# Cosmology Confronts Fundamental Questions:

From Dark Energy and Inflation to Eternity..

or.. The Good, the Bad, and the Ugly

### A Fundamental Question

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# Are Fundamental Cosmological Questions Falsifiable?



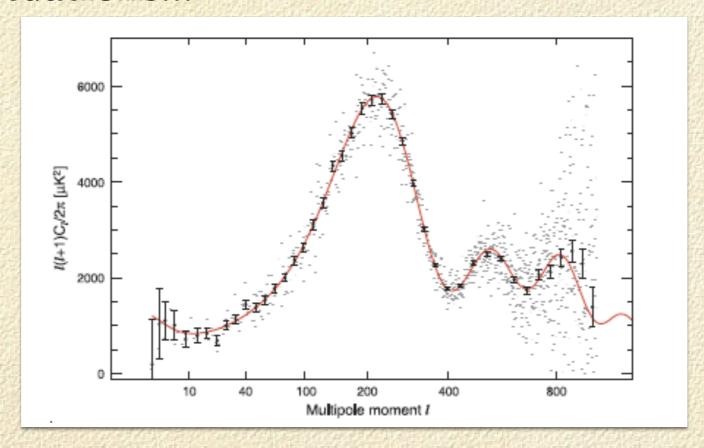
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❷ Flatness, Isotropy, Adiabatic, Gaussian Density Fluctuations...

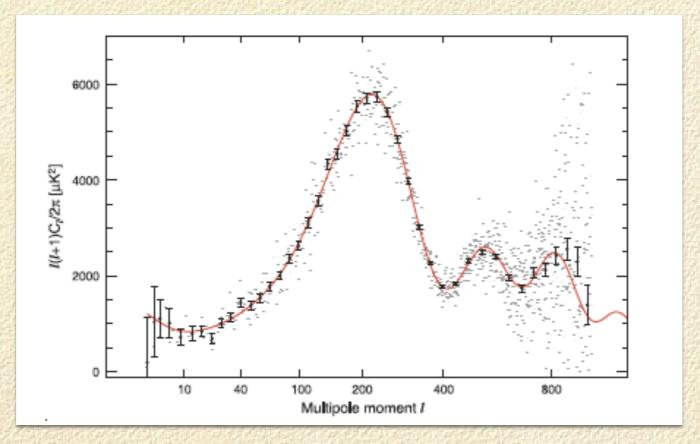
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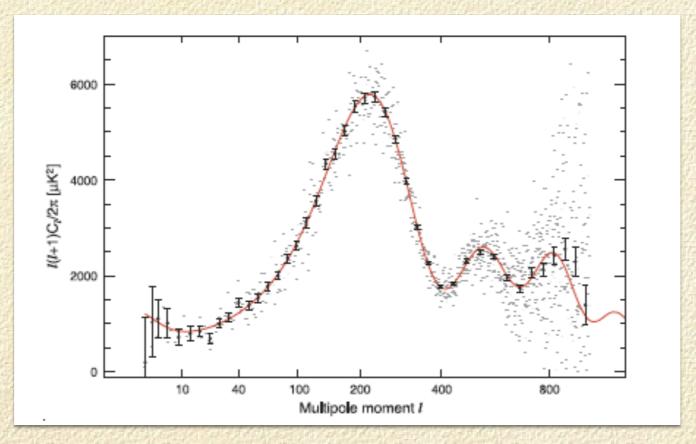
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Is Consistency Enough?

# I. The Good: "Proving" Inflation

Flatness, Isotropy, Adiabatic, Gaussian Density Fluctuations...



- Is Consistency Enough?
- Q Unique, Falsifiable Predictions?



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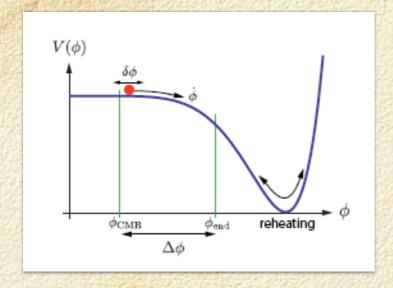
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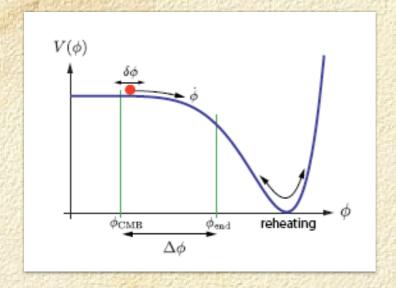
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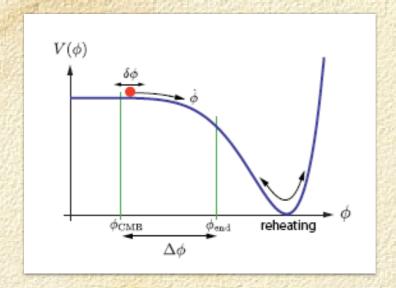


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#### A GENERIC PREDICTION OF INFLATION!

Note: Scalar density perturbations depend upon shape of potential. Tensor perturbations just depend upon scale!

