = 1.6 \text{ eV} @ 6 \text{ keV}$$



Rise Time: 90 ns



Non-Linearity < 1% @6keV



Reduction
un-resolved pile-up

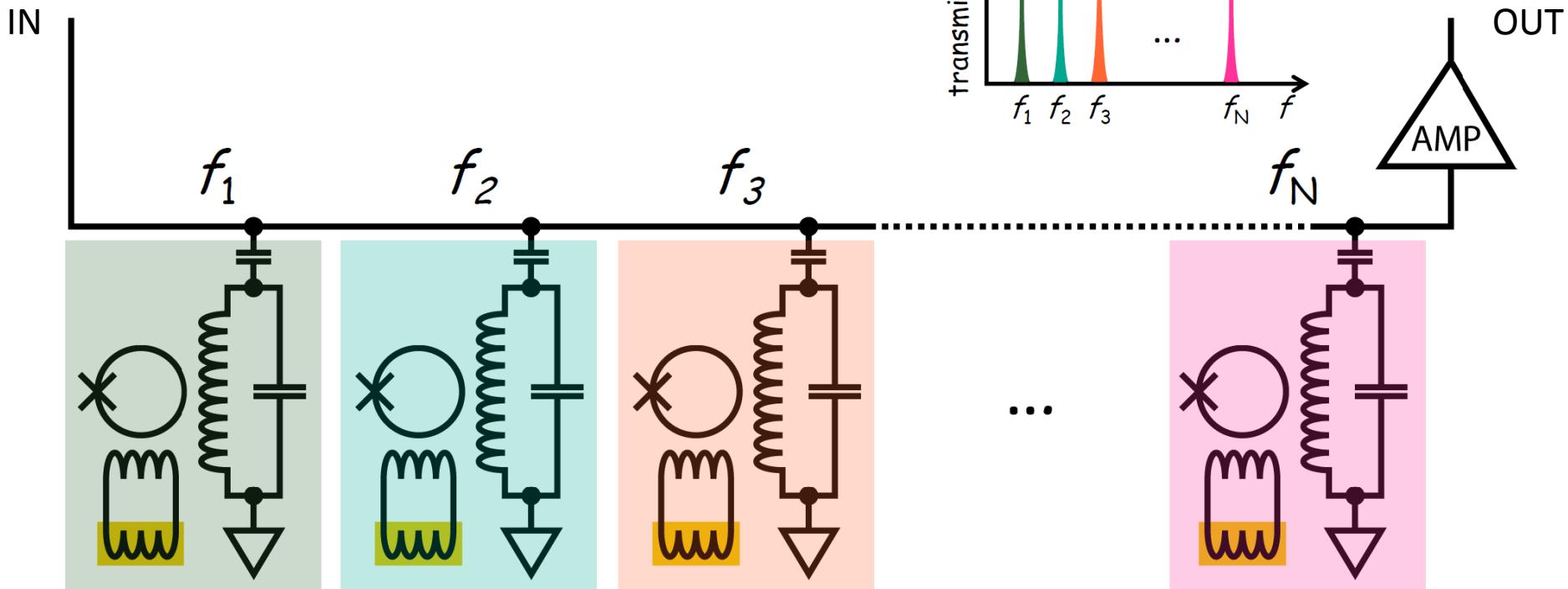
Definition
of the energy scale

Reduced smearing
in the end point region

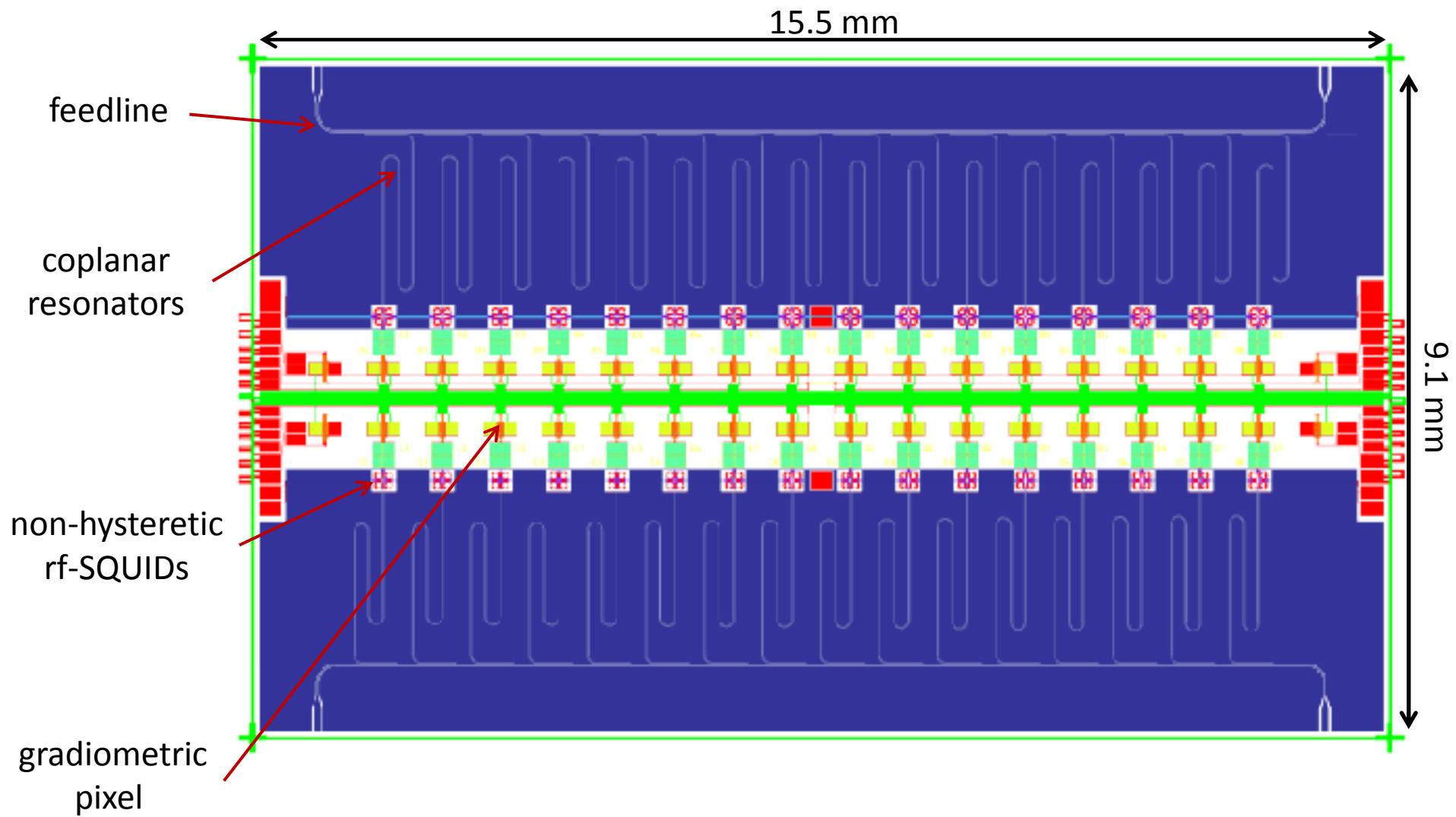
MMCs: From single pixels to arrays

Microwave SQUID multiplexing

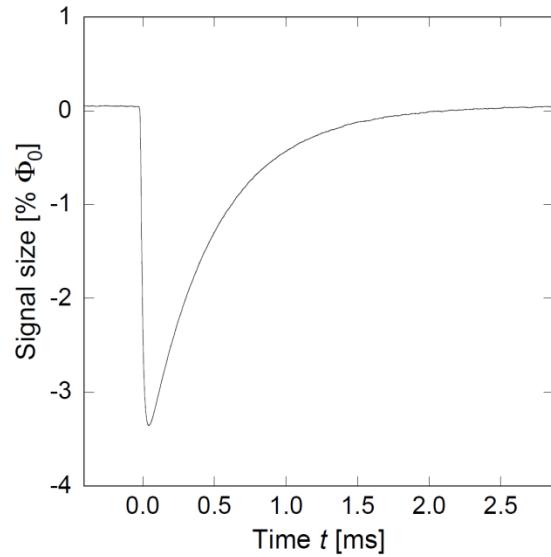
single HEMT amplifier and 2 coaxes
to read out 100 - 1000 detectors



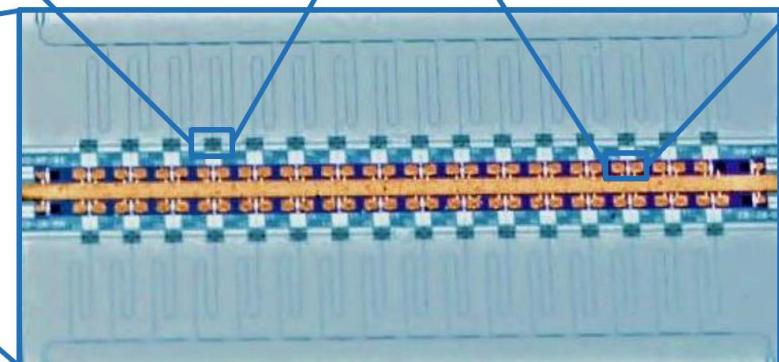
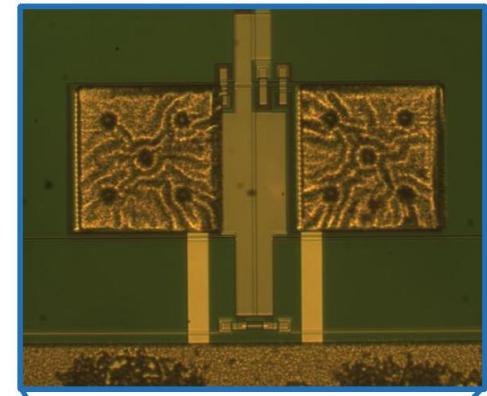
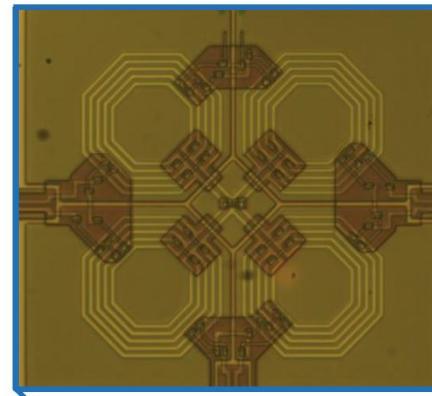
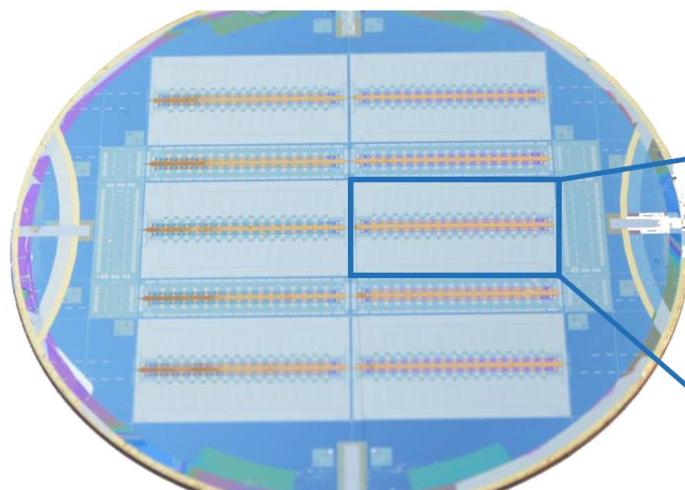
MMCs: 64 pixel array