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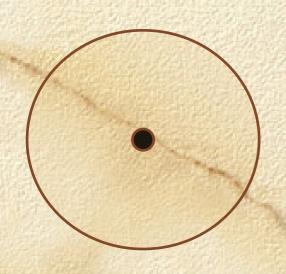
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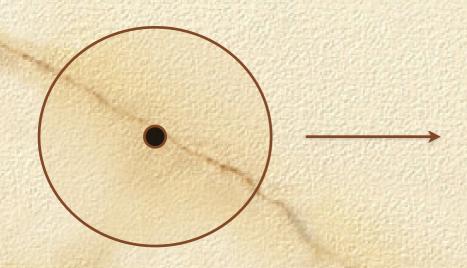
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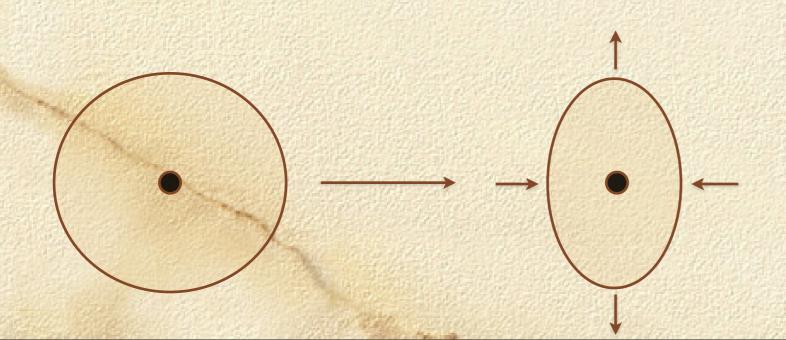
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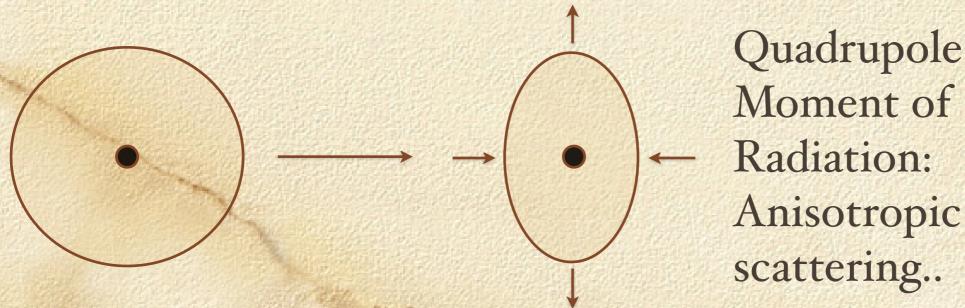
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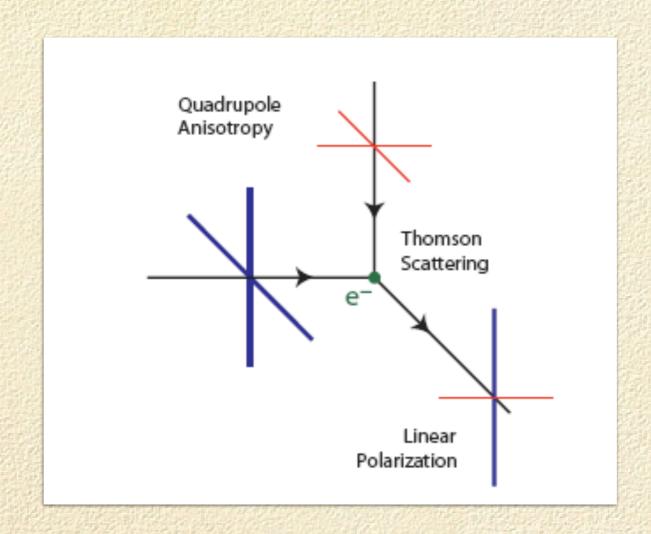
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Polarization of the CMB!:

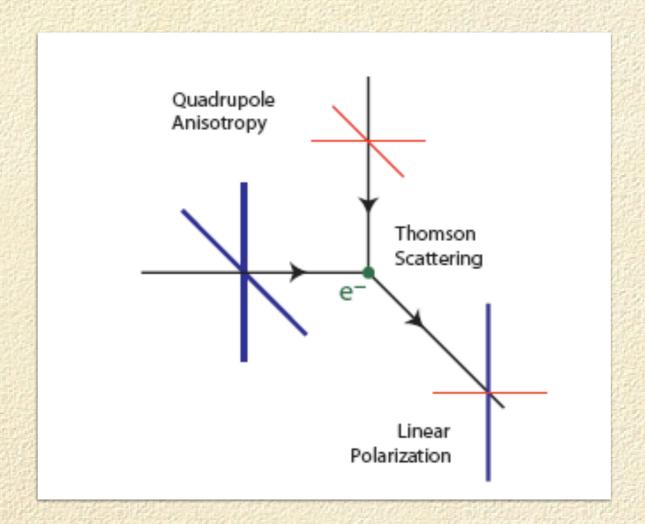


Moment of Anisotropic

### Polarization from Quadrupole



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cf:
Dodelson

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 i.e.  $r \approx 0$  for flat potential

NOTE: This actually makes it appear that r is insensitive to M, but this is not true... because S =∞

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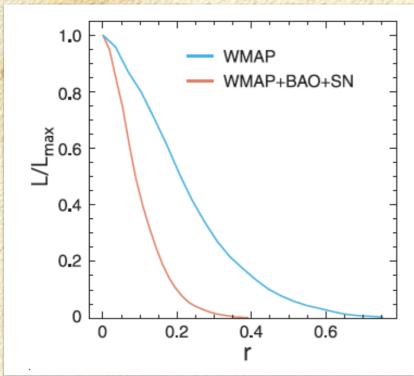
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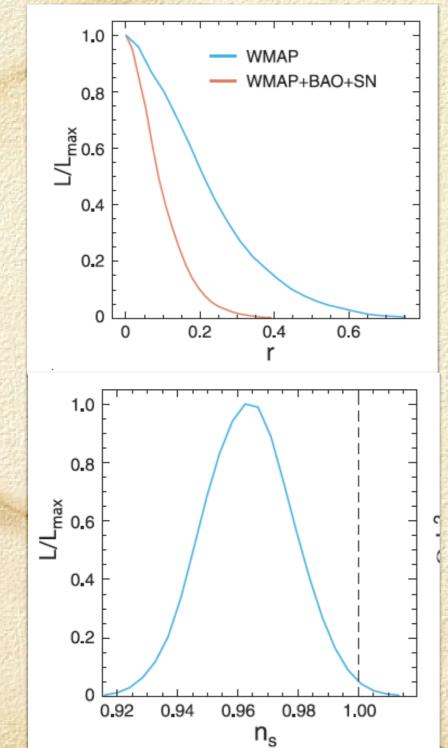
MEASURING POLARIZATION AND SPECTRUM CONSTRAIN r and PROBE INFLATION!

**CURRENT LIMITS:** 

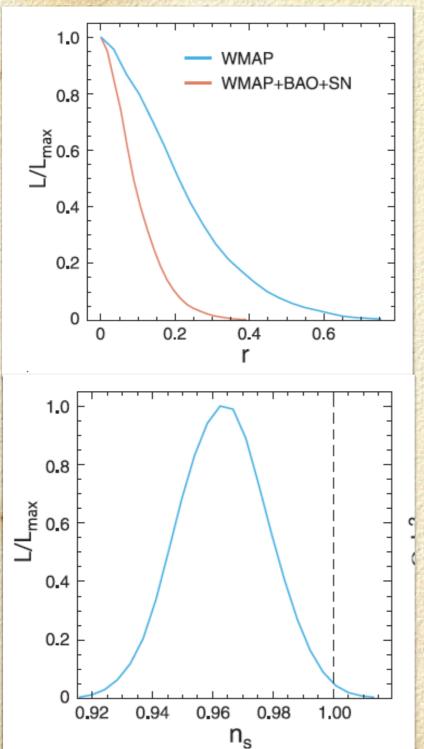
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(LB)

see

also

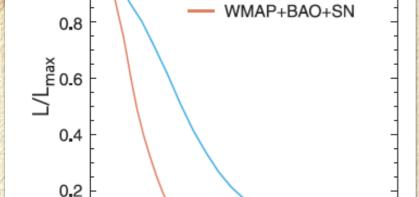
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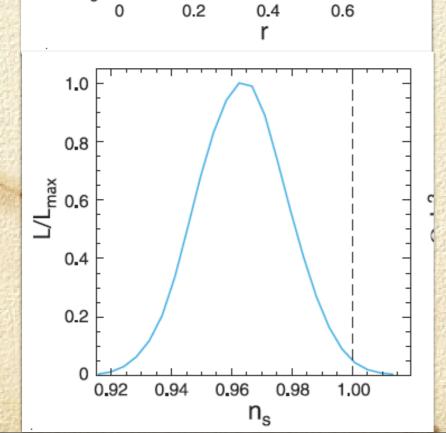
HdV