MMCs: Microwave SQUID multiplexing

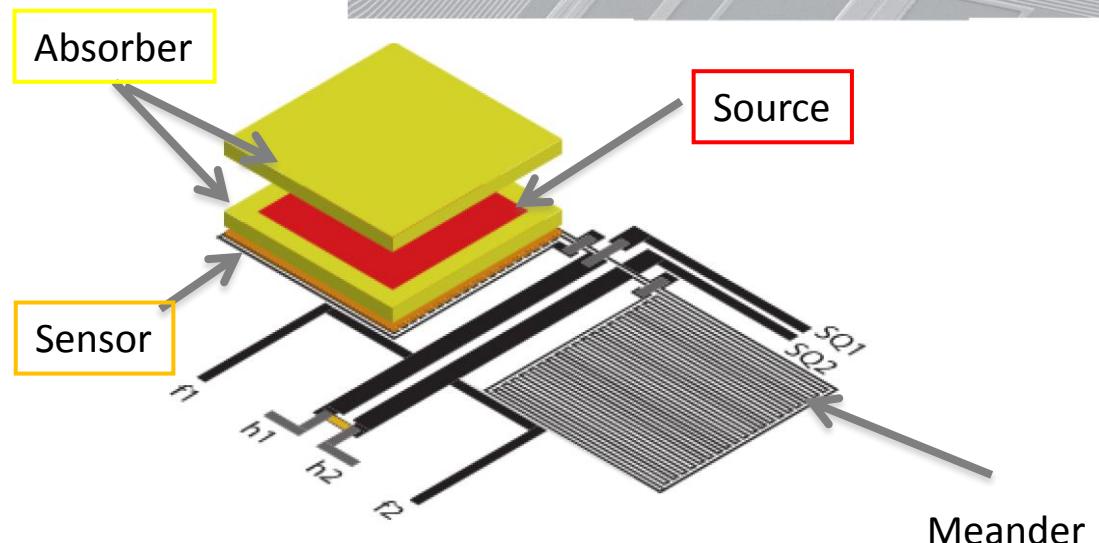
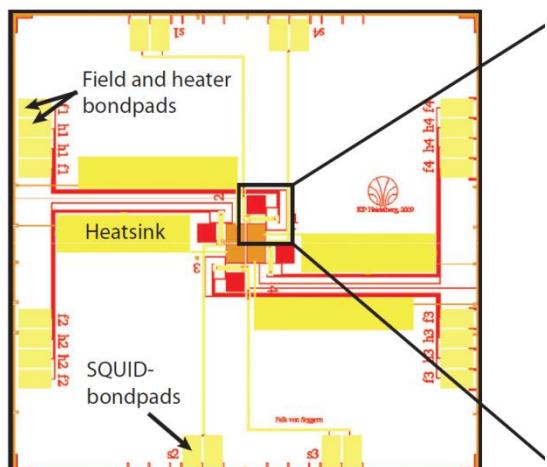
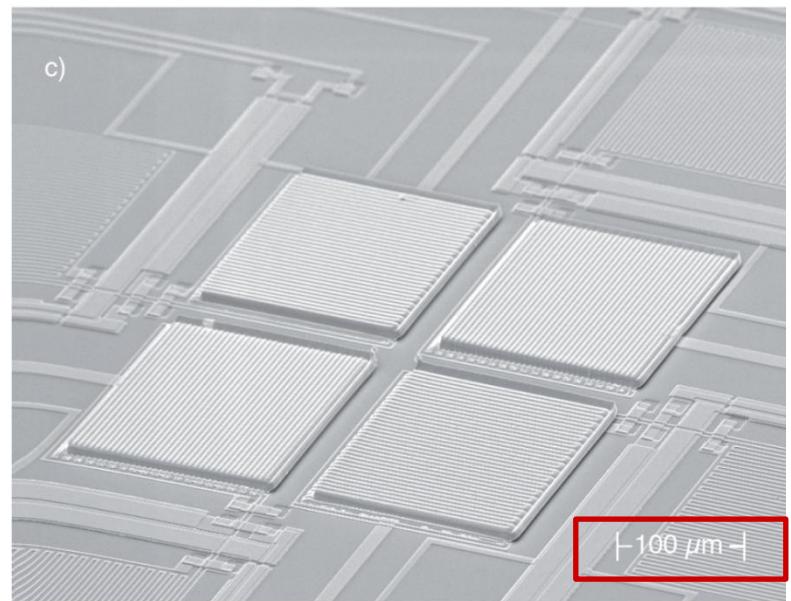


Successful production and test of the first prototype



First detector prototype

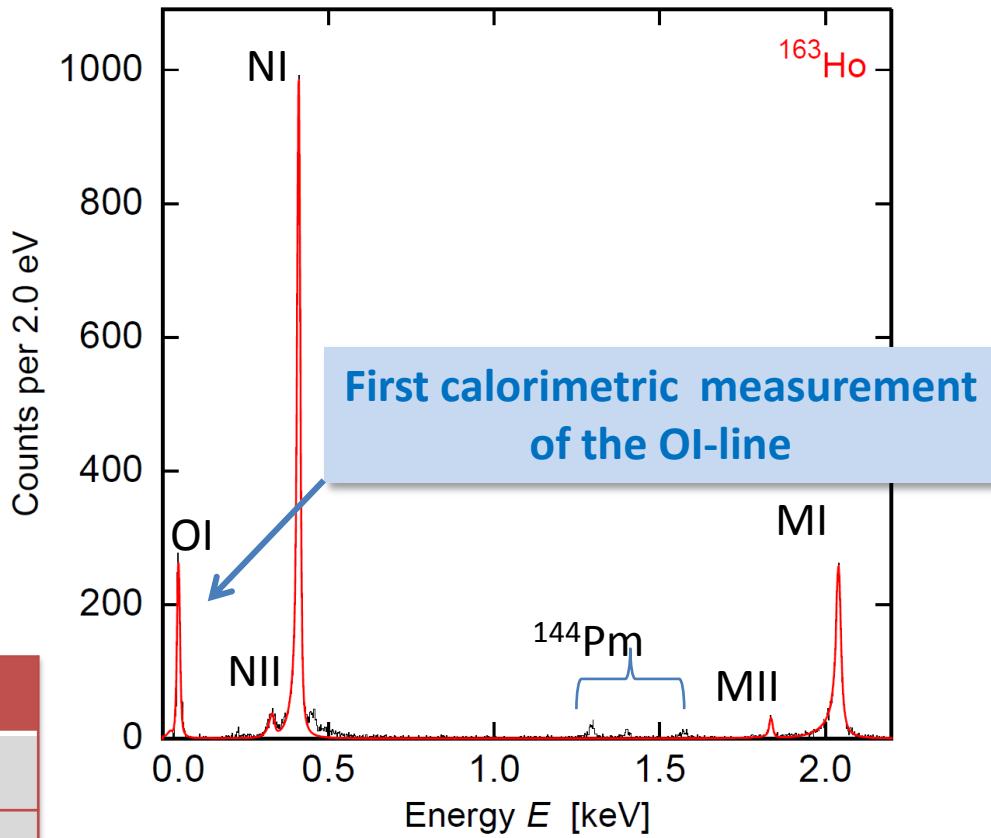
- Absorber for calorimetric measurement
→ ion implantation @ ISOLDE-CERN
- About 0.01 Bq per pixel
- Two pixels have been simultaneously measured



Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV (2013)
 $\Delta E_{\text{FWHM}} = 2.4$ eV @ 0 keV (2014)
- Non-Linearity < 1% @ 6keV
- Synchronized measurement of 2 pixels
- Presently most precise ^{163}Ho spectrum

	E_{H} bind.	E_{H} exp.	Γ_{H} lit.	Γ_{H} exp
MI	2.047	2.040	13.2	13.7
MII	1.845	1.836	6.0	7.2
NI	0.420	0.411	5.4	5.3
NII	0.340	0.333	5.3	8.0
OI	0.050	0.048	5.0	4.3



$$Q_{\text{EC}} = (2.843 \pm 0.009^{\text{stat}} - 0.06^{\text{syst}}) \text{ keV}$$

Where to improve

Detector design and fabrication:

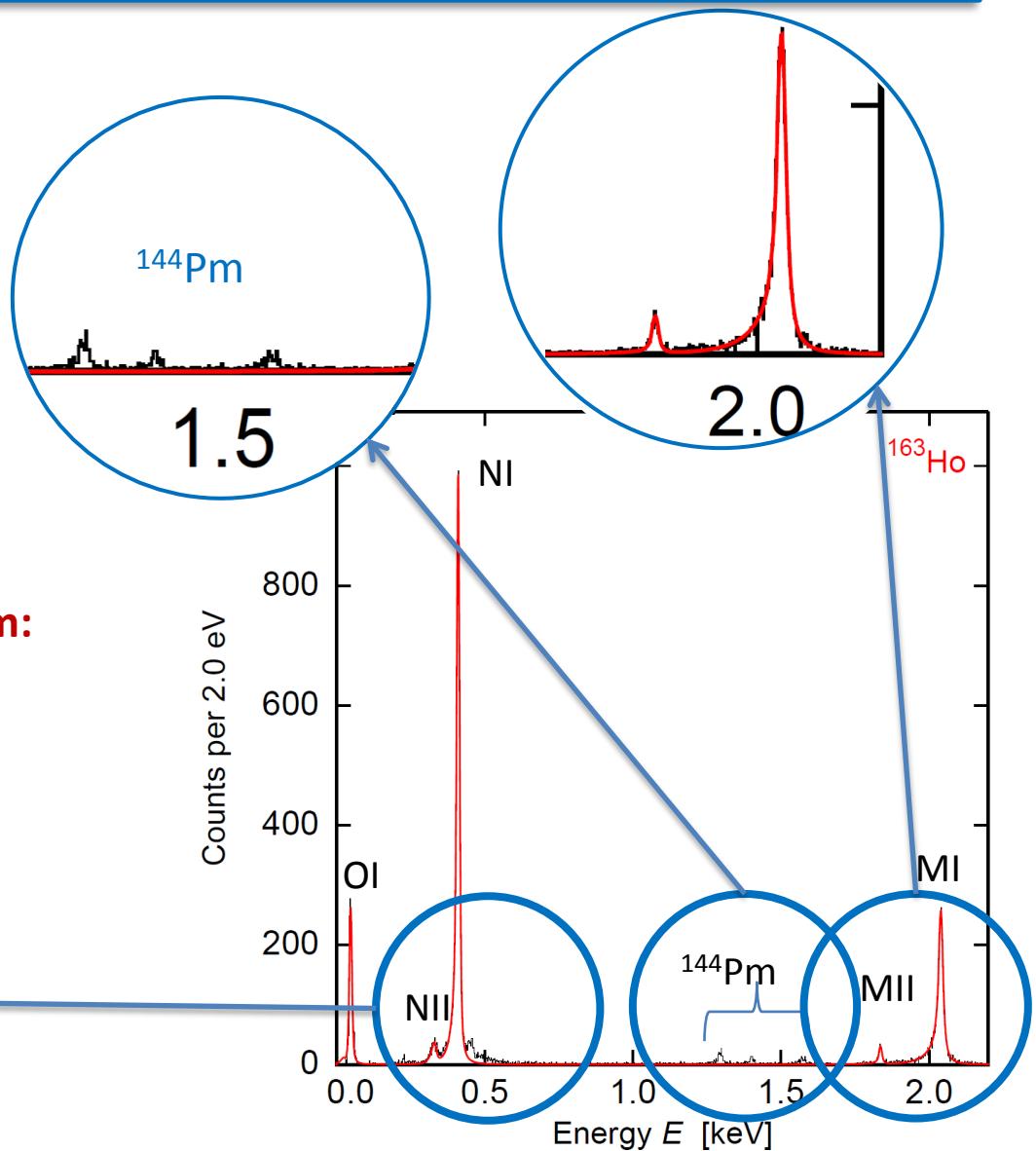
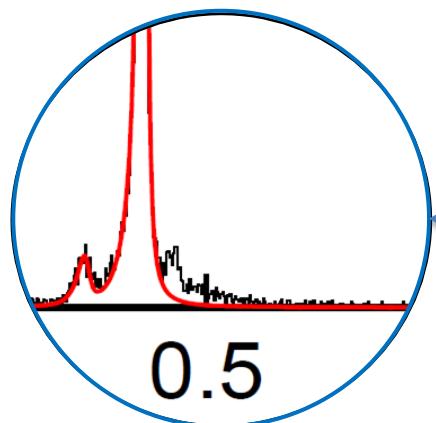
- Increase activity per pixel
- Remove low energy tail

High purity ^{163}Ho source:

- Background reduction

Understanding of the ^{163}Ho spectrum:

- Investigate undefined structures



High purity ^{163}Ho source

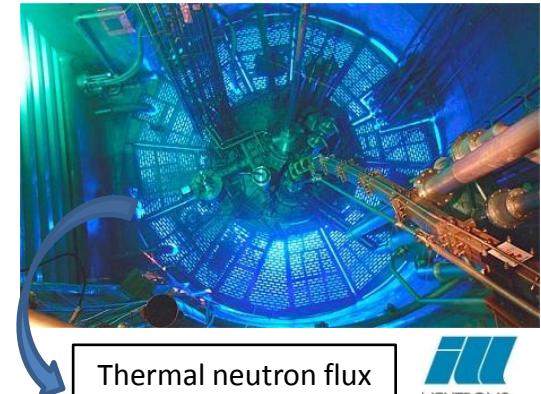
(n, γ)-reaction on 30% enriched ^{162}Er

γ spectrum of the 30 mg sample after chemical separation:

⇒ only $^{166\text{m}}\text{Ho}$ visible

Mass separation

- Use of the RISIKO mass-separator@Uni-Mainz
→ First successful test with ^{165}Ho
- Separation at CERN/ISOLDE December 2014
→ new chips to test!



Thermal neutron flux
(Φ): $1.3 \times 10^{15} \text{ cm}^{-2}\text{s}^{-1}$



Goal: $^{166\text{m}}\text{Ho}/^{163}\text{Ho} \leq 10^{-9}$

maXs-20: 16 pixel detector arrays for soft x-rays

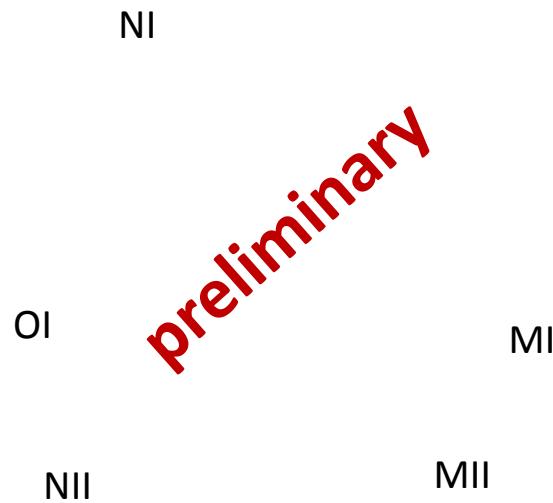


- higher activity per pixel $\approx 1 \text{ Bq}$
- no radioactive contaminants
- better energy resolution

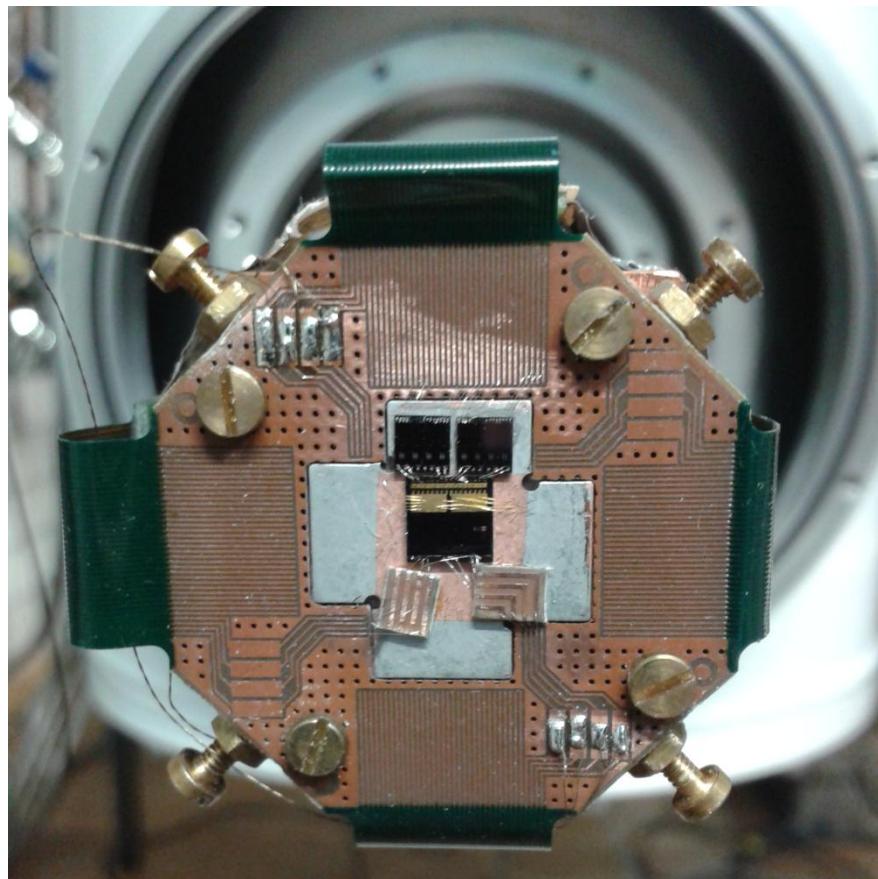


New detector ready for first tests....

At 18 mK since last Wednesday.....

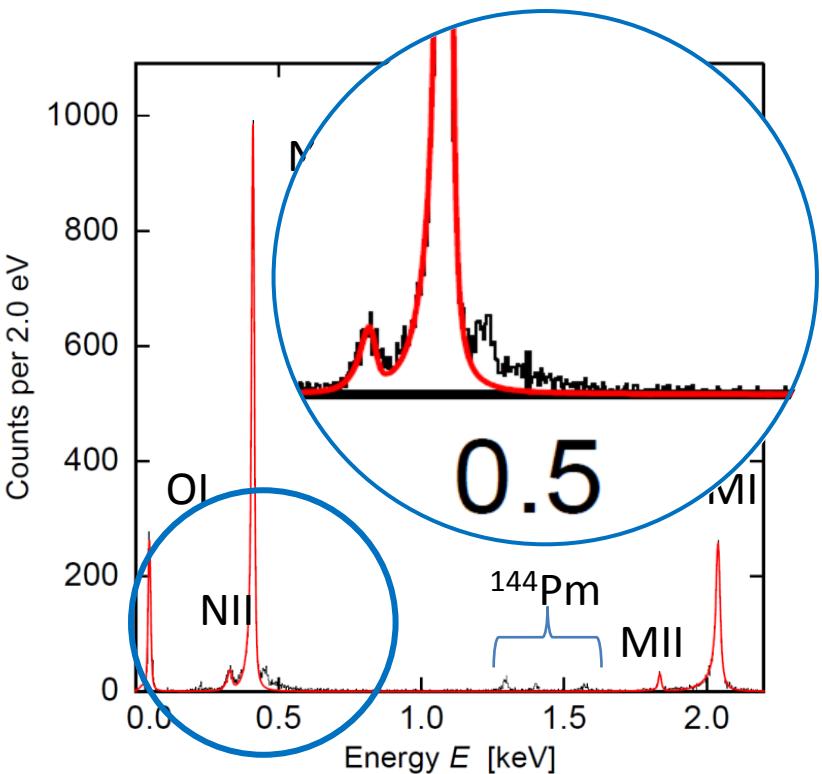


Mounted on a cold arm of a dry cryostat



- Activity per pixel : $A \sim 0.2 \text{ Bq}$
- Baseline resolution : $\Delta E_{\text{FWHM}} = 5 \text{ eV}$
- No evidence of strong radioactive contamination in the source

Characterisation of spectral shape



Estimate the effect of

- Higher order excitation in ^{163}Dy
- ^{163}Ho ion embedded in Au

A. Faessler et al.

J. Phys. G **42** (2015) 015108

R. G. H. Robertson

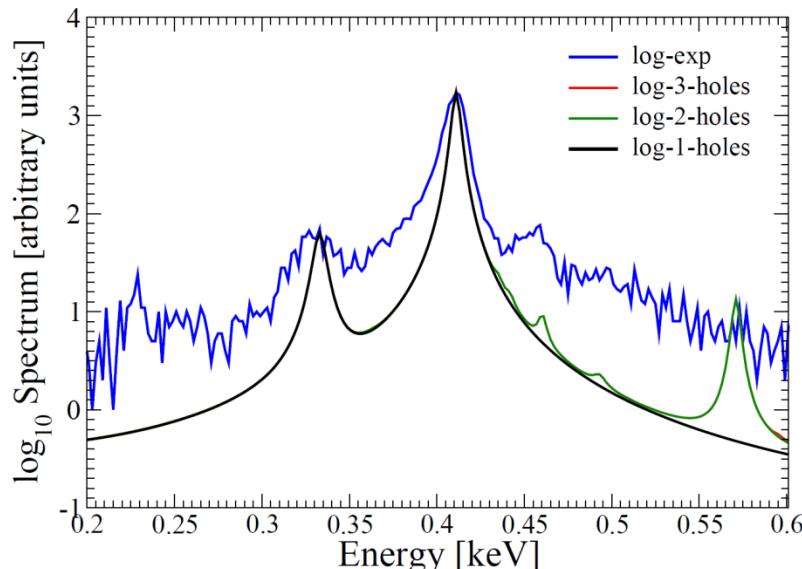
Phys. Rev. C **91**, 035504 (2015)

A. Faessler et al.

Phys. Rev. C **91**, 045505 (2015)

A. Faessler et al.

Phys. Rev. C **91**, 064302 (2015)