2008

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### 1.0 - WMAP

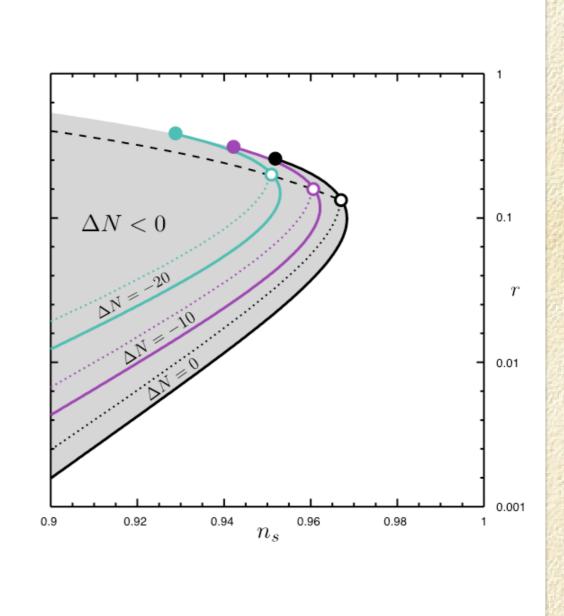




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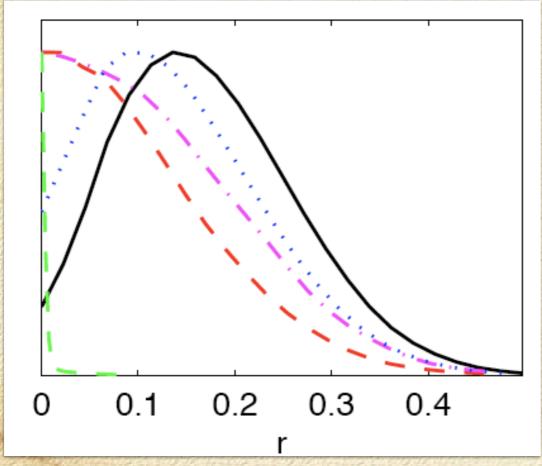
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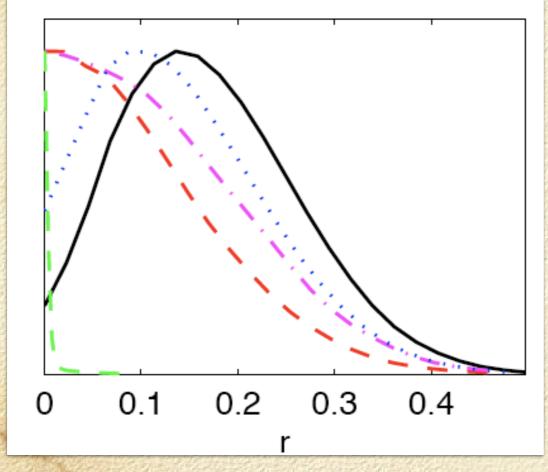
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(WV,LMK,JH, '08)

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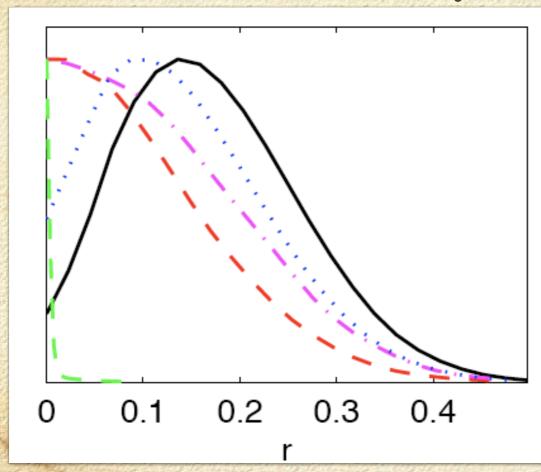


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Bayesian Complexity < 8

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LMK '92:

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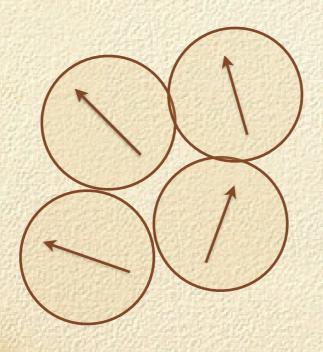
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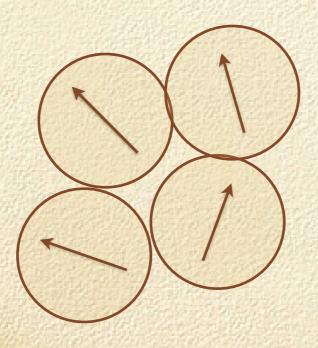
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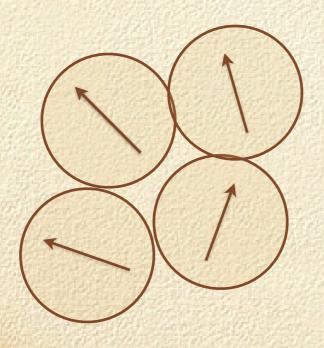


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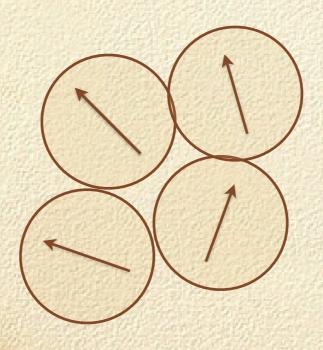
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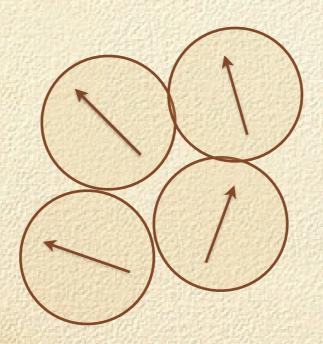
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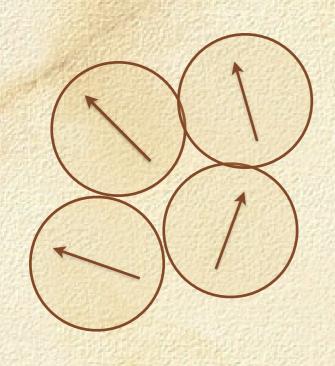