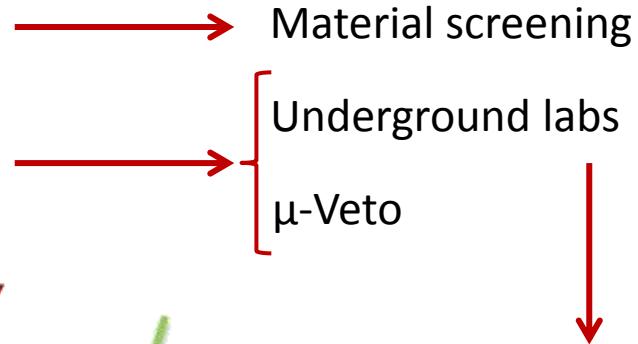
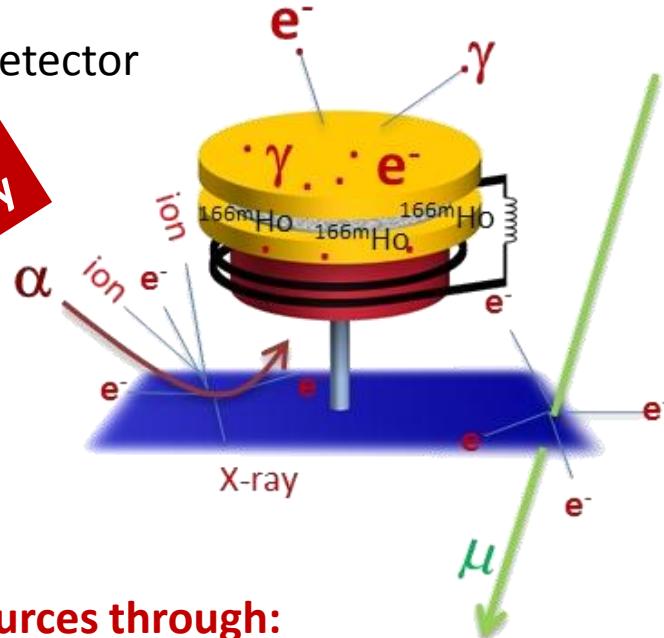


Background

Background sources:

- Environmental radioactivity
- Cosmic rays
- Induced secondary radiation by cosmic rays
- Radioactivity in the detector

Background level
 5×10^{-5} counts/eV/det/day



Study of background sources through:

- Monte Carlo simulations
- Dedicated experiments

Screening facilities

- Uni-Tübingen
- Felsenkeller

ECHO overview

- Prove **scalability** with medium large experiment **ECHO-1k**

- $A \sim 1000 \text{ Bq}$
- $\Delta E_{\text{FWHM}} < 5 \text{ eV}$
- $\tau_r < 1 \mu\text{s}$
- multiplexed arrays → microwave SQUID multiplexing
- 1 year measuring time → 10^{10} counts = Neutrino mass sensitivity $m_\nu < 10 \text{ eV}$

High purity ^{163}Ho source (produced at reactor)

Just approved

Research Unit FOR 2202/1

„Neutrino Mass Determination by Electron Capture in Holmium-163 – ECHO“



Deutsche
Forschungsgemeinschaft

- **ECHO-1M** towards sub-eV sensitivity

Sterile Neutrino and ^{163}Ho

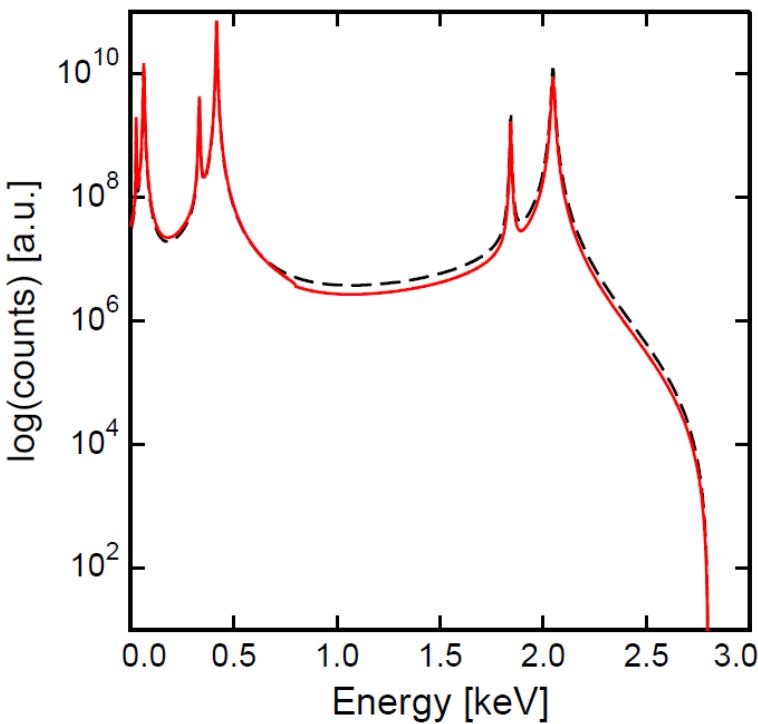


Sterile Neutrino and ^{163}Ho

How does
the existence of sterile neutrino
affect the EC spectrum?

Sterile Neutrino and ^{163}Ho

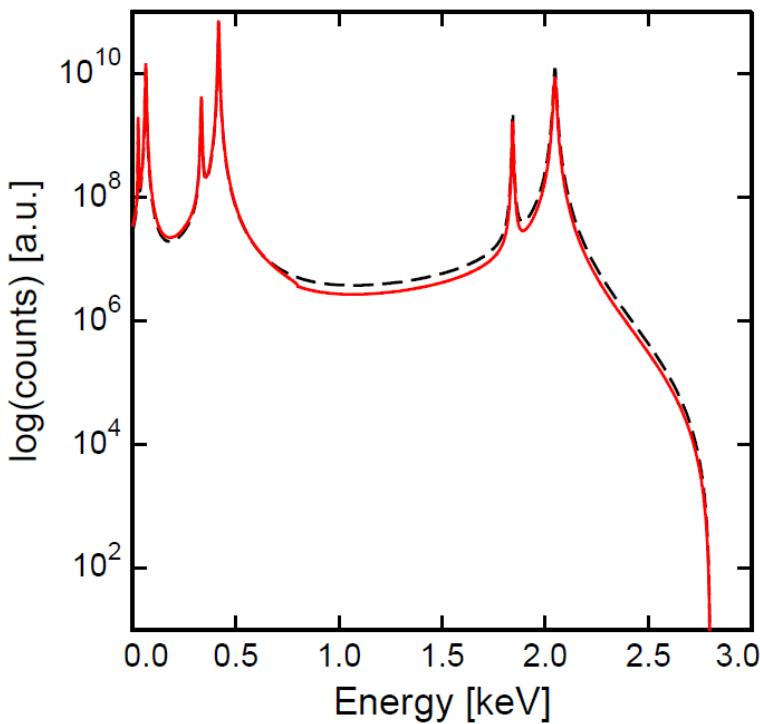
$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



Sterile Neutrino and ^{163}Ho

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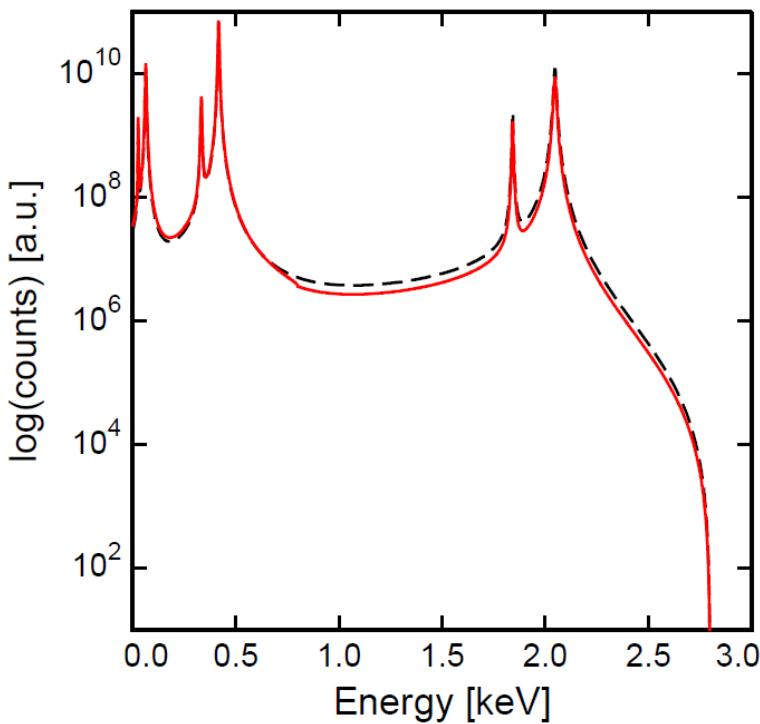


Sterile Neutrino and ^{163}Ho

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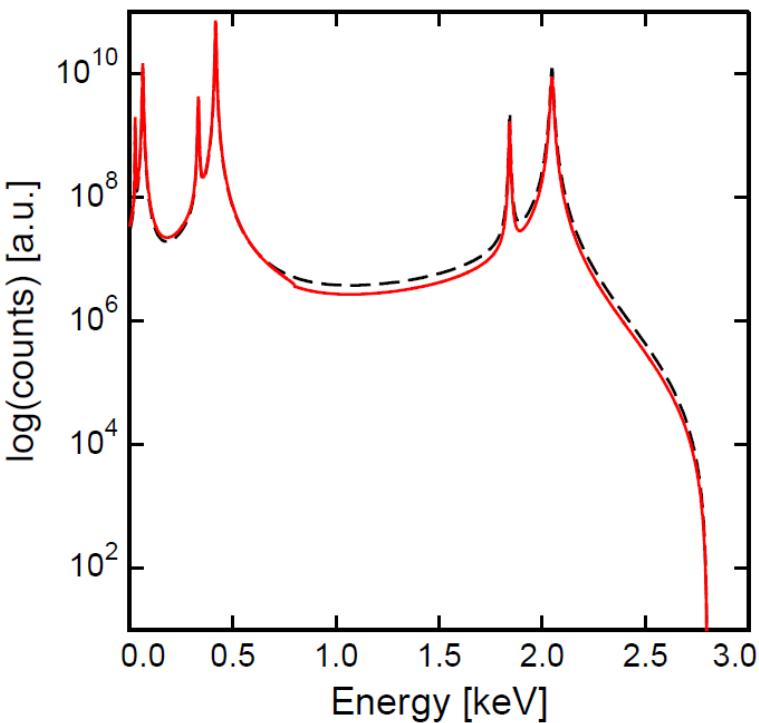


Sterile Neutrino and ^{163}Ho

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Assume: $m_a = 0$

$m_s \neq 0$

$\Rightarrow m_{1,2,3} = 0$

$m_4 \neq 0$

$m_4 \approx \text{keV}$

Sterile Neutrino and ^{163}Ho

$$\frac{dW}{dE_C} = A(Q_{EC} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{EC} - E_C)^2}} \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$

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$$|\nu_e\rangle = \sum_{i=1}^3 U_{ei} |\nu_i\rangle + U_{e4} |\nu_4\rangle$$

$$\frac{dW}{dE_C} = A(Q_{EC} - E_C)^2 \left[\left(1 - |U_{e4}|^2 \right) + |U_{e4}|^2 \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_C)^2}} H(Q_{EC} - E_c - m_4) \right] \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$