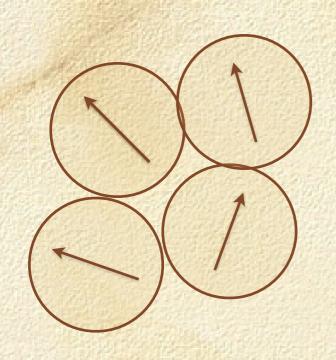
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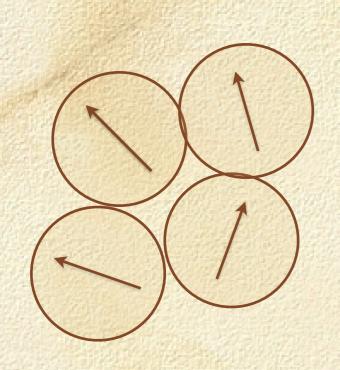
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When horizon grows, gradients relax. Use quadrupole approx to calculate generation of gravitational waves ON THE HORIZON.



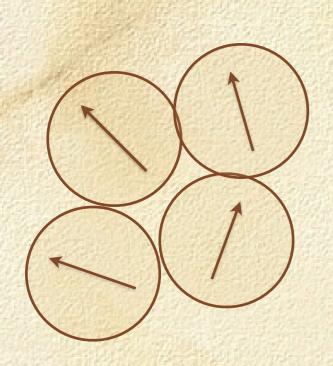


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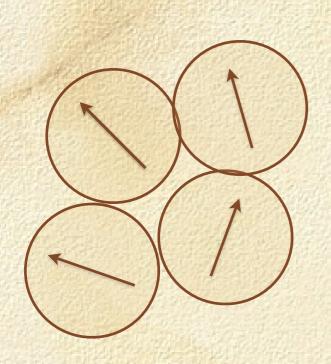


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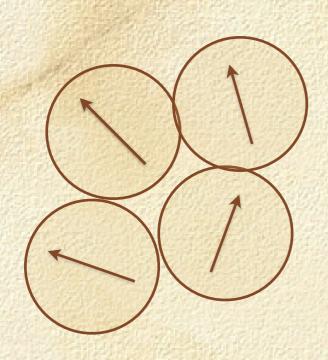


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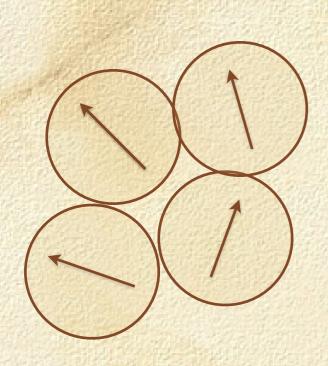
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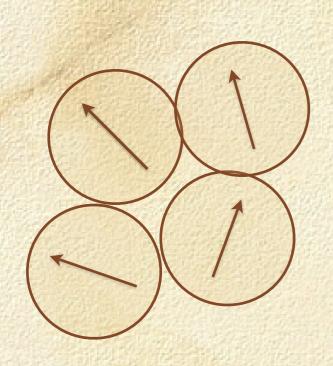
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$$h \approx H^{-1} (G^2 f^2 v^4 H^2)^{1/2} = G f v^2 \approx f \left(\frac{v}{M_{Pl}}\right)^2$$

Identical to Inflation.. Flat Spectrum..continued relaxation on H scale.. up to factor of f.. Q: What is f?

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KJ-S, LMK, HM, '08

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BUT!!!! for pt ≈ 1 C ≈ 2000! FOUR ORDERS of MAG!

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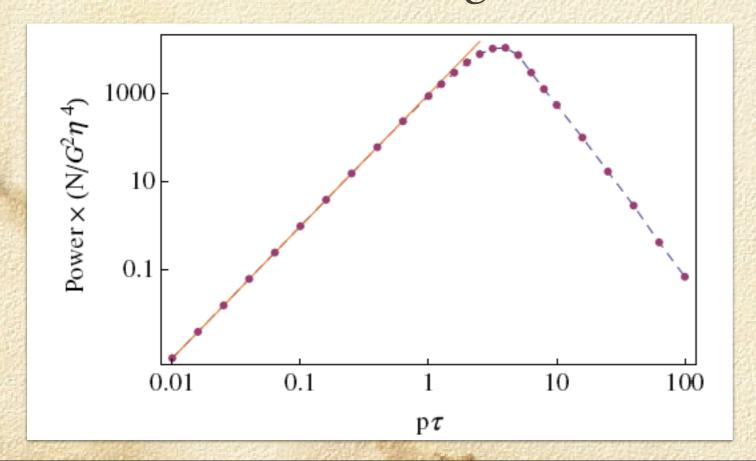
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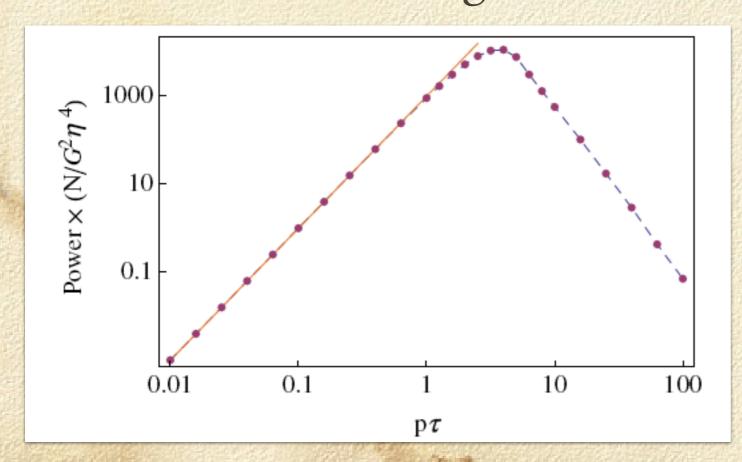




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Q: Spectrum is flat on horizon scale, but differs on superhorizon at any time. Does this produce difference in polarization???

**QINFLATION: NO UNAMBIGUOUS PREDICTION?** 

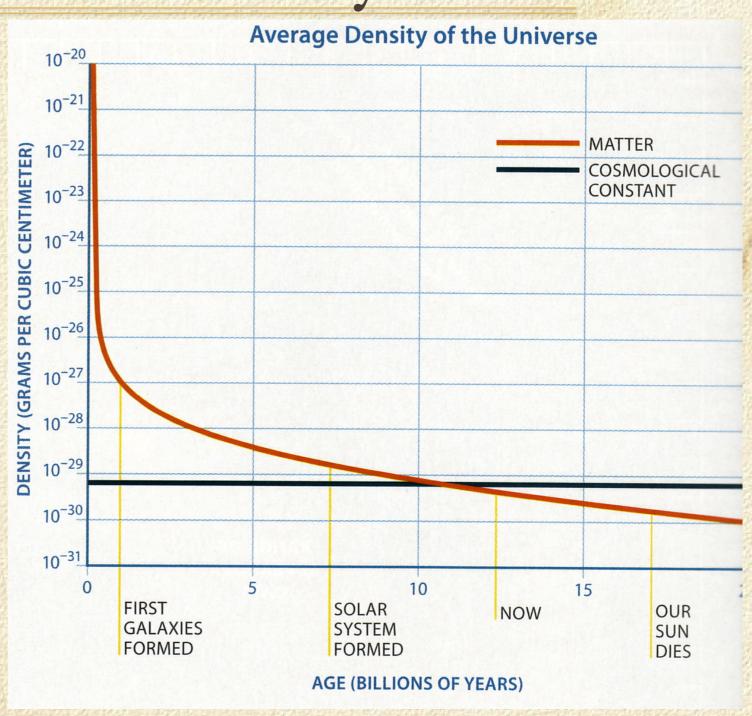
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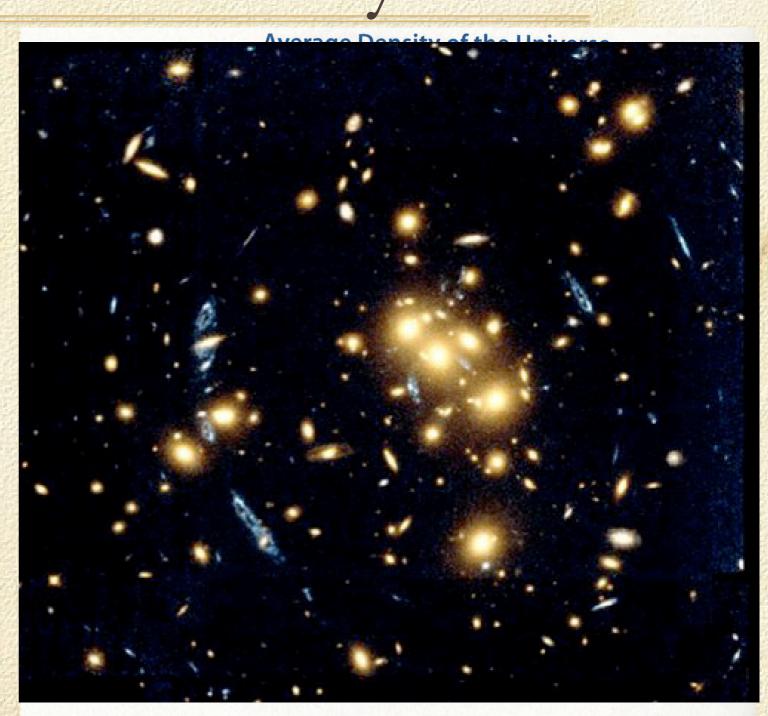
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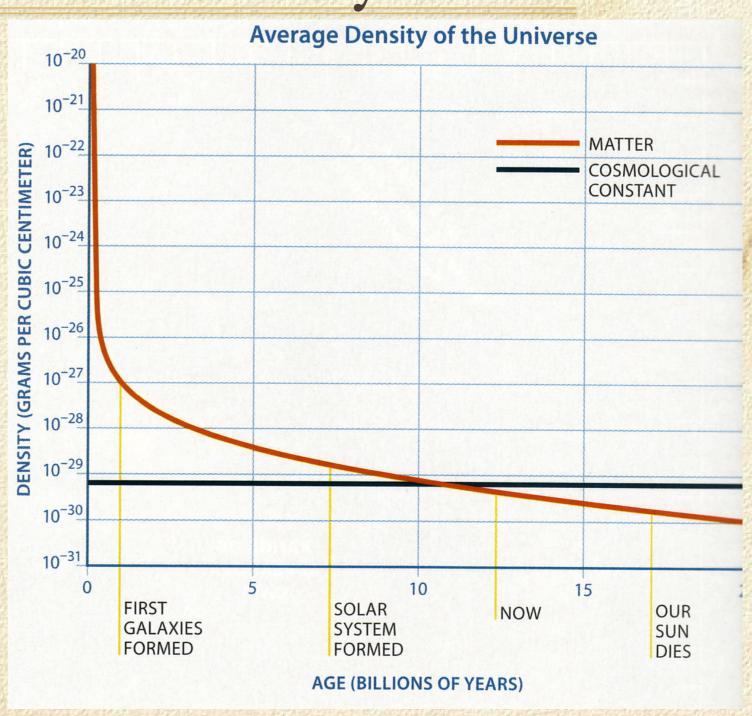
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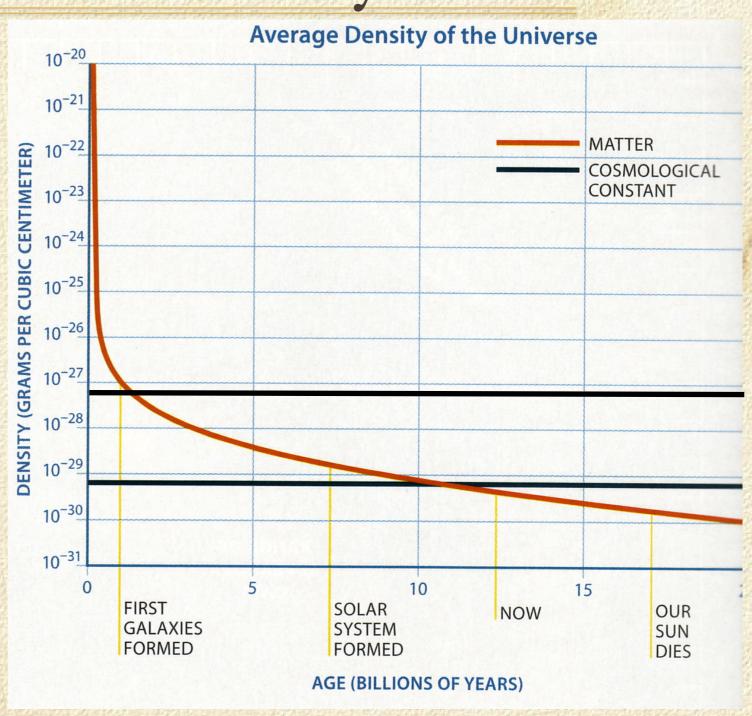
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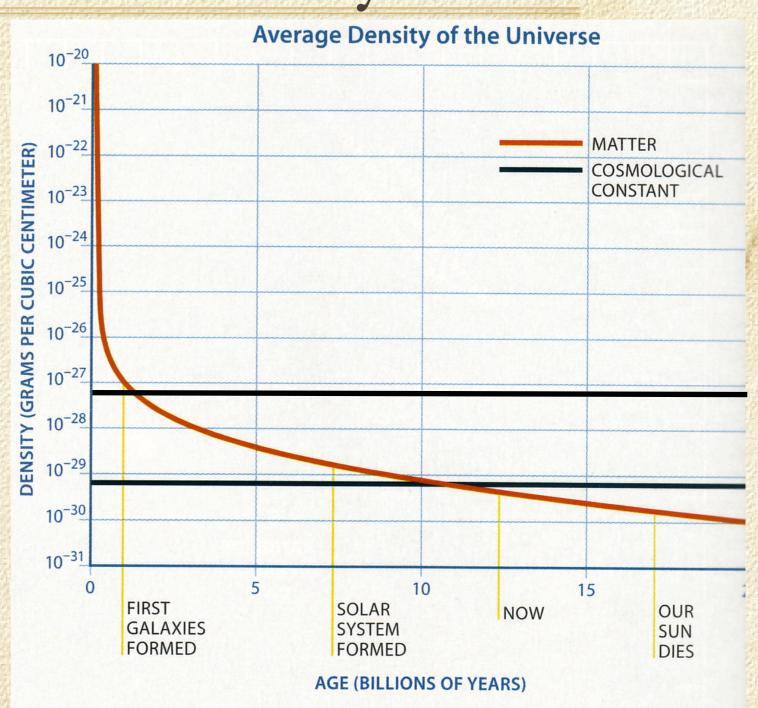
### The Bad: Dark Energy