Sterile Neutrino and ^{163}Ho

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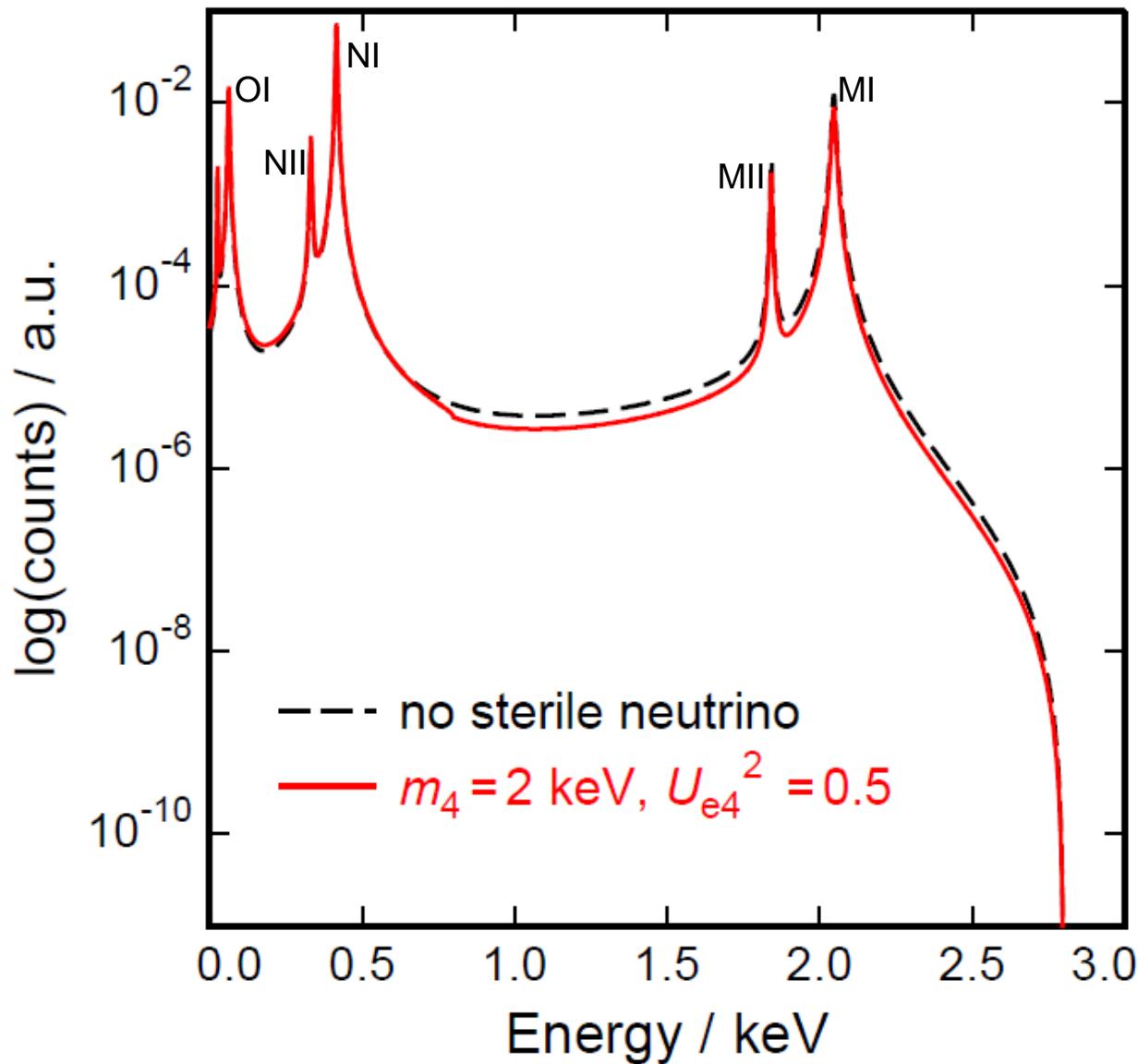
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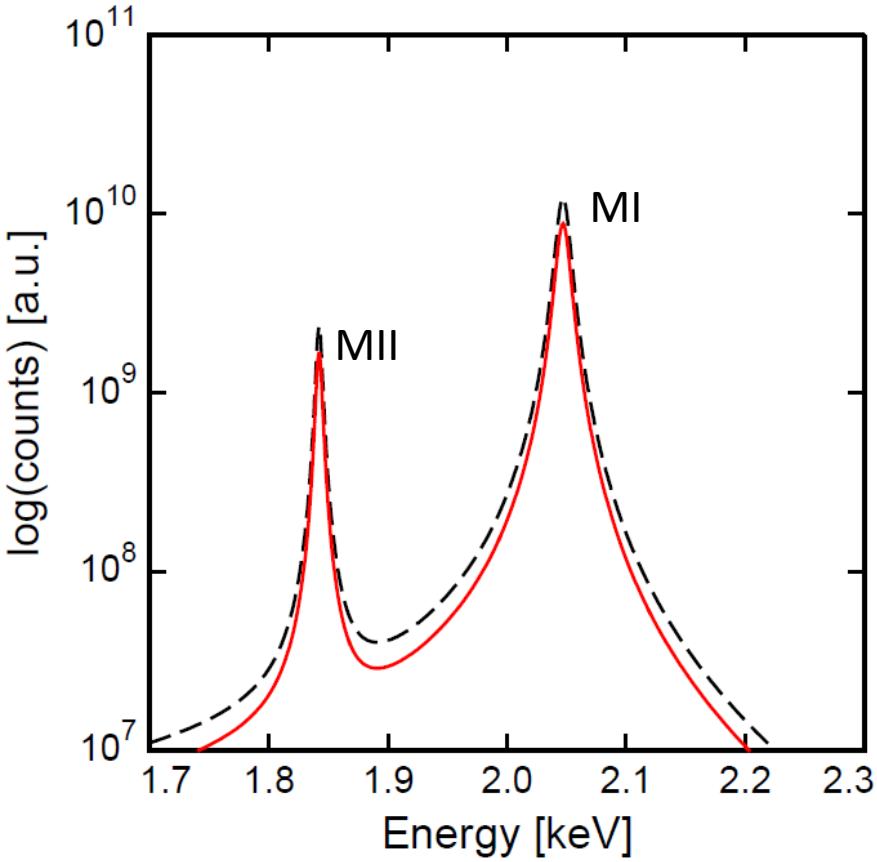
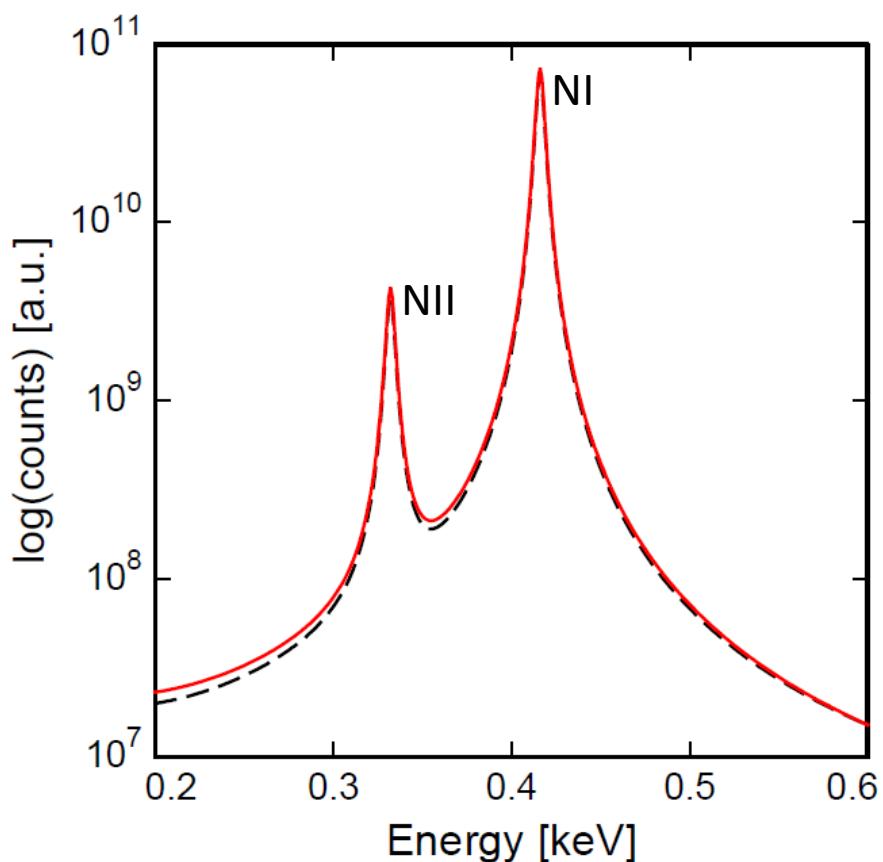
No contribution for $Q_{EC} < m_4$