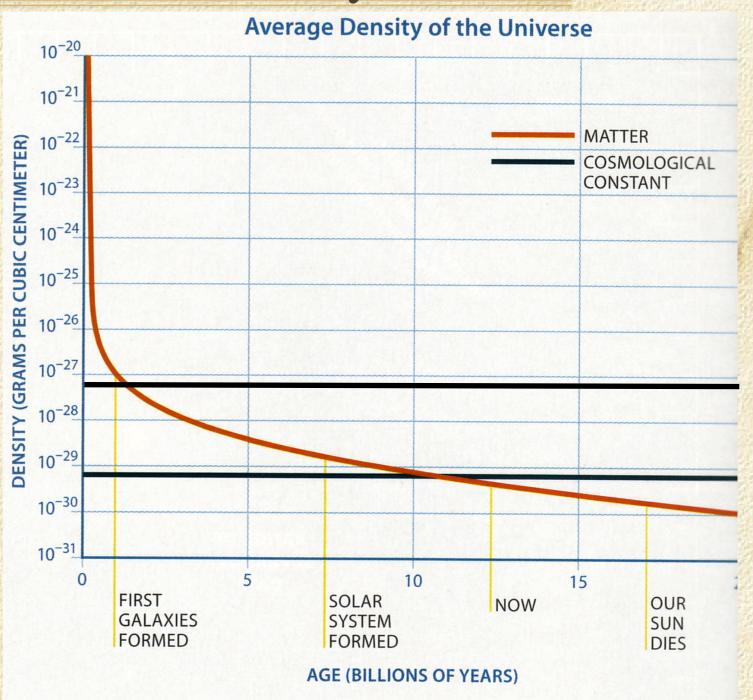






Maybe this is telling us something?

Galaxies
Never
Form?



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#### Anthropic Mania

IF there are many different universes, and the energy of empty space can vary in each one, then only those in which it is not much greater than what we measure will galaxies form... and only then will stars and planets form, and only then astronomers....

The Constants of Nature and the Puzzles of Modern Physics

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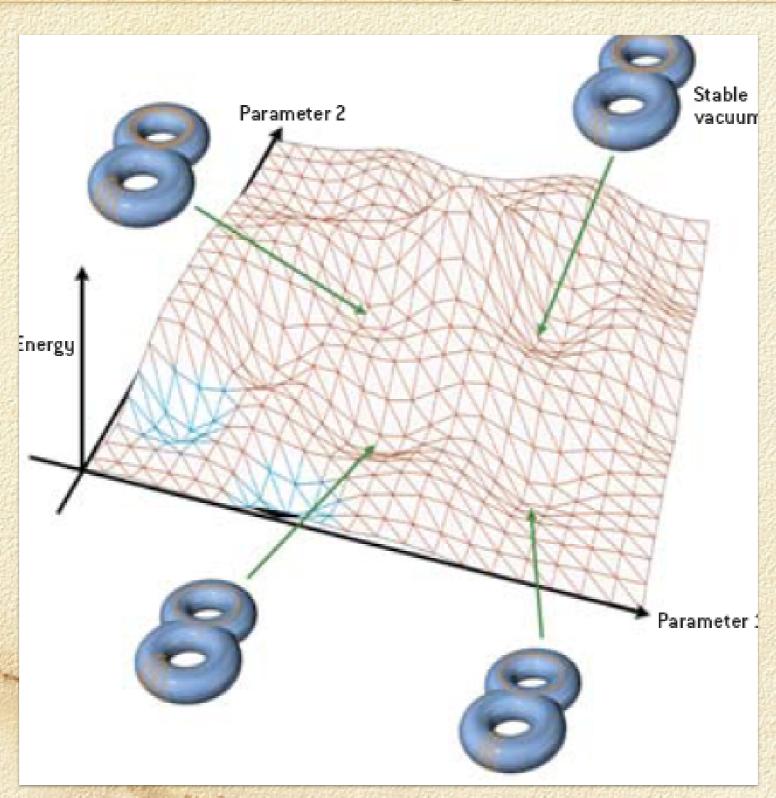
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- ...
- A theory of anything?

## The Landscape of Nothingness



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#### VASTLY DIFFERENT RESULTS DEPENDING UPON CHOICE!



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LMK, GDS 04

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Total Energy  $\approx 3 \times 10^{67}$  Joules  $\approx$  total baryonic rest mass within today's horizon.

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Moore's Law < 400 Years!

High Tech Civilizations don't last long?

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# Can Anthropic Principle "Explain" Lambda?

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As currently framed: just demonstrates correlations..