Sterile Neutrino and ^{163}Ho



Sterile Neutrino and ^{163}Ho

$m_4=2 \text{ keV}$, $U_{e4}^2=0.5$
no sterile neutrino



Sterile Neutrino and ^{163}Ho

- Amplitude of the line H for only active neutrinos

$$W_{Ha} = A(Q_{EC} - E_H)^2 B_H \varphi_H^2(0)$$

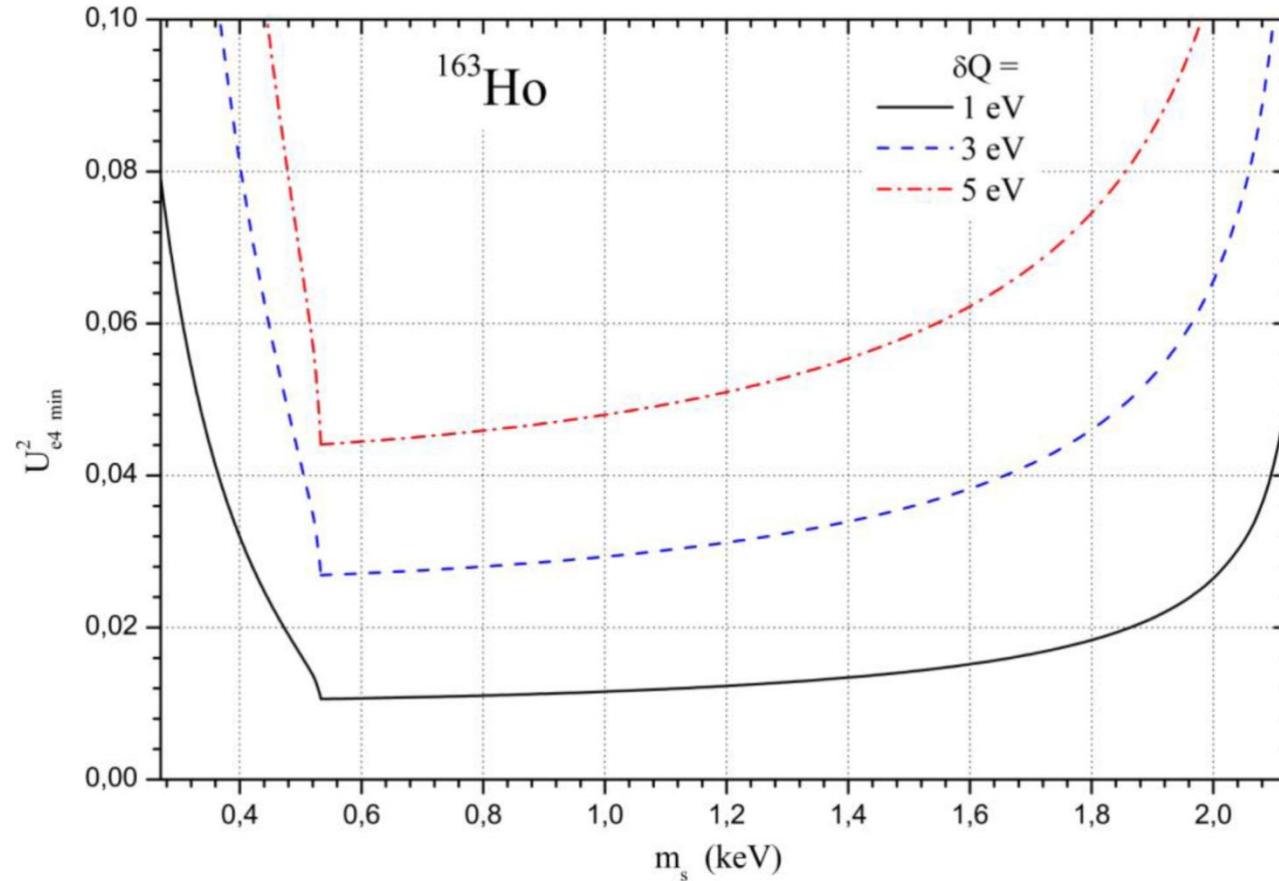
- Amplitude of the line H for 3+1 model in case of $m_a = 0 \text{ eV}$

$$W_{Hs} = A(Q_{EC} - E_H)^2 \left[\left(1 - |U_{e4}|^2 \right) + |U_{e4}|^2 \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_c)^2}} H(Q_{EC} - E_c - m_4) \right] B_H \varphi_H^2(0)$$

- Ratio between amplitudes of two lines in the spectrum for 3+1 model in case of $m_a = 0 \text{ eV}$

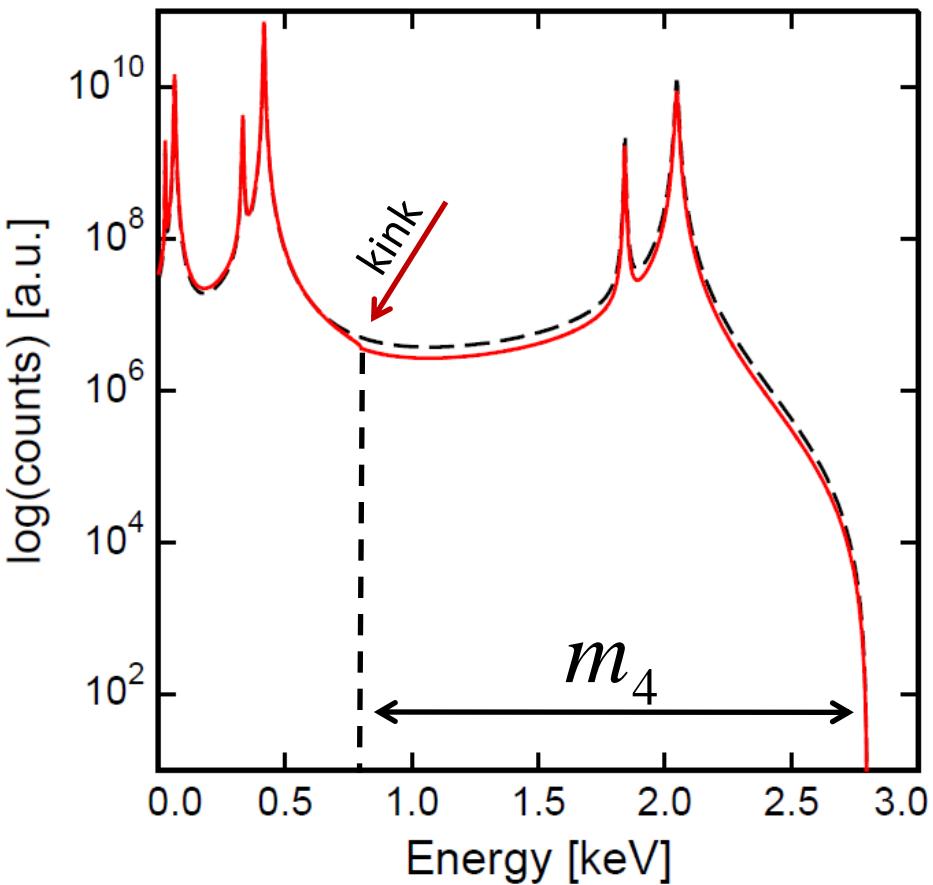
$$\left(\frac{W_{H1}}{W_{H2}} \right)_s = \left(\frac{W_{H1}}{W_{H2}} \right)_a \frac{|U_{e4}|^2 \left[H(Q_{EC} - E_1 - m_4) \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_1)^2}} - 1 \right] + 1}{|U_{e4}|^2 \left[H(Q_{EC} - E_2 - m_4) \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_2)^2}} - 1 \right] + 1}$$

Sterile Neutrino and ^{163}Ho

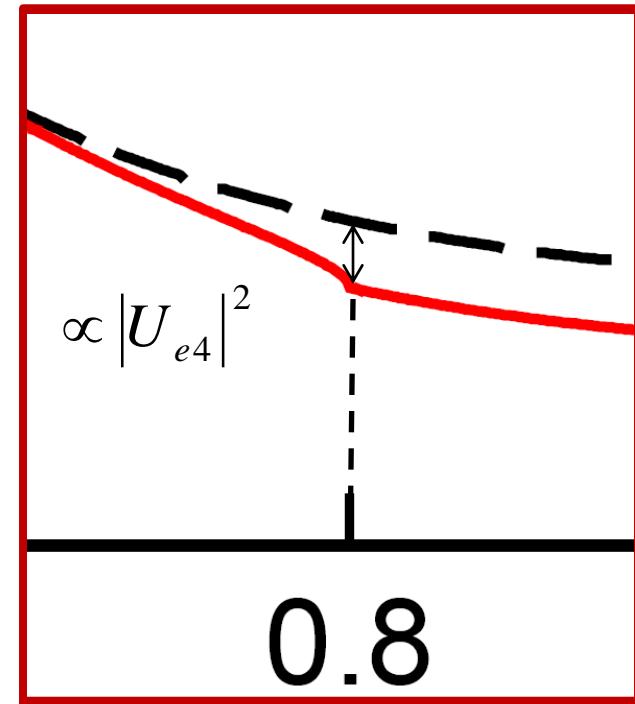


Sensitivity to the mixing matrix element at 90% CL as a function of the sterile neutrino mass achievable with about 10^{10} events in the full EC spectrum.