### Typicality is Everything!! IM, LMK, GDS 07

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What Really Matters is

IM, LMK, GDS 07

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IM, LMK, GDS 07

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IM, LMK, GDS 07

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IM, LMK, GDS 07

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IM, LMK, GDS 07

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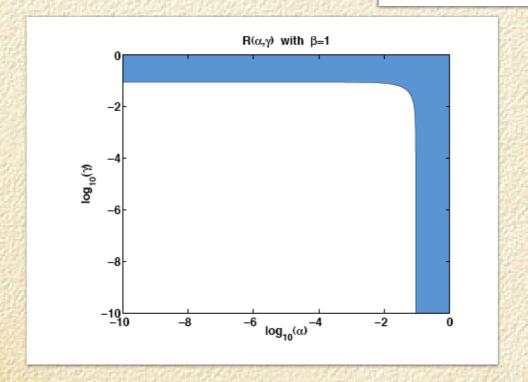
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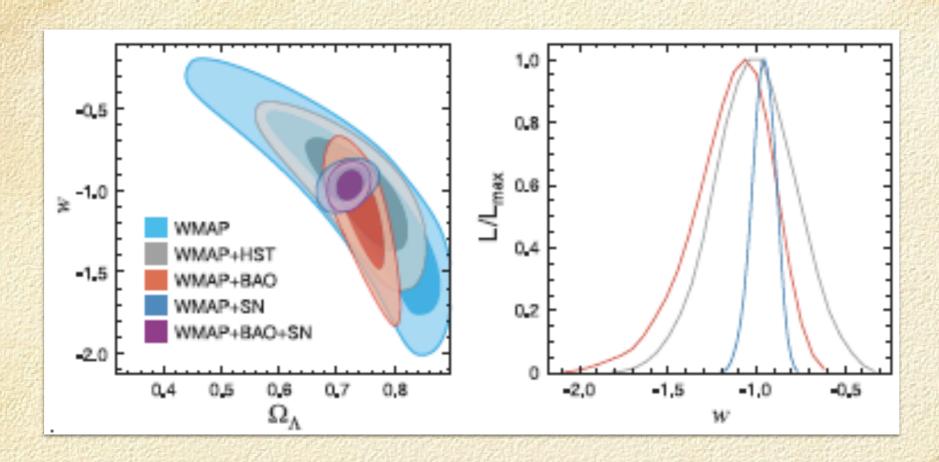
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- Observations suggest w ≈-1
- Measuring w ≈-1 therefore tells us nothing.

## The End of an Era?





## 1. Dark Energy:

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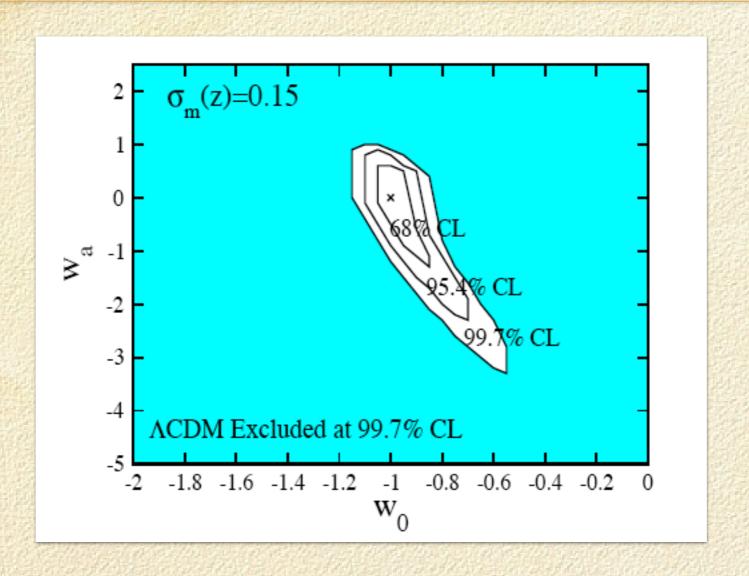
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- existing limits -1.2 < w < -0.8\* already rule out many alternative models. How much better can we do.. with existing theoretical uncertainties and expected observational accuracy?
- The PROBLEM: We DON'T HAVE ANY IDEA of w(z). Hence limits on w=constant are not appropriate.



□ 3000 SN ..LMK, D.H. K.J-S 07



## Pure theory required?

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Betting on the future!

Recall Cosmology CIRCA 1900: Static and Eternal

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What will the future bring?:

## Our Island Universe....





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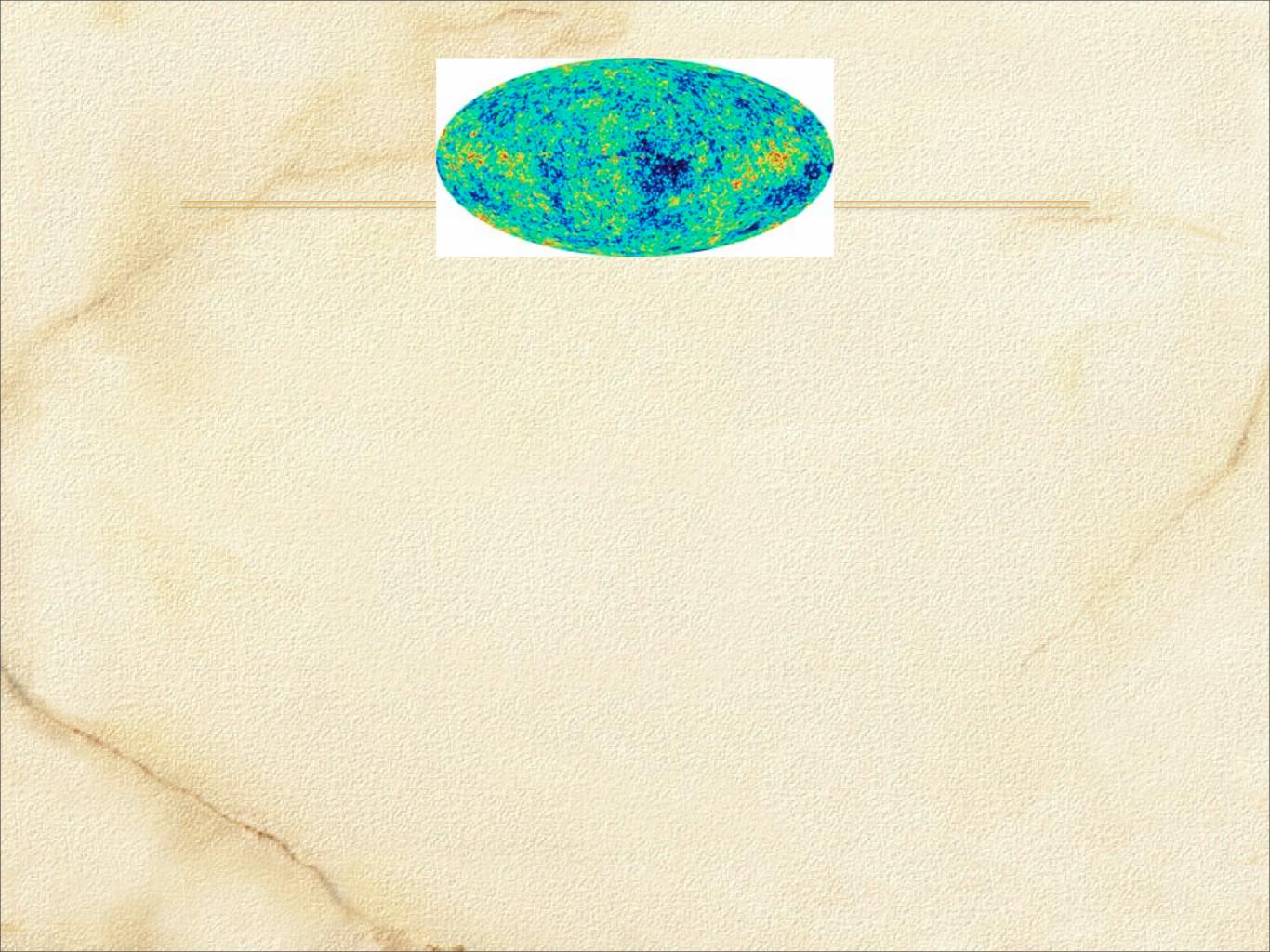
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 $\rho_{\Lambda} \ll \rho_{M}$ , but also when  $\rho_{\Lambda} \gg \rho_{M}$ .