Sterile Neutrino and ^{163}Ho

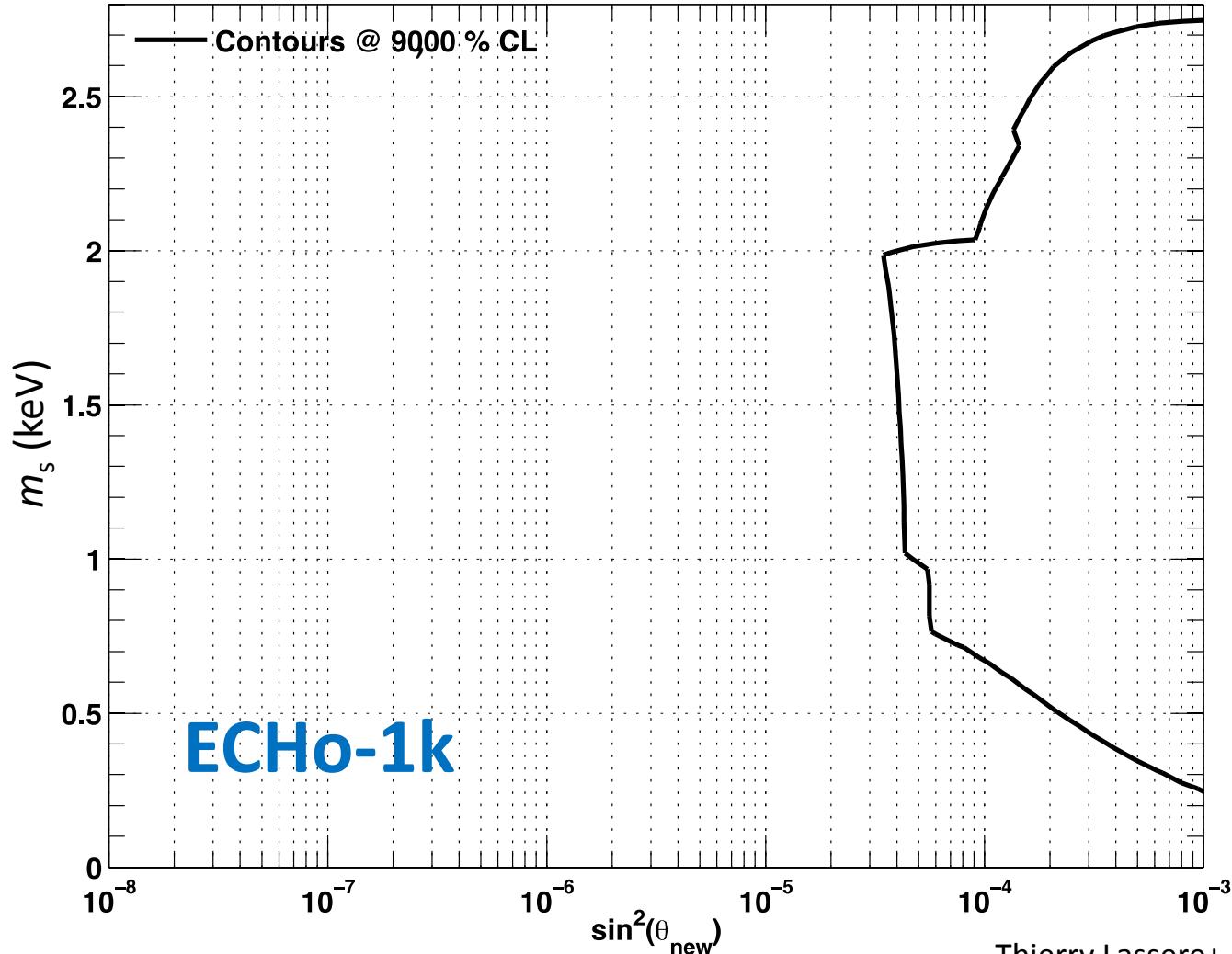


- position of kink => m_4
- depth of kink => $|U_{e4}|^2$



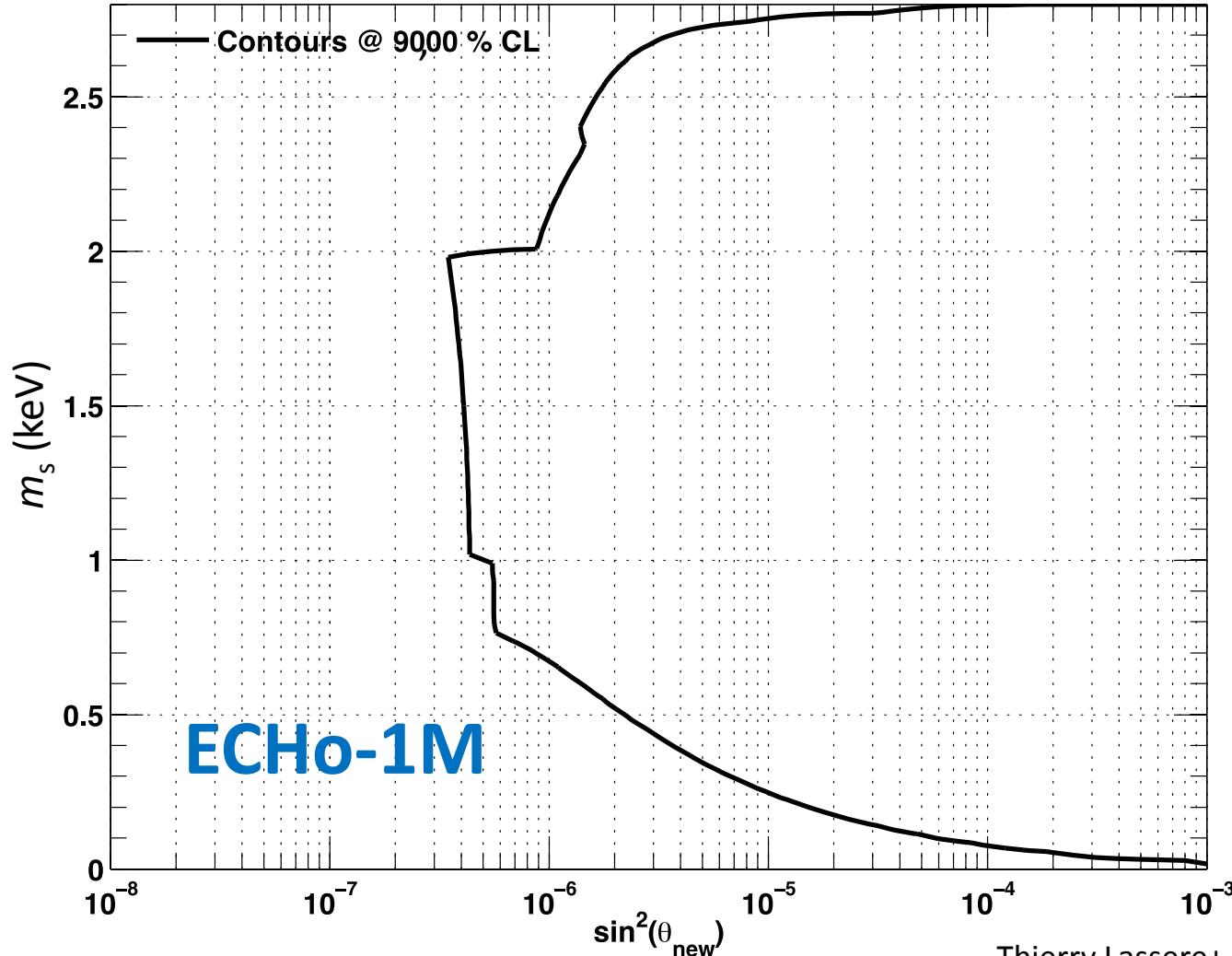
Sterile Neutrino and ^{163}Ho

Statistical Fluctuation – No Pile Up – Counts = 1e10
Theoretical Spectrum Supposed to be perfectly known

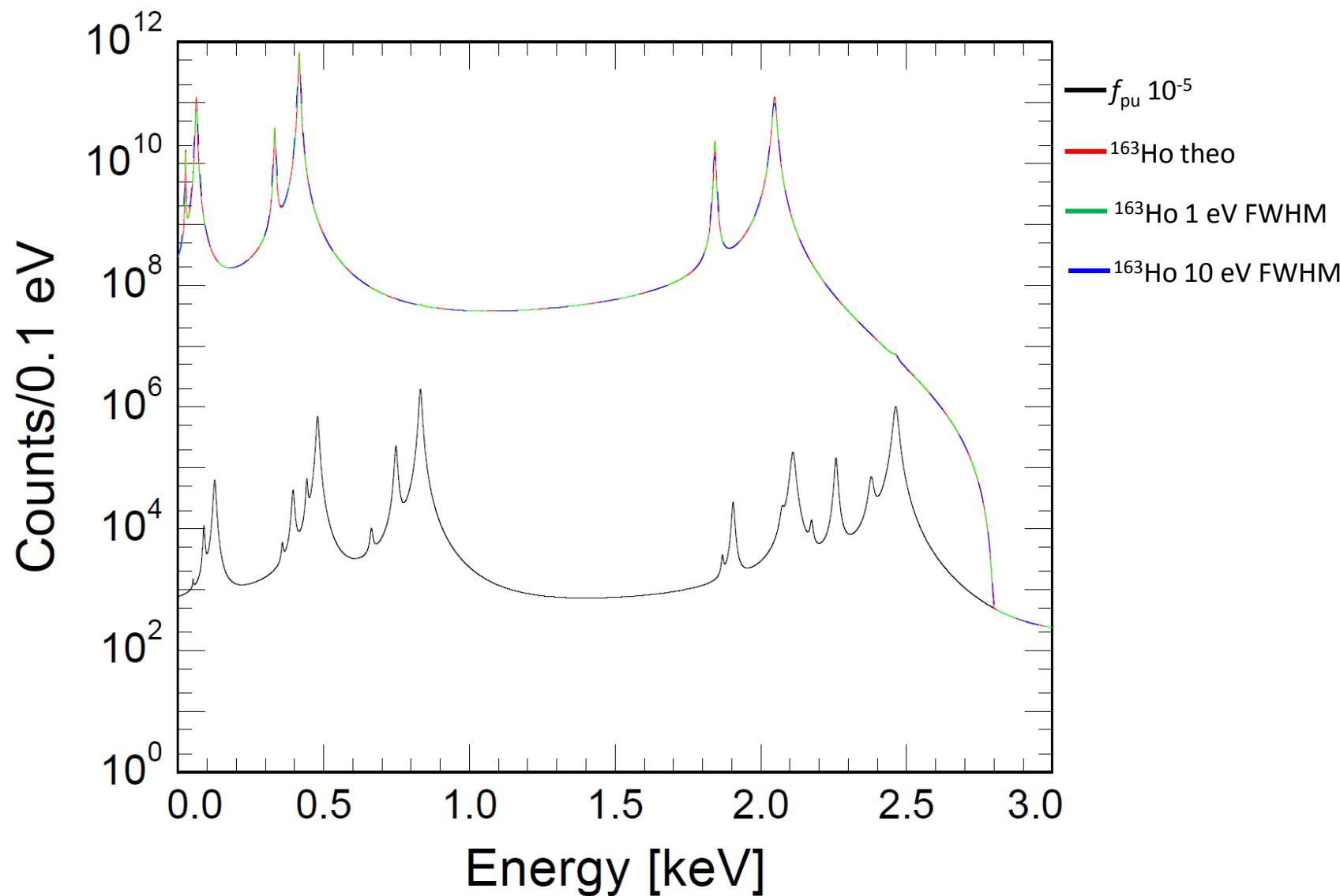


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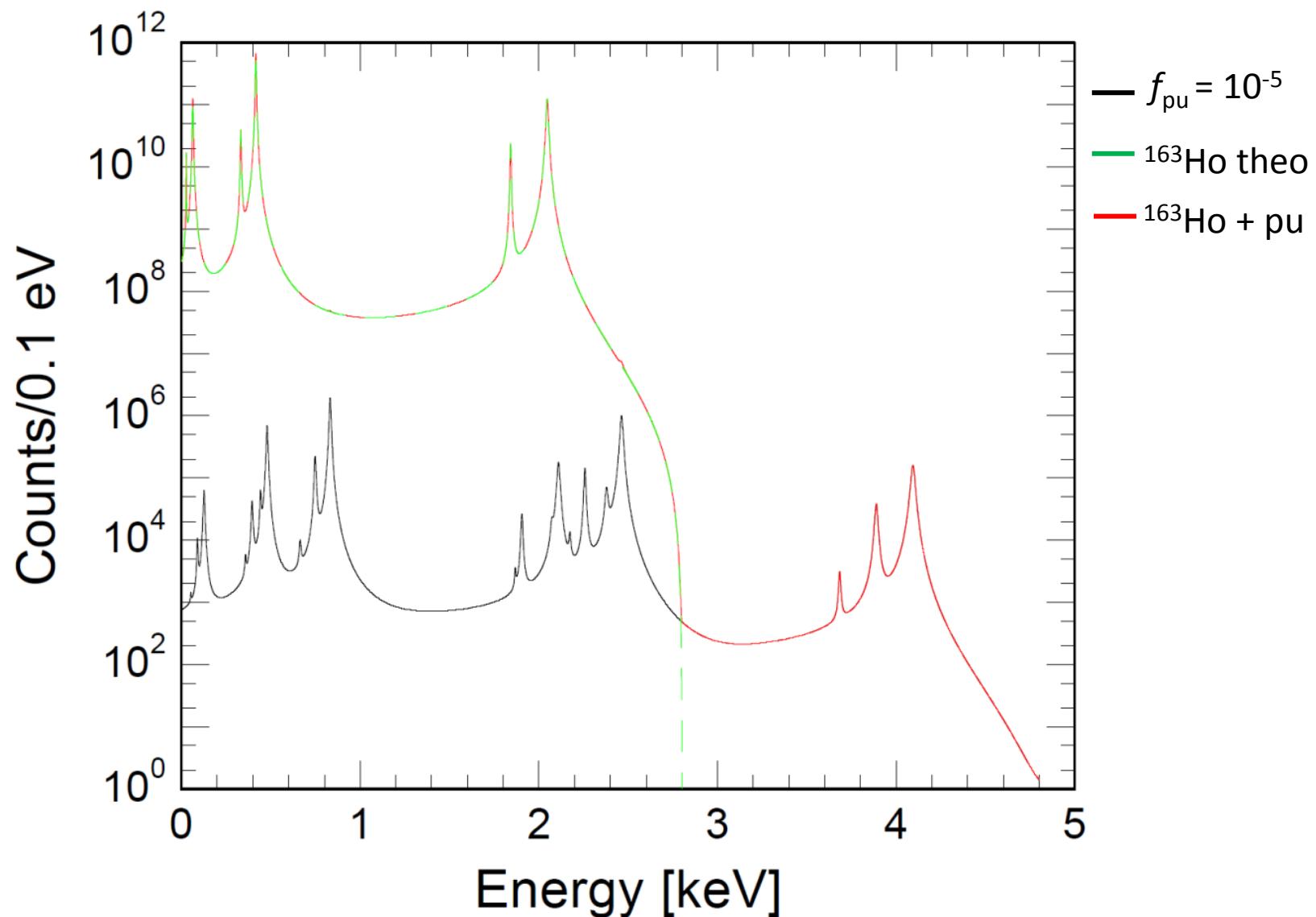
Statistical Fluctuation – No Pile Up – Counts = 1e14
Theoretical Spectrum Supposed to be perfectly known