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- achieved when universe less than 50 times its present age..



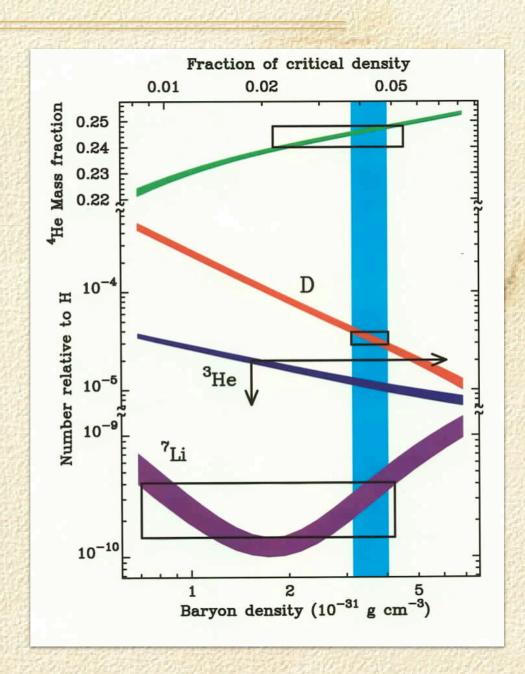
5. General Relativity no help! (Recall LeMaitre)

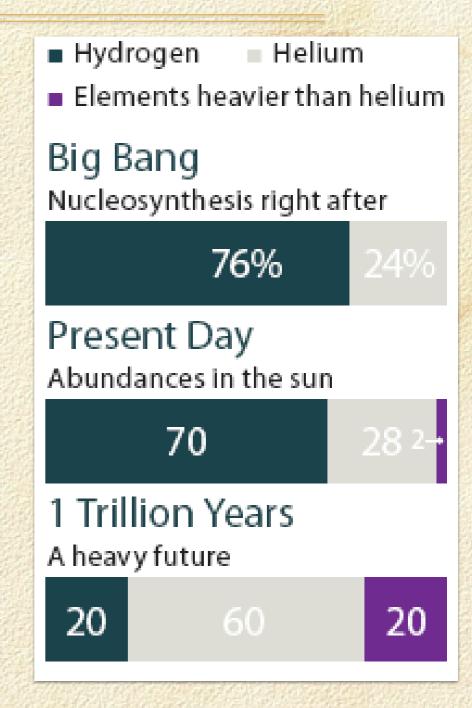
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- 7. What about finite past?

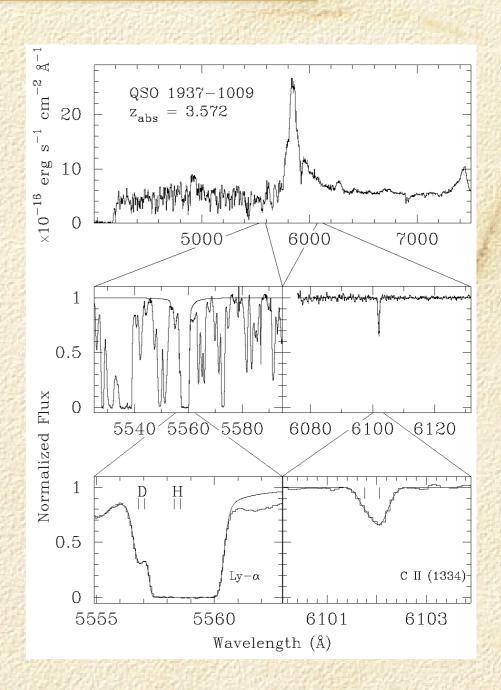






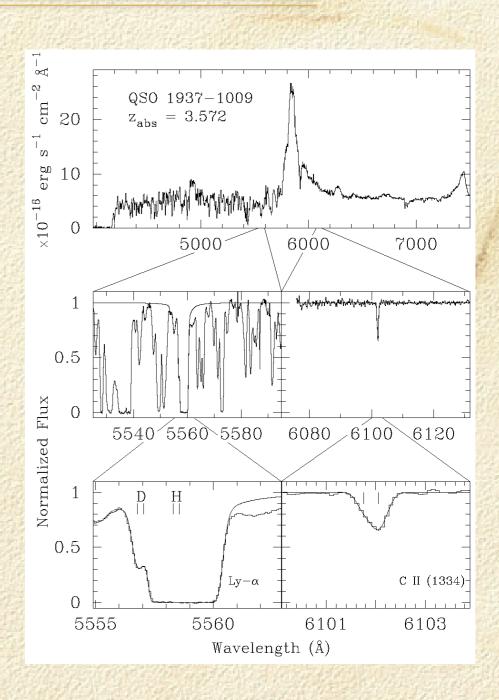
9. No primordial deuterium absorption from distant quasar

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No evidence of primordial big bang production!



### FALSIFIABLE SCIENCE: Return of Static Universe!



#### The Good News

We live in a very special time: the only time when we can observationally verify that we live at a very special time! More important, although we are certainly fortunate to live at a time when the observational pillars of the big bang are all detectable, we can easily envisage that other fundamental aspects of the universe are unobservable today. What have we already lost? Rather than being self-satisfied, we should feel humble. Perhaps someday we will find that our current careful and apparently complete understanding of the universe is seriously wanting.



#### THE GOOD:

Golden Era of Observational Cosmology has changed everything. New observations have a realistic chance of probing nature on heretofore unimaginable scales.

### THE BAD:

#### THE BAD:

# Fundamental Limits on science may be looming



#### THE UGLY?:

The nature of fundamental science may change

#### THE UGLY?:

The nature of fundamental science may change or end



## We will be lonely and ignorant, but dominant....