Pile-up and energy resolution



Pile-up and energy resolution



Sterile Neutrino (keV) and Electron Capture

Other candidates in the EC branch:

- $Q_{\text{EC}} < 100 \text{ keV}$
- Reasonable halflife

Nuclide	$T_{1/2}$	EC-transition	Q (keV) [22]	B_i (keV) [23]	B_j (keV) [23]	$ \psi_i ^2/ \psi_j ^2$	$Q-B_i$ (keV)
^{123}Te	$>2 \cdot 10^{15} \text{ y}$?	52.7(16)	K: 30.4912(3)	L _I : 4.9392(3)	7.833	22.2
^{157}Tb	71 y	$3/2^+ \rightarrow 3/2^-$	60.04(30)	K: 50.2391(5)	L _I : 8.3756(5)	7.124	9.76
^{163}Ho	4570 y	$7/2^- \rightarrow 5/2^-$	2.555(16)	M _I : 2.0468(5)	N _I : 0.4163(5)	4.151	0.51
^{179}Ta	1.82 y	$7/2^+ \rightarrow 9/2^+$	105.6(4)	K: 65.3508(6)	L _I : 11.2707(4)	6.711	40.2
^{193}Pt	50 y	$1/2^- \rightarrow 3/2^+$	56.63(30)	L _I : 13.4185(3)	M _I : 3.1737(17)	4.077	43.2
^{202}Pb	52 ky	$0^+ \rightarrow 2^-$	46(14)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	30.7
^{205}Pb	13 My	$5/2^- \rightarrow 1/2^+$	50.6(5)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	35.3
^{235}Np	396 d	$5/2^+ \rightarrow 7/2^-$	124.2(9)	K: 115.6061(16)	L _I : 21.7574(3)	5.587	8.6

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Sterile Neutrino and ^{163}Ho

- Amplitude of the line H for only active neutrinos

$$W_{Ha} = A(Q_{EC} - E_H)^2 B_H \varphi_H^2(0)$$

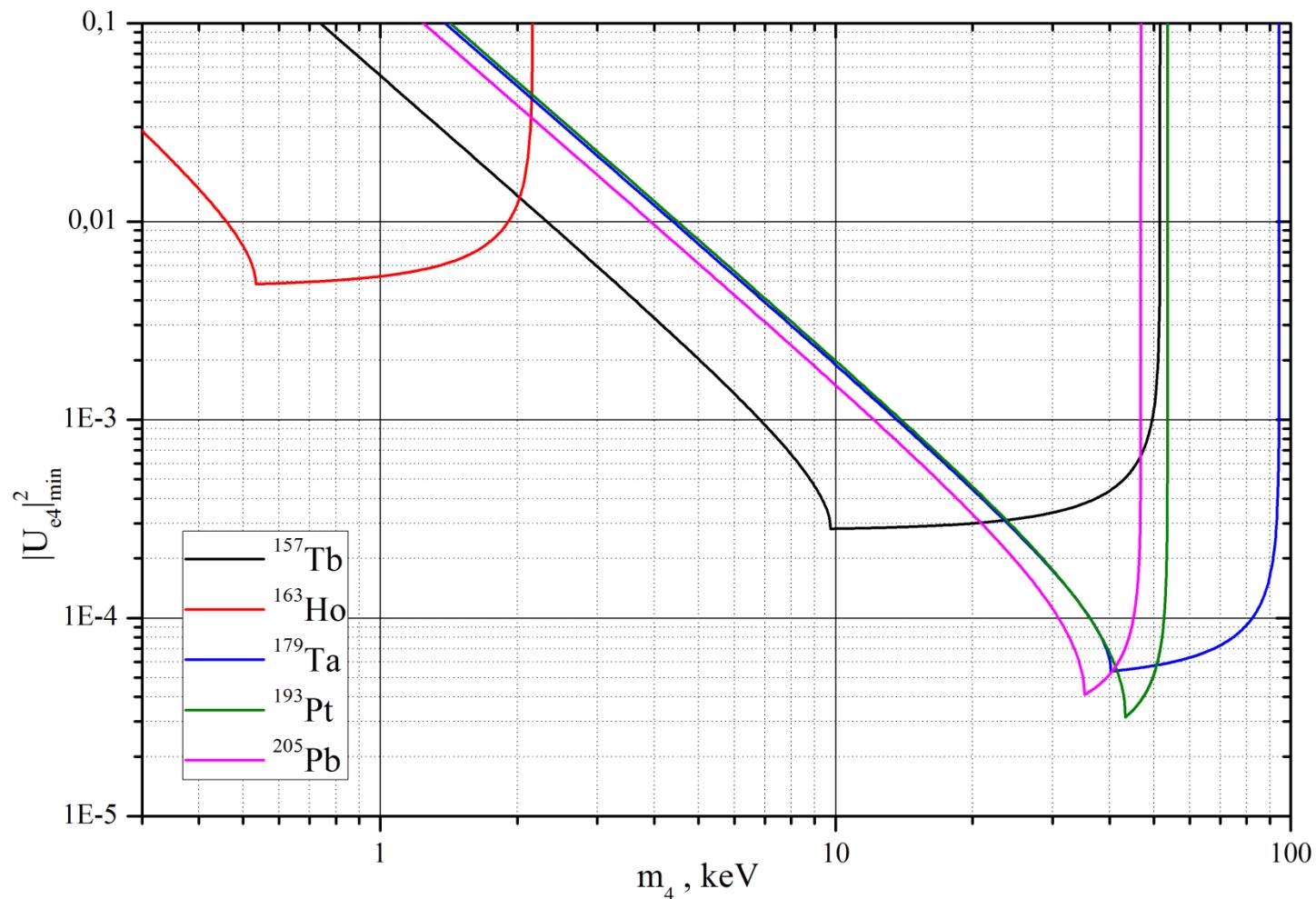
- Amplitude of the line H for 3+1 model in case of $m_a = 0 \text{ eV}$

$$W_{Hs} = A(Q_{EC} - E_H)^2 \left[\left(1 - |U_{e4}|^2 \right) + |U_{e4}|^2 \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_c)^2}} H(Q_{EC} - E_c - m_4) \right] B_H \varphi_H^2(0)$$

- Ratio between amplitudes of two lines in the spectrum for 3+1 model in case of $m_a = 0 \text{ eV}$

$$\left(\frac{W_{H1}}{W_{H2}} \right)_s = \left(\frac{W_{H1}}{W_{H2}} \right)_a \frac{|U_{e4}|^2 \left[H(Q_{EC} - E_1 - m_4) \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_1)^2}} - 1 \right] + 1}{|U_{e4}|^2 \left[H(Q_{EC} - E_2 - m_4) \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_2)^2}} - 1 \right] + 1}$$

Sterile Neutrino (keV) and Electron Capture



Same statistics + including errors : $(\delta \psi_{i,j} = 0)$ $\delta Q_{EC} = 1 \text{ eV}$ $\delta E_{i,j} = 0.1 \text{ eV}$.

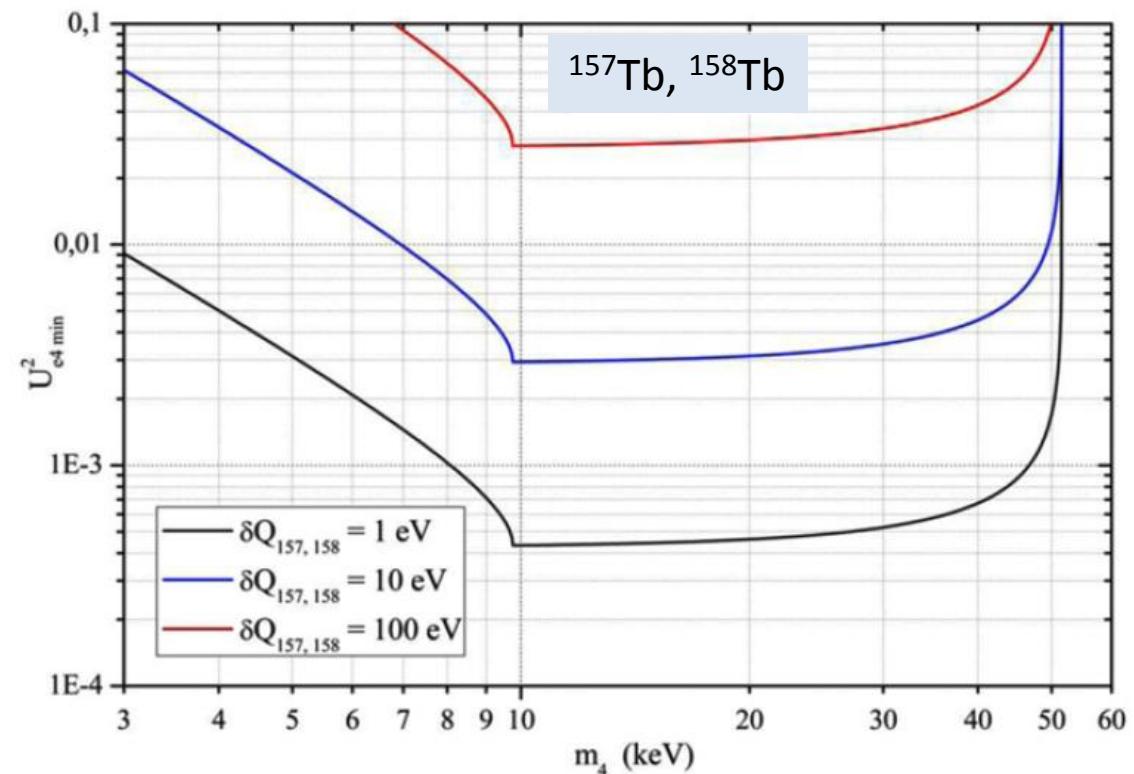
Sterile Neutrino (keV) and Electron Capture

Two EC isotopes of the same element:

$$\zeta_{\text{st}} \equiv \frac{\left(\lambda_i/\lambda_j\right)_1}{\left(\lambda_i/\lambda_j\right)_2} = \zeta_{\text{act}} \frac{\left[1 - U_{e4}^2 (1 - \omega_{i1})\right] \left[1 - U_{e4}^2 (1 - \omega_{j2})\right]}{\left[1 - U_{e4}^2 (1 - \omega_{i2})\right] \left[1 - U_{e4}^2 (1 - \omega_{j1})\right]}$$

$$\omega_{lk} \equiv H \left[(Q_k - B_l) - m_s \right] \cdot \sqrt{1 - \left(\frac{m_s}{Q_k - B_l} \right)^2}$$

$$\zeta_{\text{act}} = \left[\frac{(Q_1 - B_i)(Q_2 - B_j)}{(Q_2 - B_i)(Q_1 - B_j)} \right]^2$$



Conclusions

- The ECHo experiment can investigate the electron neutrino mass in sub-eV range
- The sensitivity of ^{163}Ho - based experiment to the sterile neutrino is limited to $m_s < Q_{\text{EC}}$.
- Other candidates can be found in the EC sector. They can cover a larger sterile neutrino mass range.

Thank you!

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